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Investigation of Stability and Reproducibility of

Perovskite Solar Cells

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Untersuchung der Stabilität und Reproduzierbarkeit von Perowskit-Solarzellen

Im vergangenen Jahrzehnt konnte die Effizienz von Photovoltaik Bauteilen auf Basis von Bleihalogenoid-Perowskit Materialien erheblich gesteigert werden. Die Stabilität und Reproduzierbarkeit von Perowskit-Solarzellen sind jedoch nach wie vor zwei der größten Herausforderungen und stehen ihrer Einführung als kostengünstige und effiziente Nachfolgegeneration für Photovoltaik Bauteile im Weg. Diese Doktorarbeit zielt darauf ab, diese beiden Aspekte in drei Teilen zu untersuchen. Zunächst veranschaulichen wir die Bedeutung der Mikrostruktur beim Abbau von Methylammonium-Blei-Triiodid-Perowskit-Filmen unter kontrollierter Atmosphäre mit variablem Sauerstoffgehalt und kontinuierlicher Beleuchtung. CH₃NH₃PbI₃-Filme mit kleinen, unregelmäßig geformten Kristallkörnern und hoher Defektdichte werden deutlich schneller und stärker abgebaut, als Filme mit großen, gleichmäßig geformten Kristallkörnern und besseren elektronischen Eigenschaften. Im zweiten Teil dieser Arbeit wird gezeigt, dass das Vorhandensein von großflächigen Inhomogenitäten in der chemischen Zusammensetzung sowie den elektronischen Eigenschaften von CH₃NH₃PbI₃-Filmen ein häufig auftretendes Problem ist. Des weiteren wird eine Möglichkeit vorgestellt, dieses Problem durch Verbesserung des Herstellungsprozesses zu mindern und eine Perowskitschicht mit einheitlicher Oberflächenzusammensetzung und einheitlichen elektronischen Eigenschaften zu erhalten. Der letzte Abschnitt zielt darauf ab, die Reproduzierbarkeit von Dreifachkationen-Perowskit-Bauteilen $(Cs_{0.05}(MA_{0.17}FA_{0.83})_{0.95}Pb(I_{0.9}Br_{0.1})_3)$ durch Optimierung des 'Antisolvent'-Verfahrens zu verbessern. Mit Hilfe von Rasterelektronenmikroskopie und photovoltaischen Messungen wird die Beziehung zwischen Antisolvent-Typ und Tropfgeschwindigkeit ermittelt und deren Effekte auf die Perowskitfilmbildung sowie die Leistungsfähigkeit der Bauteile untersucht.

Investigation of Stability and Reproducibility of Perovskite Solar Cells

Last decade has witnessed a remarkable increase in the efficiency of photovoltaic devices based on lead halide perovskite materials. However, the stability and reproducibility of perovskite solar cells remain two of the biggest challenges impeding their application as the next generation of low-cost and efficient photovoltaics. This thesis aims to investigating these two issues in three parts. First, we reveal the important role of microstructure in the degradation processes of methylammonium lead triiodide perovskite films in controlled oxygen atmospheres under continuous illumination. $CH_3NH_3PbI_3$ films with small, irregular grains and high defect density degrade much faster and more severely than films with large uniform grains and better electronic properties. The second part of this thesis demonstrates that the presence of large-scale chemical compositional and electronic inhomogeneities in $CH_3NH_3PbI_3$ films is a common issue, and suggests one way to mitigate this problem by improving the fabrication process to obtain a perovskite layer with a uniform surface composition and electronic properties. The last section aims to improve the reproducibility of triple cation perovskite $(C_{5_{0.05}}(MA_{0.17}FA_{0.83})_{0.95}Pb(I_{0.9}Br_{0.1})_3)$ devices by optimizing the antisolvent treatment procedure. Scanning electron microscopy and photovoltaic measurements elucidate the relationship between antisolvent type and dripping speed on perovskite film formation and device performance.

Contents

1	Intr	oductic)n	4			
2		Theoretical Background of Hybrid Organic-Inorganic Metal Halide Perovskite Solar Cells					
	2.1		iew				
	2.2	Introd	uction to Solar Cells	7			
		2.2.1	What is a Solar Cell?				
		2.2.2	The Solar Spectrum	8			
		2.2.3	Characteristic Performance Parameters				
		2.2.4	External Quantum Efficiency	12			
		2.2.5	Efficiency Limits	13			
	2.3	Perovs	skite Solar Cells	15			
		2.3.1	What is Perovskite?	15			
		2.3.2	Evolution of Perovskite Solar Cells	15			
		2.3.3	Material Properties and Compositional Advances	17			
		2.3.4	Fabrication Methods	20			
		2.3.5	Device Architecture and Charge Selective Contacts	23			
		2.3.6	Device Stability and Degradation	26			
~	-	•		29			
3	•						
	3.1	-	iew				
	3.2		and Device Fabrication				
		3.2.1	Substrates				
		3.2.2	Hole Transport Layers				
		3.2.3	Perovskite Layers				
		3.2.4	Electron Transport Layers				
		3.2.5	Materials				
	3.3		and Device Characterization				
		3.3.1	Profilometery				
		3.3.2	Atomic Force Microscopy (AFM)				
		3.3.3	Scanning Electron Microscopy (SEM)				
		3.3.4	X-Ray Diffraction (XRD)				
		3.3.5	UV-vis Spectroscopy (UV-vis)				
		3.3.6	Photothermal Deflection Spectroscopy (PDS)				
		3.3.7	Photoemission Spectroscopy	38			
		3.3.8	XPS and UPS Mapping	40			
		3.3.9	Photoluminescence Quantum Efficiency (PLQE)				
		3.3.10	Solar Simulator and J-V measurement Setup	42			

		3.3.11External Quantum Efficiency (EQE)3.3.12Environmental Rig				
4	Oxy	gen and Light Induced Photodegradation	45			
	4.1	Overview	. 45			
	4.2	Introduction	. 45			
	4.3	Degraded Films in Oxygen under Illumination	. 46			
		4.3.1 Optical Properties				
		4.3.2 Surface Compositional Properties				
	4.4	Photovoltaic Performance of Degraded films	. 53			
	4.5	Discussion—Role of Microstructure				
		4.5.1 Evolution of Microstructure in Degradation				
		4.5.2 Role of Grain Boundaries				
	4.6	Conclusions				
5	on	ects of Large-Scale Compositional and Electronic Inhomogeneities Device Performance	59			
	5.1	Overview				
	5.2	Introduction				
	5.3	XPS and UPS Mapping				
		5.3.1 I/Pb Maps				
		5.3.2 IP maps \ldots				
		5.3.3 Influence of Storage Conditions				
		5.3.4 Possible Causes for the Surface Inhomogeneities				
	5.4	Effect on Device Performance				
	5.5	Conclusions	. 68			
6	The Relationship between Antisolvent Type and Dripping Speed,					
		ermining the Device Performance	69			
	6.1	Overview				
	6.2	Introduction				
	6.3	The Relationship between Antisolvent type and Dripping Speed $\ . \ .$				
		6.3.1 Device Performance				
		6.3.2 Surface Chemical Composition	. 72			
		6.3.3 Microstructure				
	6.4	Discussion				
	6.5	Conclusions	. 81			
7	7 Conclusions and Future Work					
A	open	dices	84			
Appendix A Figures						
A	Appendix B Tables					

Appendix C Calculate the Pbl_2 Amount from SEM Images	112
Appendix D Abbreviations	115
References	117
List of my Publications	135
Acknowledgements	136

1 Introduction

As an important source of renewable energy, solar energy has been harnessed in various ways, one of which is photovoltaics that convert the energy of sunlight into electricity by the photovoltaic effect. The great potential of photovoltaic devices, or solar cells, was foreseen by Werner von Siemens in 1885, commenting on the discovery of the photovoltaic effect: 'However great the scientific importance of this discovery may be, its practical value will be no less obvious when we reflect that the supply of solar energy is both without limit and without cost, and that it will continue to pour down upon us for countless ages after all the coal deposits of the earth have been exhausted and forgotten.' [1]

Crystalline silicon solar cells, as the first generation photovoltaic devices, still dominate more than 90% of the commercial market, despite the long payback period. The second generation is the thin-film solar cell, produced by depositing semiconducting materials (such as gallium arsenide (GaAs), cadmium telluride (CdTe), copper indium gallium diselenide (CIGS) and silicon) on a substrate. The world record of the highest efficient single junction solar cell is held by GaAs thin film solar cells at 29.1% [2]. The third generation of solar cells (i.e., organic solar cells, dye-sensitized solar cells(DSSCs), perovskite solar cells and quantum dot solar cells), often also described as 'emerging photovoltaics', are promising to achieve high efficiency at lower production cost. These materials are still under development, with full commercialization promising in the near future.

Perovskite solar cells originated as dye-sensitized solar cells (DSSCs), and only held an efficiency of 3.8% when discovered in 2009 [3]. Its efficiency was increased to 9.7% in 2012 by replacing the liquid with a solid hole transporting material [4]. Almost at the same time, perovskite was shown to be able to transport both electrons and holes, and reached an efficiency of 10.9% [5]. Since then, the following years witnessed a rapid growth of perovskite photovoltaic studies. The breakthrough of an efficiency exceeding 20% was achieved by Seok's group in 2015 [6] employing intramolecular exchange, which was later applied to the fabrication of advantageous mixed cation perovskites [7]. In 2016, Saliba et al. reported a 'triple cation' perovskite, yielding an efficiency of 21.1% and improved stability [8]. As of August 2019, the record efficiency of 25.2% was reported by the Korean Research Institute of Chemical Technology (KRICT) and Massachusetts Institute of Technology (MIT), certified by the National Renewable Energy Laboratory (NREL) [2].

Despite a decade of research, many problems still remain unsolved. One of the biggest challenge is the poor stability of the perovskite solar cell, which has been reported to degrade under various operating conditions such as moisture, oxygen, light, heat and electrical stress. In addition, large variations in the photovoltaic

performance have been reported by different research groups, but the origin is merely discussed. Furthermore, some high-functioning perovskite solar cells are fabricated using an 'antisolvent' treatment, but the requirements of a 'good' antisolvent that results in a high-quality perovskite film remain unclear, although a wide range of different solvents have been reported to produce good perovskite solar cells. In this thesis, we will investigate these issues.

In Chapter 2, a brief background of perovskite materials and their application in solar cells is presented. First, an introduction to solar cells is given, followed by the characteristic parameters of solar cells and their efficiency limits. In the second section, a brief introduction to perovskite materials and its development in the application of solar cells are briefly reviewed, followed by summaries of the material properties, fabrication methods, selective contacts and the stability issues.

Chapter 3 presents the experimental methods used in this work. The film and device fabrication methods and procedures are described, followed by an explanation of the techniques used for characterizing perovskite films and devices. These techniques are Profilometery, Atomic Force Microscopy (AFM), Scanning Electron Microscopy (SEM), X-ray Diffraction (XRD), UV-vis Spectroscopy (UV-vis), Photothermal Deflection Spectroscopy (PDS), X-Ray Photoemission Spectroscopy (XPS), Ultra-Violet Photoemission Spectroscopy (UPS), XPS and UPS Mapping, Photoluminescence Quantum Efficiency (PLQE) measurements, Solar Cell measurements (current density-voltage (J-V) and External Quantum Efficiency (EQE)) and Environmental Rig set-up.

Chapter 4 focuses on oxygen-induced degradation of $CH_3NH_3PbI_3$ perovskite films with different microstructures under continuous illumination. We find that the microstructure of perovskite films has a tremendous effect on the oxygen and light induced degradation. Perovskite films with small, irregular grains and high defect density degrade much faster and more severely than films with large uniform grains and better electronic properties.

Chapter 5 presents the large-scale chemical compositional and electronic inhomogeneities of $CH_3NH_3PbI_3$ perovskite films. We demonstrate that these inhomogeneities are not limited to certain fabrication methods, and are commonly observed by other research groups. We find that variations of photovoltaic performance parameters across the entire sample are related to the extent of the observed inhomogeneities, which explains the wide spread of device performance reported in the literature.

Chapter 6 demonstrates the relationship between antisolvent type and dripping speed, and their effects on photovoltaic performance of triple cation perovskite $(Cs_{0.05}(MA_{0.17}FA_{0.83})_{0.95}Pb(I_{0.9}Br_{0.1})_3)$ devices. We find that the solubility of organic precursors in the antisolvent has a tremendous effect on the perovskite formation, microstructure, and consequently, the final device performance. Additionally, we present that using certain antisolvents which possess a low solubility of organic precursors is able to improve device reproducibility by widening the dripping speed 'window'.

Finally, Chapter 7 presents the conclusions of this work and discusses briefly the

future directions of research.

2 Theoretical Background of Hybrid Organic-Inorganic Metal Halide Perovskite Solar Cells

2.1 Overview

The first part of this Chapter briefly describes the working principles of a solar cell, followed by an explanation of the characteristic parameters of solar cells. Lastly, the theoretical efficiency limits of a single junction solar cell are described.

In the second part, an introduction to the perovskite material and its development in photovoltaic applications are briefly reviewed. Next, perovskite material properties and advances in the perovskite composition are presented. Following this, fabrication methods of the perovskite film are introduced, together with a short review of device architectures and the selective contacts used in this work. Finally, the device stability is reviewed and two major degradation processes are briefly discussed.

2.2 Introduction to Solar Cells

2.2.1 What is a Solar Cell?

A solar cell is an electrical device that is capable of converting solar energy into electricity by the photovoltaic effect. Firstly, photons of the sunlight are absorbed by the active semiconducting material, generating bound electron-hole pairs (excitons). Next, the electron-hole pairs dissociate to form two separate unbound charge carriers. Lastly, the free electrons and holes are transported and extracted at respective contacts by an external load.

The inorganic semiconductor solar cell is configured as a p-n junction. When an ntype material (i.e., an intrinsic semiconductor doped with an electron donor element) is brought into contact with a p-type material (i.e., an intrinsic semiconductor doped with an electron acceptor element), free electrons diffuse from the n-type material across the junction and recombine with holes on the p-type side. Likewise, holes are also diffused and recombined with electrons on the n-type side. Due to the diffusion of the majority charge carriers (electrons for the n-type material and holes for the ptype semiconductor), a depletion region is formed near the junction interface, with an electric field that counteracts the diffusion (Figure 2.1). If a positive voltage (forward bias) is applied between the two ends of the p-n junction, electrons in

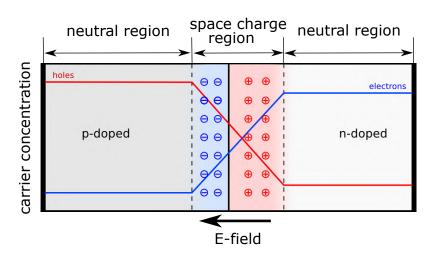


Figure 2.1: Schematic of a p-n junction in equilibrium without an external circuit. Adapted with permission from [9]. Copyright TheNoise, licensed under Creative Commons Attribution-Share Alike 3.0 Unported license.

the n-type region (and holes in the p-type region) are pushed towards the junction, neutralizing the depletion region and further generating a current flow.

Organic semiconductor solar cells usually are constructed in a bulk heterojunction consisting of mixed donor (p-type) and acceptor (n-type) organic materials, in order to facilitate charge separation and compensate the short exciton diffusion length of many organic material systems to improve their device performance.

The most commonly used material for perovskite solar cells is methylammonium lead halide, which has a tunable band gap controlled by the halide content [10, 11], and exhibits an electron-hole diffusion length over one micrometer [12]. The exciton binding energy in this material system has been reported to be low enough [13, 14] that free electrons and holes are present at room temperature. These excellent properties enable the perovskite material to function efficiently as an intrinsic active layer in a thin-film device.

2.2.2 The Solar Spectrum

The total solar irradiation received at the earth's surface depends on the length of light travelling through the atmosphere, which is quantified by air mass (denoted by AM). Air mass is defined as the direct optical length through the earth's atmosphere at a given location, and has value 1 for the normal incidence. To compensate for the variation over seasons, the solar spectrum is standardized by the American Society for Testing and Materials (ASTM). As illustrated in Figure 2.2a, AM 0 spectrum describes the extraterrestrial solar spectral irradiance distribution outside the earth's atmosphere and is given by the ASTM E-490. AM 1.5G spectrum describes the total terrestrial spectral irradiance distribution arriving at the surface of earth after passing through 1.5 times a standard air mass at zenith angle of 48.2°, while AM 1.5D represents only the direct contribution to the total global spectrum. Both

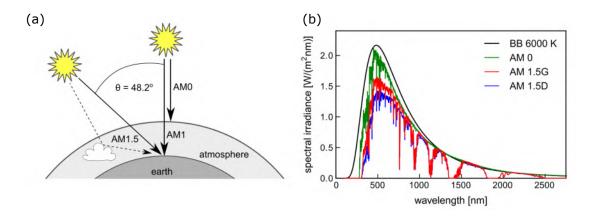


Figure 2.2: (a) Schematic representation of the solar spectral irradiance outside the earth's atmosphere (AM 0), and two terrestrial solar spectral irradiance on the earth's surface (AM 1.5D (the direct normal spectrum, solid arrow) and AM 1.5G (the total global spectrum, solid and dashed arrows)).
(b) Spectral irradiance of AM 0, AM 1.5D and AM 1.5G, in comparison to the black body spectrum with a surface temperature of 6000 K used by Shockley and Queisser (BB 6000K). Reproduced with permission form [15]. Copy right 2016 Elsevier Ltd.

spectra are given by the ASTM G-173-03. Figure 2.2b demonstrates the spectra of AM 0, AM 1.5G, AM 1.5D and a black body emission with a surface temperature of 6000 K. The Standard Test Conditions (STC) specifies a standard testing condition of solar cells: under AM 1.5G illumination (with an irradiance of 100 mW/cm^2) at 25 °C.

2.2.3 Characteristic Performance Parameters

The current of a solar cell consists of three components, a dark generation current I_0 , a recombination current I_{rec} , and a photocurrent I_{ph} . Divided by the illuminated active area on a solar cell, the current is converted to the current density (J = I/A) with A representing the illuminated area). Therefore, the total current density of a solar cell is:

$$J = J_{ph} + J_0 - J_{rec} (2.1)$$

The difference between the dark generation current density and the recombination current density equals the current density injected to a solar cell in the dark, which can be described with the Shockley diode equation [16]:

$$J_{dark} = J_0 - J_{rec} = J_0 \cdot (exp(\frac{qV}{nK_BT}) - 1)$$
(2.2)

with the parameter n being the ideality factor, describing the specific exponential voltage dependence of the current.

Figure 2.3 shows a typical current density-voltage (J-V) characteristic of a solar cell (for V > 0) with the following expression derived from equation 2.1 and 2.2:

$$J = J_{ph} - J_0 \cdot (exp(\frac{qV}{nK_BT}) - 1)$$
(2.3)

Open-circuit and short-circuit are two important operating conditions of a solar cell. The open-circuit voltage (V_{OC}) and the short-circuit current density (J_{SC}) under illumination are described by the expressions derived from equation 2.3:

$$V_{OC} = V(J = 0) = \frac{nK_BT}{q} \cdot ln(\frac{J_{ph}}{J_0} + 1)$$
(2.4)

$$J_{SC} = J(V = 0) = J_{ph}$$
(2.5)

The electrical power density output of a solar cell is the product of current density and voltage. The maximum power density is achieved at the maximum power point (MPP), with current density J_{MPP} and the voltage V_{MPP} .

The third important parameter is the fill factor (FF), defined as the ratio of the maximum power density at the MPP (dark green rectangle in Figure 2.3) and the theoretical maximum power density (light green rectangle in Figure 2.3), which is the product of V_{OC} and J_{SC} .

$$FF = \frac{J_{MPP} \cdot V_{MPP}}{J_{SC} \cdot V_{OC}}$$
(2.6)

The fourth characteristic parameter is the power conversion efficiency (PCE), which describes the maximum efficiency of a solar cell converting the incoming photon energy to the electrical power at the MPP.

$$PCE = \frac{P_{MPP}}{P_{in}} = \frac{FF \cdot J_{SC} \cdot V_{OC}}{I_E^{AM1.5G}}$$
(2.7)

with P_{in} being the incoming photon power density, $I_E^{AM1.5G} = 100 \,\mathrm{mW/cm^2}$ the integrated power density of AM 1.5G spectrum.

The equivalent circuit of a real solar cell is shown in Figure 2.4a, which consists of a photocurrent source I_{ph} , a series resistance R_s , a shunt resistance R_{sh} and a diode. The influence of these two resistances on the J-V curve is simulated in Figure 2.4b and c. The series resistance represents the contact resistance between electrodes and adjacent layers, mainly causing the loss in the fill factor. When $R_{sh} \rightarrow \infty$, FF drops for a larger R_s , whereas V_{OC} is unaffected. The shunt resistance represents the possible alternative pathways, such as pinholes in the active layer, and is responsible for the low fill factor and V_{OC} in the solar cell. When $R_s = 0$, V_{OC} and FF decrease with R_{sh} , while J_{SC} remains intact.

Derived from the equation 2.3, the J-V characteristics of a real solar cell can be described as:

$$J = J_{ph} - J_0 \cdot (exp(\frac{q(V+JR_s)}{nK_BT}) - 1) + \frac{V+JR_s}{R_{sh}}$$
(2.8)

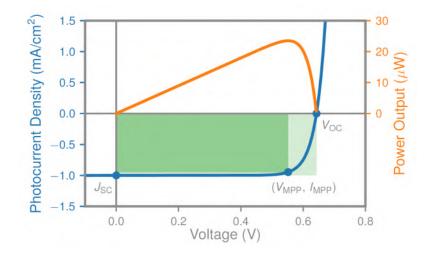


Figure 2.3: A typical J-V curve (blue line) and electrical power output (orange line) of a solar cell. Open-circuit voltage and short-circuit current conditions are denoted. The power is calculated as the product of current and voltage with an illuminated area ($A = 4.5 \text{ mm}^2$), whose maximum occurs at the maximum power point (MPP). Adapted from [17] with permission.

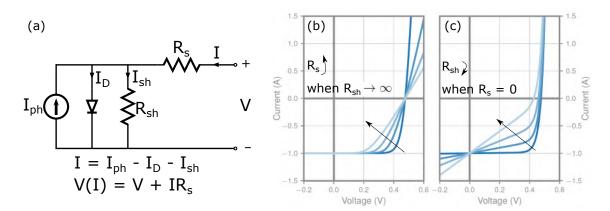


Figure 2.4: (a) An equivalent circuit of a solar cell with a photocurrent source I_{ph} , a diode, a series resistance R_s and a shunt resistance R_{sh} . (b-c) Modelling of the effect of two resistances on the J-V curve. The J-V curve changes from dark to light blue, when (b) R_s increases with $R_{sh} \rightarrow \infty$, and (c) R_{sh} decreases with $R_s = 0$. Adapted from [17] with permission.

2.2.4 External Quantum Efficiency

Another important parameter describing the efficiency of a photovoltaic device is the external quantum efficiency (EQE), which is defined as the ratio of the number of extracted charge carriers collected at the electrodes of a solar cell to the number of photons incident on the device:

$$EQE = \frac{\# \text{ of extracted charge carriers}}{\# \text{ of incident photons}}$$
(2.9)

It is obtained by measuring the short-circuit current density under illumination by monochromatic light as a function of wavelength λ :

$$EQE(\lambda) = \frac{J_{SC}(\lambda)/q}{\Phi(\lambda)} = \frac{J_{SC}(\lambda)/q}{P_{in}/E_{ph}(\lambda)}$$
(2.10)

where $\Phi(\lambda)$ is the photon flux (numbers of photons per second per m²) incident on the device at a wavelength λ , $E_{ph}(\lambda) = \frac{hc}{\lambda} = q \cdot \frac{1240}{\lambda(nm)}$ is the energy of a single photon at a given wavelength λ (with c and h being the velocity of light and plank constant, respectively). Hence, EQE can be written as:

$$EQE(\lambda) = \frac{hc}{q\lambda} \cdot \frac{J_{SC}(\lambda)}{P_{in}} = \frac{1240}{\lambda(nm)} \cdot \frac{J_{SC}(\lambda)}{P_{in}}$$
(2.11)

The spectral response (denoted by $SR(\lambda)$) is defined as the ratio of photocurrent generated by a solar cell to the incident light power at a given wavelength:

$$SR(\lambda) = \frac{J_{SC}(\lambda)}{P_{in}}$$
(2.12)

It can be described in terms of EQE:

$$SR(\lambda) = \frac{q\lambda}{hc} \cdot EQE(\lambda) = 0.808 \cdot \lambda(\mu m) \cdot EQE(\lambda)$$
(2.13)

The theoretical estimated short circuit current is obtained by integrating the product of EQE and photon flux over all relevant wavelength and described as:

$$J_{SC}^{predicted} = q \cdot \int EQE(\lambda) \cdot \Phi_{AM1.5}(\lambda) \,\mathrm{d}\lambda \tag{2.14}$$

with $\Phi_{AM1.5}$ being the photon flux under AM 1.5G spectrum.

If the predicted short-circuit current density $J_{SC}^{predicted}$ derived from the EQE measurement does not match the J_{SC} obtained by the J-V measurement, the light intensity of the solar simulator should be adjusted to one sun intensity (100 mW/cm²) according to the spectral mismatch correction factor.

Spectral Mismatch Correction Factor

It is important to note that solar simulators do not produce the exact AM 1.5G spectrum. Therefore, corrections are made with a spectral mismatch factor M to correct for the light intensity differences between lamp spectrum and AM 1.5G at the given wavelengths, in order to standardize the measurements among different labs. It is calculated with four terms [18]:

$$M = \frac{J_{ref}^{AM1.5}}{J_{dev}^{AM1.5}} \cdot \frac{J_{dev}^{Sim}}{J_{ref}^{Sim}}$$
(2.15)

where the four short-circuit current densities are calculated by multiplying the spectral response of the device (denoted by dev) and a calibrated reference diode (denoted by ref) with the spectral irradiance of the AM 1.5G (denoted by AM1.5) and the solar simulator (denoted by Sim):

$$J_{ref}^{AM1.5} = \int SR_{ref}(\lambda) \cdot I_{AM1.5}(\lambda) \,\mathrm{d}\lambda \tag{2.16}$$

$$J_{dev}^{AM1.5} = \int SR_{dev}(\lambda) \cdot I_{AM1.5}(\lambda) \,\mathrm{d}\lambda \tag{2.17}$$

$$J_{ref}^{Sim} = \int SR_{ref}(\lambda) \cdot I_{Sim}(\lambda) \,\mathrm{d}\lambda \tag{2.18}$$

$$J_{dev}^{Sim} = \int SR_{dev}(\lambda) \cdot I_{Sim}(\lambda) \,\mathrm{d}\lambda \tag{2.19}$$

Dividing the measured short-circuit current density of the test device by the mismatch factor M, the real J_{SC} is obtained.

2.2.5 Efficiency Limits

The theoretical maximum power conversion efficiency of a single junction solar cell is known as the Shockley-Queisser limit (SQ limit) or the detailed balance limit, presented in 1961 by William Shockley and Hans-Joachim Queisser, where they used the emission of a black body with a surface temperature of 6000 K as the approximated solar spectrum [21]. Using the standard AM 1.5G spectrum, Ruehle calculated a maximum efficiency of 33.7% at a band gap of 1.34 eV [15], still referred to as the Shockley-Queisser limit (black curve in Figure 2.5a). Up to now, the highest efficiency for a single junction solar cell is 29.1%, achieved with the thin film GaAs [2], reaching more than 85% of its limit (∇ symbol in Figure 2.5a). The record

Chapter 2: Theoretical Background of Hybrid Organic-Inorganic Metal Halide Perovskite Solar Cells

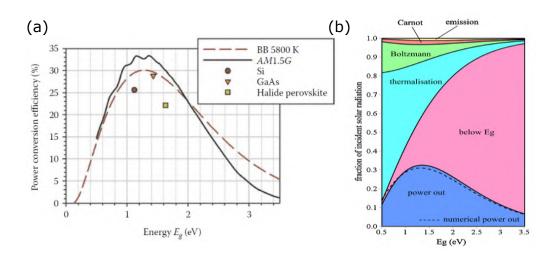


Figure 2.5: (a) The maximum PCE limit for a single junction solar cell operated at 298.15 K and illuminated with the AM 1.5G (solid curve) spectral irradiance and a black body radiation with a surface temperature of 5800 K (dashed curve) as a function of the band gap energy. The denoted points are the efficiency records for Si, GaAs and halide perovskite solar cells in 2018. Adapt with permission from [19] (b) Intrinsic losses of a solar cell. Power out is dependent on the band gap E_g. Adapt from [20] with permission. Copyright 2010 John Wiley & Sons, Ltd.

efficiency of the single junction perovskite solar cell is 25.2% [2], improved nearly by 3% since 2018 (\Box symbol presented in Figure 2.5a).

Figure 2.5b illustrates five types of intrinsic losses occurring in a single junction solar cell. The non-absorption of photons with energy below the band gap and the emission loss (i.e., absorbers are also emitters according to Kirchhoff's law) limit the maximum current of a solar cell, while the thermalisation (i.e., excited carriers relaxing to the band gap edge), Boltzmann and Carnot losses cause reductions in the maximum voltage [20]. In addition, the calculation of the SQ limit did not take non-radiative recombination into consideration, which also reduces the maximum efficiency [15]. There are two types of non-radiative recombination pathways in semiconductors, namely Auger and Shockley-Read-Hall (SRH) recombination. The latter is also known as trap-assisted recombinations. This occurs when localised energy states (i.e., trap states) are created within the band gap by dopants or defects, where the charge carriers are trapped and then recombine non-radiatively with energy exchanged in the form of lattice vibration. Auger recombination involves a third charge carrier, which is excited to a higher energy level and then losses its excess energy to thermal vibrations after the recombination.

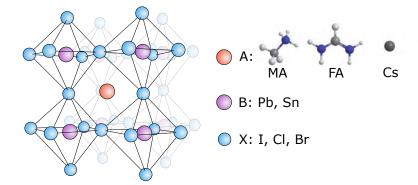


Figure 2.6: Sketch of a cubic ABX₃ perovskite structure. Courtesy of Dr. Alex Taylor. The atomic structures of the A site cations are reproduced with permission from [11]. Copyright 2014, The Royal Society of Chemistry.

2.3 Perovskite Solar Cells

2.3.1 What is Perovskite?

The name 'perovskite' is used to describe any materials with the same crystal structure as calcium titanium oxide (CaTiO₃), which is well known as the perovskite structure. It is first discovered by Gustav Rose in 1839 and named after the Russian mineralogist L. A. Perovski. In general, the chemical formula of perovskite material is ABX₃, with A and B being two different cations with different radii, and X an anion bonding to both.

The perovskite material employed in solar cell application is commonly the hybrid organic-inorganic metal halide perovskite, where A is usually an organic cation (such as methylammonium (CH₃NH₃⁺ or MA⁺) and formamidinium (CH₃(NH₂)₂⁺ or FA⁺), B a divalent metal (e.g. Pb²⁺, Sn²⁺, etc.), and X a halide anion (i.e., Cl⁻, Br⁻ and I⁻) (Figure 2.6). In this work, we focus on two representative types of the organolead halide perovskite. One is the mostly employed CH₃NH₃PbI₃ (MAPbI₃) perovskite (Chapter 4 and 5), and the other is the 'triple cation' perovskite Cs_{0.05}(MA_{0.17}FA_{0.83})_{0.95}Pb(I_{0.9}Br_{0.1})₃ (CsMAFA) (Chapter 6).

2.3.2 Evolution of Perovskite Solar Cells

Miyasaka's group were the first to report the use of organic-inorganic halide perovskite in photovoltaic devices with an efficiency of 3.8% [3], where an organic electrolyte solution was used as the hole transporting material. Following their work, Park and co-workers doubled the efficiency to 6.5% in 2011 [23], by optimizing perovskite coating procedure and electrolyte composition. However, the liquid electrolyte was believed to cause stability issues, and therefore, the research focus turned to preparing solid state perovskite-sensitized solar cells.

In 2012, Park, Grätzel and co-workers boosted the efficiency to 9.7 % [4] by introducing Spiro-OMeTAD (2,20,7,70-Tetrakis[N,N-di(4-methoxyphenyl)amino]-9,90-spi-

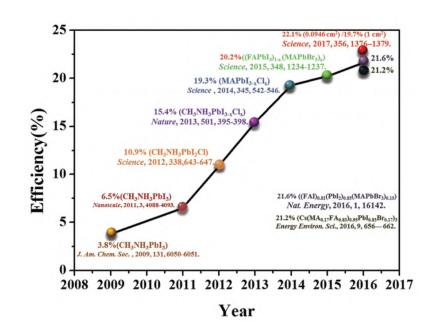


Figure 2.7: Efficiency development of perovskite solar cells from 2008 to 2017. Adapted with permission from [22]. Copyright 2018 John Wiley & Sons, Ltd.

robifluorene) as a solid state hole transporting material, avoiding the problem of dissolving perovskite in the liquid-based solar cells. This is the first prototype of the mesoporous architecture. At about the same time, Snaith and colleagues reported a 10.9% mixed halide $CH_3NH_3PbI_{3-x}Cl_x$ device also using Spiro-OMeTAD as the hole transporting material [5]. They replaced the nanoporous TiO₂ with a non-conducting Al₂O₃ scaffold layer, improved the V_{OC} by a few hundred millivolts and boosting the efficiency to over 10%. This demonstrated that perovskite is able to harvest light and transport both electrons and holes, revealing a border application of perovskite than just as sensitizers in the dye-sensitized solar cells. This study paved the path for a brand new device structure without any metal oxide scaffolds.

In 2013, Snaith's group reported a planar perovskite solar cell via a co-evaporation technique, exhibiting a PCE of 15.4% [24]. They also managed to achieve a device with 11.4% PCE by solution deposition in a planar structure [25]. Both studies demonstrated that the mesoporous oxide layer is not necessary.

In the same year, Chen's group for the first time reported a perovskite solar cell in the inverted architecture with a PCE of 3.9 % [26], where PEDOT:PSS (poly (3, 4-ethylenedioxythiopene)-poly (styrenesulfonate)) was used as the hole transporting material at the transparent conductive oxide side and an electron transporting layer was deposited between the perovskite layer and the metal cathode. PCEs of photovoltaic devices employing this architecture were boosted to over 16 % by improving the perovskite film quality [27–29]. So far, three types of perovskite solar cell architectures, namely mesoporous, planar and inverted structures, have been reported. (Details are described in the Section 2.3.5).

The breakthrough in the PCE exceeding 20% was achieved by Seok's group in 2015 [6], where $CH(NH_2)_2PbI_3$ (FAPbI_3) films were deposited via intramolecular exchange of dimethyl sulfoxide (DMSO) molecules intercalated in PbI₂ with $CH(NH_2)_2I$ (FAI). This methodology was later applied to the fabrication of advantageous mixed cation/mixed halide perovskite (FAPbI_3)_{0.85}(MAPbBr_3)_{0.85} [7]. In 2016, Saliba et al. reported a triple cation perovskite $Cs_{0.05}(MA_{0.17}FA_{0.83})_{0.95}Pb(I_{0.83}$ $Br_{0.17})_3$ by adding 5% caesium (Cs) to the mixed cation/mixed halide perovskite, yielding a PCE of 21.1% with improved stability [8]. As of August 2019, the record efficiency of perovskite solar cells is 25.2%, certified by the National Renewable Energy Laboratory (NREL)[2]. Figure 2.7 summarizes a brief history of the efficiency development of perovskite solar cells from 2008 to 2017 [22].

2.3.3 Material Properties and Compositional Advances

In general, the ABX₃ perovskite structure consists of a 3D BX₆⁴⁻ octahedral sublattice and A⁺ cation filling the empty spaces and balancing the charges. The valence band maximum (VBM) and conduction band minimum (CBM) are determined by the BX₆⁴⁻ lattice, whereas the A⁺ affects the band gap through the lattice expansion or contraction caused by its size variation [30].

The hybrid organic-inorganic metal halide perovskite has a strong optical absorption over the visible spectrum [31], and a tunable band gap from 1.1 eV to 2.3 eV, altered by interchanging the cations [6, 32], metals [33] and halides [10, 11].

CH₃NH₃PbI₃ Perovskite

The band gap of $CH_3NH_3PbI_3$ perovskite is around 1.55 eV [34], close to the optimal band gap of a solar cell, which yields a maximum efficiency of 33.7% at 1.34 eV, estimated by the Shockley-Queisser efficiency limit.

It has been reported that $CH_3NH_3PbI_3$ perovskite has a relatively low exciton binding energy on the order of 16 meV at low temperatures, and even a few millielectronvoltages at room temperature [14, 35, 36], much smaller than the most organic semiconductors. Huang's group has shown the direct evidence for the non-excitonic nature of organolead trihalide perovskite [37], where they measured a giant dielectric constant of over 500 at 20 Hz for $CH_3NH_3PbI_3$. The low exciton binding energy and high dielectric constant enable instant exciton dissociation into free electrons and holes under illumination, and consequently contribute to the excellent photovoltaic performance. The charge carrier diffusion length in $CH_3NH_3PbI_3$ perovskite is at least 100 nm [12, 38] and can be improved to more than 1μ m by increasing Cl content in $CH_3NH_3PbI_{3-x}Cl_x$ [12], which can be attributed to the combination of high charge mobility and low bimolecular recombination [39]. This long diffusion length makes sure that the generated carriers reach the transporting layers before recombination.

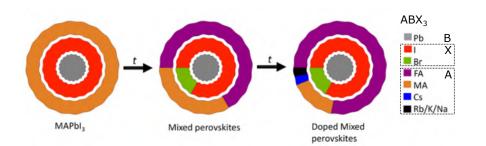


Figure 2.8: Illustration of the compositional trend toward better performing perovskites. Adapted with permission from [41]. Copyright 2018 American Chemical Society.

Mixed Halide Perovskite

The first mixed halide perovskite $CH_3NH_3PbI_{3-x}Cl_x$ was reported by Snaith's group, where they spin coated a mixture solution of CH_3NH_3I (MAI) and $PbCl_2$ (at a molar ratio of 3:1) [5]. Although the CI^- substitution is allowed only at low concentrations (x < 3-4%) [40], it increases the charge carriers mobility and diffusion length without changing the material's band gap or optical properties [12, 39, 40], and consequently enhances the photovoltaic performance [5, 12, 40]. In addition, it has shown to improve the environmental stability of the perovskite as well [5].

The incorporation of Br in $CH_3NH_3PbI_3$ changes the band gap dramatically and by the management of Br percentage in $CH_3NH_3Pb(I_{1-x}Br_x)_3$ ($0 \le x \le 1$), one can manipulate the band gap to cover nearly the entire visible spectrum (from 786 nm to 544 nm) [10].

From Single Cation to Mixed Cation

The effective radius of A^+ cation (in the ABX₃) influences perovskite's band gap by expanding or contracting the BX₆⁴⁻ octahedral cage, rather than determining the band structure [30]. The formability of the 3D perovskite is restricted by the geometric tolerance factor $t = \frac{r_A + r_B}{\sqrt{2}(r_B + r_X)}$ based on the effective ionic radii r_i of A, B and X [42]. Given a particular B and X, in most popular case APbI₃, the 'black phase' perovskite is formed for tolerance factors ranging from 0.8 to 1.0; only MA⁺(CH₃NH₃⁺), FA⁺(CH₃(NH₂)₂⁺) and Cs⁺ meet this condition [43–45].

MAPbI₃ was reported to have three structural phases: a cubic phase (α -phase) above 330 K, a tetragonal phase (β -phase) from 160 K to 330 K, and an orthorhombic phase (γ -phase) below 160 K [33, 46]. The phase transition from the tetragonal to cubic phase at 330 K (around 56 °C) does not affect the device performance [47]. However, the thermal decomposition of the cubic phase MAPbI₃ into PbI₂ at 85 °C (within the range of the operating temperature), even in inert conditions, hampers the long-term device stability [48, 49]. In addition, MAPbI₃ was shown to suffer inevitable degradation upon exposure to moisture, oxygen and light [47, 50–58]. (Details are described in the Section 2.3.6).

FAPbI₃ exhibits a photoinactive yellow hexagonal phase (δ -phase) at room temperature, and a black polymorph α -phase which requires a higher annealing temperature (150 °C) to form, and is retained after cooling down to room temperature [11]. Unfortunately, this black polymorph phase is unstable especially in an ambient humid atmosphere [33, 59]. Nevertheless, FAPbI₃ is a promising candidate to improve the J_{SC}, due to a reduced band bap of 1.48 eV and therefore an extended absorption edge of 840 nm, originating from the longer ionic radius of FA⁺ in comparison to MA⁺ [11, 59].

 $CsPbI_3$, with a band gap of 1.73 eV, forms a photoinactive yellow phase at room temperature and only transforms to the black perovskite phase above $300 \degree C$ [60], making it not suitable for photovoltaic applications.

As a result of the environmental or structural instability of the single cation perovskite, numerous attempts have been carried out with mixed cations and halides, in order to achieve a more stable perovskite structure. For example, the undesirable δ -phase in the FAPbI₃ can be avoided by using a mixed-cation perovskite system. Pellet et al. [32] stabilized the perovskite crystal structure by substituting MA⁺ with FA⁺, and prevented the formation of the δ -phase by the dominating effect of MA⁺ cation, due to the stronger interaction of MA⁺ cation and PbI₆⁴⁻ octahedral cage [61]. Therefore, this double cation perovskite exhibits a reduced band gap with broader absorption range relative to MAPbI₃, and subsequently a boosted J_{SC} [32]. Furthermore, a similar approach with the incorporation of both mixed cations and mixed anions is developed to form (FAPbI₃)_{0.85}(MAPbBr₃)_{0.15}, resulting in a highly crystalline and stable perovskite without the formation of the δ -phase [7].

Furthermore, the partial substitution of Cs^+ for FA^+ in the FAPbI₃ is found to enhance the light and moisture stability along with photovoltaic performance. This can be attributed to the stabilized perovskite lattice structure by the enhanced FA-I interaction, due to the contraction of the unit cell [62], and also the favourable entropy gain [63]. In addition, Cs has been found to improve the stability of MAPbI₃ perovskite [64].

Therefore, Cs is introduced in the $(\text{FAPbI}_3)_{0.85}$ (MAPbBr₃)_{0.15} system and a PCE of 21.1% is achieved by an optimized formulation of the triple cation perovskite $\text{Cs}_{0.05}$ (MA_{0.17}FA_{0.83})_{0.95}Pb(I_{0.83}Br_{0.17})₃ (abbreviated as CsMAFA) [8]. Saliba et al. reported an improved device stability of this CsMAFA perovskite, retaining 85% of its initial PCE after 250 hours under illumination operated at the maximum power point. Later, they incorporated the even smaller cation Rb⁺ into the Cs-MAFA system, and found that the 'quadruple cation' perovskite RbCsMAFA exhibits a stabilized PCE of 21.6% with a V_{OC} of 1.24 V at a band gap of 1.62 eV [65]. Moreover, the choice of A⁺ has been further extended. For instance, Park's group reported an inorganic perovskite solar cell with potassium integration into the Cs-based perovskite [66]. Cations with larger radii ($r_i > r_{FA^+}$) were investigated, such as guanidinium (GA⁺), which was reported to significantly suppress the nonradiative recombination [67].

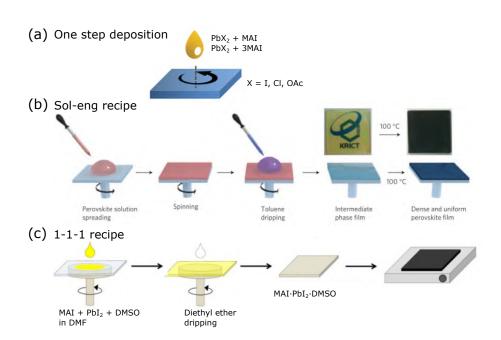


Figure 2.9: Schematic representation of different fabrication methods. (a) One step deposition. Adapted with permission from [73]. Copyright 2016 John Wiley & Sons, Ltd. (b) Sol-eng recipe. Toluene dripping during spin coating the perovskite precursor solution. Reproduced with permission from [74]. Copyright 2014 Macmillan Publishers Limited. (c) 1-1-1 recipe. Diethyl ether dripping during spin coating the perovskite precursor solution (equimolar MAI, PbI₂ and DMSO in DMF). Adapted with permission from [75]. Copyright 2015 American Chemical Society.

2.3.4 Fabrication Methods

Perovskite films can be fabricated by solution processing at low temperatures and also by solution-free processing (e.g. thermal co-evaporation from multiple perovskite precursors [24]). The simplest solution processing method is spin coating, which can be divided into different categories by the choice of perovskite precursors, deposition steps, and whether employing an antisolvent treatment. Other solution processing techniques (e.g. dip-coating [68], blade coating [69], and slot-die coating [70]) and printing techniques (such as ink-jet printing [71]), opens the door for upscaling of perovskite solar cells and modules for commercial application in industry [72].

In this work, we use the spin coating method with a one step recipe, where all the perovskite precursors are mixed together and dissolved in a high boiling point polar solvent (dimethylformamide (DMF), dimethyl sulfoxide (DMSO) or γ -butyrolactone (GBL)), or a mixture of thereof. This precursor solution is spin coated in a single step and the perovskite crystallizes during the thermal annealing process. We also employ the treatment of a droplet of the antisolvent on the wet film during the spin coating process to facilitate the fast formation of a high-quality perovskite film.

One Step Deposition

Early reports on the mesoporous perovskite solar cells were mostly prepared by the one step deposition [4, 5, 76]. However, it results in films with poor morphology when fabricated on planar substrates with a stoichiometric $CH_3NH_3I:PbI_2$ precursor solution, due to the presence of PbI_2 rod-shaped nanoparticles in the precursor solution. It is actually a colloidal dispersion in a mother solution, rather than a pure solution [77]. Yan et al. also showed that adding excess methylammonium halide reduces the colloidal size and improves the coverage and film morphology [77]. Petrov et al. presented that the $CH_3NH_3I:PbI_2$ ratio in the precursor solution determines the final perovskite film morphology, which correlates with the crystal shapes of different intermediate phases [78]. PbI_2 -rich and stoichiometric solutions form a ribbon-like phase that produces a needle-like crystal shape, while the $CH_3NH_3I:PbX_2$ (X = I, Cl, etc.) is commonly used for the single step deposition and the general reaction process is summarized as [79, 80]:

$$PbX_2 + 3CH_3NH_3I \rightarrow CH_3NH_3PbI_3 + 2CH_3NH_3X$$

$$(2.20)$$

The excess of organic cation is removed during perovskite formation by sublimation of the volatile by-product salt (i.e., CH_3NH_3X in Equation 2.20) [80, 81]. Therefore, the thermal property of this by-product salt CH_3NH_3X has a great impact on the annealing time and temperature required for perovskite crystallization, and on the film morphology [79, 80, 82].

In 2015, Zhang et al. introduced the non-halide lead source lead acetate trihydrate (Pb(OAc)₂·3H₂O), which forms the by-product salt CH₃NH₃(OAc) with volatility higher than CH₃NH₃I and CH₃NH₃Cl [82]. The hydration water in the Pb(OAc)₂·3H₂O improves the film morphology and subsequently photovoltaic performance due to the reduced nonradiative pathways [82, 83]. The acetate system reduces the activation energy for nucleination and crystal growth [80], resulting in a much faster crystallization and much smoother perovksite films with less pinholes [82]. Zhang et al. also demonstrated that the addition of hypophosphorous acid (HPA) into the perovskite solution assists the formation of high-quality perovskite films with much reduced non-radiative recombination centres, due to the retarding of I₂ back to I⁻, which improves the stoichiometry and reduces the metallic lead species in the perovskite film [84]. Hence, this much improved film quality leads to an enhanced photoluminescence intensity and subsequently improved photovoltaic performance [84]. This method is abbreviated as the **Pb(OAc)₂ recipe** in the following discussion.

Using Antisolvent

The antisolvent treatment was first introduced by Xiao et al. [85] and Jeon et al. [74] in 2014. During spin coating of the perovskite precursor solution, a droplet of the

antisolvent is dripped onto the rotating sample, resulting in a fast formation of the perovskite film upon annealing. Xiao et al. fabricated a uniform perovskite film with full coverage employing a fast deposition crystallization procedure, which introduces a high density of nucleation sites by dripping chlorobenzene during spin coating the perovskite precursor solution dissolved in DMF [85]. Jeon et al. reported a controlled crystallization of dense and uniform perovskite films using an antisolvent treatment during spin coating of perovskite precursors dissolved in a mixed solvent of DMSO and GBL [74] (Figure 2.9b). At the initial stage of spinning, the wet film is composed of CH₃NH₃I and PbI₂ dissolved in DMSO and GBL. With the rapid evaporation of GBL (GBL here works only as a solvent that evaporates faster than DMSO), the film components begin to concentrate. The dripping of an antisolvent (i.e., toluene) immediately removes the excess DMSO solvent and forms an intermediate phase of CH₃NH₃I-PbI₂-DMSO, with DMSO retarding the fast reaction between CH₃NH₃I and PbI₂ during solvent evaporation. After thermal annealing, this intermediate phase converts into a uniform crystalline perovskite film. This technique is denoted by the sol-eng recipe.

This method was further developed with a better understanding of the Lewis base adduct formation. Solvents with Lewis base properties such as DMSO donate a pair of free electrons to the Pb^{2+} ion, and form a Lewis adduct- $PbI_2(DMSO)$ complex [86]. Iodide (I^-) in CH₃NH₃I also serves as a donor and forms a Lewis adduct with PbI_2 , which explains the increased solubility of PbI_2 in the presence of CH_3NH_3I in DMF. Hence, Ahn et al. proposed a reproducible deposition method employing the Lewis adduct of CH₃NH₃I·PbI₂·DMSO by spin coating equimolar CH₃NH₃I, PbI₂ and DMSO in DMF [75] (Figure 2.9c). A high-quality perovskite film is converted from this adduct film upon annealing due to the removal of volatile DMSO [86]. They compared different solvents and found diethyl ether to be the best in terms of reproducibility. Toluene and chlorobenzene are miscible with DMSO and will violate the stoichiometry of the CH₃NH₃I·PbI₂·DMSO adduct, whereas diethyl ether selectively washes away only DMF and therefore forms a well-defined adduct. They also pointed out the importance to keep the stoichiometric ratio of DMSO, because excess DMSO leads to the formation of extra PbI_2 upon exposure to air, and therefore, results in less stable perovskite films. This method is abbreviated as the 1-1-1 recipe.

Using the PbI₂(DMSO) complex, Yang et al. reported a deposition method to fabricate uniform perovskite films with full coverage by an intramolecular exchange between $CH(NH_2)_2I$ (FAI) and DMSO molecules intercalated in the PbI₂(DMSO) complex, due to the higher affinity of FAI toward PbI₂ relative to DMSO [6]. The smooth morphology of the perovskite film is confirmed by the uniform and dense film of pre-deposited PbI₂(DMSO) complex with unaltered inorganic PbI₂ frame. Later, this methodology is applied to the deposition of mixed cation/mixed halide perovskites, with a breakthrough in PCE of 21.1 % with a triple cation perovskite $Cs_{0.05}(MA_{0.17}FA_{0.83})_{0.95}Pb(I_{0.83}Br_{0.17})_3$, with chlorobenzene used as the antisolvent during spinning [8].

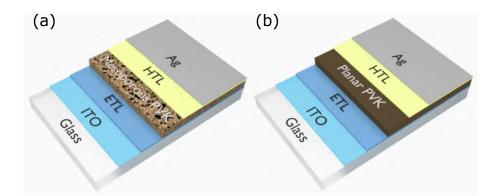


Figure 2.10: Schematic of n-i-p mesoporous (a) and n-i-p planar(b) architectures of perovskite solar cells.

Choice of Antisolvents

The choice of antisolvents in early times were always among toluene [74], diethyl ether [75] and chlorobenzene [8, 87]. Interestingly, these three solvents produce highquality perovskite films with high photovoltaic efficiencies, despite their different physiochemical properties (such as boiling point and polarity (Table B5 in the Appendix)). In 2017, two reports compared a whole series of different solvents and presented two promising new candidates, namely ethyl acetate [88] and trifluorotoluene [89], which yielded champion efficiencies of 19.4% and 20.3%, respectively. One year later, Zhao et al. introduced anisole as an antisolvent for CsMAFA perovskite fabrication, and reached a champion PCE of 19.8% [90]. The use of anisole as the antisolvent expanded the processing window with a dripping time ranging from 5 to 25 s, an antisolvent volume varying from 0.1 to 0.9 ml, and a DMF:DMSO volume ratio ranging from 6:4 to 9:1 [90]. Interestingly, even 'poor' solvents (traditionally considered for perovskite) can yield decent efficiencies. For example, xylene, immiscible with DMF and DMSO, produces a PCE of 17.8% [89]. In addition, a mixed solvent system ('good' + 'poor' solvents) is able to produce a plausible PCE of 19.2%, when 6% isopropanol (also called 2-propanol) is added to chlorobenzene [91], whereas the isopropanol alone gives only 14.8% PCE [88]. Nevertheless, there are no clear requirements for a 'good' antisolvent (such as high boiling point or strong polarity). Therefore, we will elucidate this issue with a systematic study comparing a large number of solvents in the Chapter 6.

2.3.5 Device Architecture and Charge Selective Contacts

The perovskite layer is sandwiched between an n-doped layer that functions as electron selective contact (ESC) and a p-doped layer working as hole selective contact (HSC), where the selectivity is achieved by the conductivity of only one type of the charge carriers and the matched energetic level. The n-i-p or p-i-n solar cells are named according to the layer stacking order, where incident light passes through the ESC or HSC first. The very first perovskite solar cells were fabricated by depositing

Chapter 2: Theoretical Background of Hybrid Organic-Inorganic Metal Halide Perovskite Solar Cells

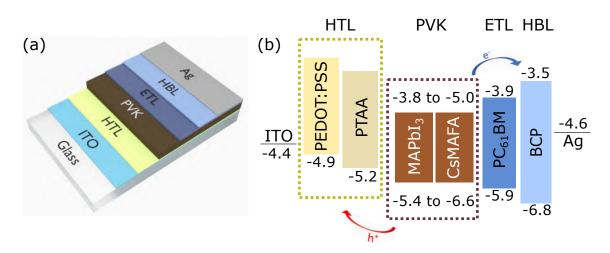


Figure 2.11: (a) Schematic of a planar p-i-n perovskite solar cell. (b) Energy diagram of the anode (i.e., indium tin oxide (ITO)), hole transporting layer (HTL), perovskite layer [92–97], electron transporting layer (ETL), hole blocking layer (HBL) [98] and cathode (i.e., Ag) used in the device architecture (a). [99, 100]

perovskite on a mesoporous electron transport layer (ETL) [3, 4, 23] or an ETL with a porous scaffold [5], whose structure is referred to as the n-i-p mesoporous architecture (Figure 2.10a). The n-i-p planar architecture (Figure 2.10b) describes the type of perovskite solar cells that the perovskite film is coated on a planar ETL.

In this work, we employ the planar p-i-n architecture (also known as the inverted device architecture), where the perovskite film is coated on a planar hole transporting layer (HTL), as illustrated in Figure 2.11a. This device structure originates from the regular architecture used in the organic photovoltaics (OPVs). Therefore, the charge transport materials that are well studied and widely used in OPVs can be utilised in perovskite solar cell as well. Figure 2.11b shows the energy diagram of the anode (i.e., indium tin oxide (ITO)), hole transporting layer (HTL), perovskite layer [92–97], electron transporting layer (ETL), hole blocking layer (HBL) [98] and cathode (i.e., Ag) in this planar p-i-n architecture [99, 100].

Hole Transport Layers (HTLs)

Although functioning HTL-free inverted perovskite cells have been reported [101], the photovoltaic performance is heavily hindered by poor hole extraction due to the mismatch of the work function (WF) between ITO and the valence band of perovskite [100, 102]. Therefore, HTLs are employed between the anode and perovskite layer for better hole extraction with a matched energy level near the valence band of perovskite.

The first inverted perovskite solar cell employed poly(3,4-ethylenedioxythiophene): poly(styrenesulfonate) (denoted by PEDOT:PSS) as the HTL, and had an efficiency of 3.9 % [26]. Recently, highly efficient photovoltaic devices with over 16 % PCE have been produced in the same architecture by improving the perovskite film quality [27–29]. The PEDOT:PSS film is deposited by spin coating followed by a low temperature annealing process, making it ideal for up-scaling and application on flexible substrates. However, it has a work function in a range from 4.9 to 5.2 eV (depending on the ratio of PEDOT to PSS) [103], lower than the literature reported ionization potential values of perovskite (5.4-6.6 eV [92–97]), limiting the V_{OC} of the inverted devices [103, 104]. Another possible reason for the V_{OC} loss for PEDOT:PSS based perovskite solar cells is non-radiative recombination at the PEDOT:PSS/perovskite interface [105, 106]. Moreover, the acidic and hygroscopic nature of the PEDOT:PSS corrodes the ITO electrode [107, 108], and it might react with perovskite, causing degradation and hindering the long-term device stability [109, 110].

Research to improve the V_{OC} of perovskite solar cells has been investigated in three approaches: modifying the PEDOT:PSS (e.g. by adding PSS-Na to the PEDOT:PSS to achieve a higher WF [111]), employing p-type metal oxides with a deep lying valence band (such as NiO_X [104, 110, 112]), and using p-type organic polymers [109, 113–115].

In this work, poly(bis(4-phenyl)(2,4,6-trimethylphenyl)amine) (PTAA) is employed as an alternative to PEDOT:PSS in Chapter 6. The boost of V_{OC} to 1.10 V is partially ascribed to a higher work function of PTAA than PEDOT:PSS [114]. More importantly, the non-wetting surface of the PTAA film results in enhanced grain quality of perovskite films with reduced grain boundaries. Therefore, charge recombination at the HTL/perovskite interface is much reduced, which is the dominating factor in the V_{OC} improvement [114, 115].

In order to overcome the coverage issue of the hydrophilic perovskite precursor solution on the surface of the hydrophobic PTAA layer, a amphiphilic interfacial compatibilizer poly [(9,9-bis(3'-(N,N-dimethylamino)propyl)-2,7-fluorene)-alt-2,7-(9,9-dioctylfluorene)] (PFN-P₂) is coated between the PTAA and perovskitelayer [116]. This ultra thin PFN-P₂ layer modifies the surface energy of the PTAAlayer, and facilitates the formation of a high-quality perovskite film without affectingthe electronic properties of the perovskite.

Electron Transport Layers (ETLs)

[6,6]-phenyl C61 butyric acid methyl ester (PC₆₁BM) has already been used as ETL in the early reported inverted perovskite solar cells [26, 117] to facilitate electron extraction and transport. It has be demonstrated that the photocurrent hysteresis is significantly reduced by the incorporation of PC₆₁BM between the perovskite layer and metal cathode in inverted solar cells [118–121]. Some attribute this reduced hysteresis to the fact that PC₆₁BM passivates any charge trapping states at the surface and grain boundaries of the perovskite layer [118, 119]. Another reason might be the inhibited ion migration by PC₆₁BM through chemically bonding with the under-coordinated atoms of the perovskite [121] or defective halides [119, 122]. In addition, Petrozza's group proposed that iodine ions that migrate to the PC₆₁BM layer chemically dopes the PC₆₁BM, increasing its electron conductivity, which further contributes to the stabilized photovoltaic output and hysteresis-free behaviour [123].

Hole Blocking Layers (HBLs)

With the incorporation of a hole blocking layer (HBL) such as bathocuproine (BCP) between the $PC_{61}BM$ layer and metal cathode, the leakage current is much reduced due to a smoother interface between the perovskite/ $PC_{61}BM$ and cathode, which consequently increases the fill factor of the solar cell [124]. In addition, using metal oxides (such as ZnO) as the HBL prevents iodide infiltration into the metal electrode, and consequently hinders the iodide migration induced degradation of the electrode [125].

2.3.6 Device Stability and Degradation

Unencapsulated perovskite solar cells are reported to be stable up to 500 h under one sun illumination [4] and 1000 h under full sunlight [126] in ambient air, still far from practical and commercial standards. Various causes, such as moisture, oxygen, light, heat, and electric stress, are responsible for the instability and cause degradation. Here we briefly discuss two atmospheric causes - moisture and oxygen.

Moisture induced degradation

Although moisture induced degradation has been reported to cause unencapsulated perovskite to decompose and is responsible for the poor device stability [10], others claim that a proper humidity level (of 30%-35%) facilitates the perovskite crystallization, and therefore improves the film morphology [27, 127, 128]. Due to the hygroscopic nature of MAI [129], the absorption of moisture within grain boundaries helps to merge adjacent grains together, which significantly increases the grain size and suppresses the formation of pinholes, and consequently reduces the nonradiative recombination, enhancing the charge carrier life time and photovoltaic performance [27, 128].

At the initial stages of degradation (e.g. short time exposure to a low moisture concentration), monohydrate MAP $bI_3 \cdot H_2O$ is formed. This process is reversible once it is dehydrated [130]. With prolonged exposure to moisture or higher humidity, dihydrate (MA)₄PbI₆ \cdot H₂O starts to form, often accompanied by the formation of PbI₂, which makes this process partially irreversible [130, 131]. Upon exposure to an even higher H₂O vapour pressure for longer periods, it further decomposes into PbI₂ and volatile MAI, becoming completely irreversible [132]. The three following equations describes these three stages of moisture-induced degradation [130, 132]:

$$CH_3NH_3PbI_3 + H_2O \rightleftharpoons CH_3NH_3PbI_3 \cdot H_2O \tag{2.21}$$

$$(4 - n)CH_3NH_3PbI_3 \cdot H_2O + nCH_3NH_3PbI_3 \cdot H_2O + (2 - n)H_2O \approx (CH_3NH_3)_4PbI_6 \cdot 2H_2O + 3PbI_2$$
(2.22)

$$(CH_3NH_3)_4PbI_6 \cdot 2H_2O \rightarrow PbI_2 + 4CH_3NH_3I_3 + 2H_2O$$
 (2.23)

The choice of the hole-transporting materials plays an important role in the device long-term stability [131]. For instance, hole-transporting materials that employ the hygroscopic lithium salt Li-TFSI (such as Spiro-OMeTAD) should be avoided, as it tends to absorb and introduce water to the perovskite, accelerating the degradation process [131–133]. Habisrentinger et al. presented that using a single-walled carbon nanotube (SWNT) wrapped with a monolayer of a hole-transporting molecular as the HTL dramatically increases the device stability and boosts its lifetime [50].

Oxygen (and Light) induced degradation

Oxygen induced degradation of perovskite has a detrimental effect on the device stability, and is found to degrade perovskite devices much faster (on time scales of minutes to a few hours) in comparison to degradation in humid nitrogen (85% relative humidity) [55]. The oxygen induced degradation mechanism is first proposed by Aristidou et al. [53]. When MAPbI₃ is exposed to light and oxygen, superoxide (O_2^-) is generated through electron transfer from the photo-excited MAPbI₃ to molecular oxygen [53]. This reactive superoxide deprotonates the photo-excited MAPbI₃, leading to its decomposition into PbI₂, methylamine (CH₃NH₂), iodine and water, as described by the following equations [53]:

$$CH_3NH_3PbI_3 \xrightarrow{light} CH_3NH_3PbI_3^*$$
 (2.24)

$$O_2 \xrightarrow{CH_3NH_3PbI_3^*} O_2^- \tag{2.25}$$

$$CH_3NH_3PbI_3^* + O_2^- \xrightarrow{deprotonation} CH_3NH_2 + PbI_2 + \frac{1}{2}I_2 + H_2O$$
(2.26)

In addition, their ab initio study demonstrates that the iodide vacancies are the energetically preferred sites for the superoxide species [134]. Depositing a layer of iodide salt onto the perovskite film significantly lowers the amount of superoxide species and increases the photoluminescence lifetime. Therefore, iodide salt are able to passivate the perovskite film and reduce the superoxide formation, and consequently improve the film and device stability [134]. Other reports have highlighted the importance of efficient electron extraction in reducing the formation of superoxide species [53–55, 135]. Device stability can be much improved if the photon-generated electrons are removed by an efficient ETL before their reaction with O_2 to form O_2^- . We demonstrated that devices using $PC_{61}BM$ as ETL degraded at similar speeds in dry air and nitrogen thanks to its efficient removal of electrons [125].

Aristidou et al. have also shown that light-induced superoxide species form much more on perovskite films with smaller grain size. Therefore, devices with poorer perovskite crystallinity degrade at a faster speed and are less stable [134]. In Chapter 4, we will reveal the correlation of the oxygen induced degradation and the microstructure and grain boundaries in the perovskite film.

3 Experimental Methods

3.1 Overview

This chapter describes the processes involved in fabricating perovskite solar cells: substrate cleaning, precursor solution preparation, film deposition via spin-coating, and thermal evaporation of anode. Then, the main techniques for characterizing perovskite films and devices are described.

3.2 Film and Device Fabrication

Spin coating was the main technique employed for thin film deposition in this thesis, where a small amount of solution with coating materials dissolved is applied on top of the sample, before its rotation at a certain speed. The centrifugal force from the rotation spreads the material evenly and forms a neat thin film following the thermal annealing process. Film thickness depends on rotation speed and solution concentration. Metal contacts were deposited by thermal evaporation. We only used Ag as our cathode.

3.2.1 Substrates

Indium-tin oxide (ITO) coated glass substrates were used for film and device fabrication. They were purchased from PsiOTec Ltd. and pre-patterned with one ITO stripe $(7 \text{ mm} \times 12 \text{ mm})$ centred on the $12 \text{ mm} \times 12 \text{ mm}$ soda-lime glass substrate. The ITO serves as anode in the inverted device and has a sheet resistance of $15 \Omega/\text{sq}$. On each substrate, eight pixels $(3 \text{ mm} \times 1.5 \text{ mm})$ were fabricated, increasing the total number of solar cells taken into statistics. ITO substrates were cleaned sequentially with 2 % Hellmanex detergent, deionized water, acetone and 2-propanol in an ultrasonic bath. After drying with an N₂ gun, they were treated with oxygen plasma at 100 mW for 10 min.

3.2.2 Hole Transport Layers

PEDOT:PSS, short for poly (3, 4-ethylenedioxythiopene)-poly (styrenesulfonate): It was used as purchased. Before spin coating, it was sonicated and filtered with a $0.45 \,\mu\text{m}$ polyvinylidene difluoride (PVDF) filter. Cleaned and plasma-treated ITO substrates were covered with PEDOT:PSS solution, spin-coated at 4000 rpm for 30 s and annealed at 150 °C for 10 min in ambient air.

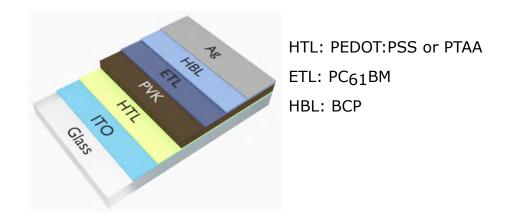


Figure 3.1: Schematic of an inverted perovskite solar cell used in this work.

PTAA, abbreviated for poly(triaryl amine), poly[bis(4-phenyl)(2,4,6-trimethylphenyl)amine]: It was dissolved in anhydrous toluene with a concentration of 1.5 mg/ml and stirred at 70 °C for 2-3 hours until fully dissolved. After cooling down to room temperature, $30 \,\mu$ l of PTAA solution was applied to clean substrates and spin-coated at 2000 rpm for 30 s. Samples were annealed at 100 °C for 10 min in a dry air filled glovebox.

PFN-P₂, short for poly [(9,9-bis(3'-(N,N-dimethylamino)propyl)-2,7-fluorene)alt-2,7-(9,9-dioctylfluorene)]: It was used together with PTAA as a modifier for better wetting and higher voltage output. It was dissolved in anhydrous methanol with a concentration of 0.5 mg/ml and fully mixed by stirring overnight at room temperature. $30 \,\mu$ l of PFN-P₂ solution was spin coated onto PTAA-coated ITO substrates at 5000 rpm for 30 s in a dry air filled glovebox.

3.2.3 Perovskite Layers

Two types of perovskite were studied in the following chapters. One is methylammonium lead triiodine (CH₃NH₃PbI₃). The other is a 'triple cation' perovskite with the composition of $Cs_{0.05}(MA_{0.17}FA_{0.83})_{0.95}Pb(I_{0.9}Br_{0.1})_3$, containing three cations (i.e., Cs^+ , MA^+ and FA^+) and two anions (i.e., Br^- and I^-). They were prepared by spin coating of the perovskite precursor solution and thermal annealing in a dry air filled glovebox (relative humidity < 1.0%).

Single Cation Perovskite CH₃NH₃Pbl₃

Three various fabrication methods were employed to prepare $CH_3NH_3PbI_3$ perovskite films in chapter 4 and 5.

$Pb(OAc)_2$ recipe

The perovskite precursor solution for the $Pb(OAc)_2$ recipe was prepared by dissolving lead acetate trihydrate ($Pb(OAc)_2 \cdot 3H_2O$) and methylammonium iodide (CH_3NH_3I) at a molar ratio of 1:3 in anhydrous dimethylformamide (DMF) with a weight concentration of 40 % and an addition of hypophosphorous acid 50 % aqueous solution $(1.7 \,\mu\text{l} \text{ per 100 mg CH}_3\text{NH}_3\text{I})$. The perovskite solution was spin coated at 2000 rpm for 60 s. After drying at room temperature for 5 min, the as-spun films were annealed at 100 °C for 5 min.

sol-eng recipe

The perovskite precursor solution for the sol-eng recipe was prepared by dissolving PbI₂ and CH₃NH₃I at a molar ratio of 1:1 in a solution mixture of anhydrous γ -butyrolactone (GBL) and dimethylsulfoxide (DMSO) (7:3, v/v), with a weight concentration of 40 %, and stirring at 70 °C for 2-3 hours. The perovskite solution was spin coated first at 1000 rpm for 15 s and then at 4000 rpm for 25 s. During the second step, 600 μ l of toluene was dripped onto the rotating sample 5 s before the end of the second spinning step. The as-spun films were annealed directly at 100 °C for 15 min.

1-1-1 recipe

The perovskite precursor solution for the 1-1-1 recipe was prepared by dissolving PbI₂, CH₃NH₃I and DMSO (1:1:1, molar ratio) in anhydrous DMF, with a weight concentration of 50 %, and stirring at 70 °C for 2-3 hours. The perovskite solution was spin coated at 2000 rpm for 25 s, and 500 μ l of diethyl ether was dripped onto the rotating sample 3s before the end of the spinning step. The as-spun films were annealed directly at 60 °C for 1 min and then at 100 °C for 2 min.

Triple Cation Perovskite $Cs_{0.05}(MA_{0.17}FA_{0.83})_{0.95}Pb(I_{0.9}Br_{0.1})_3$

The precursor solution for the triple cation perovskite contains five different powders of PbI₂, PbBr₂, CsI, methylammonium iodine (MAI) and formamidinium iodine (FAI) mixed at a specific molar ratio in a solution mixture of DMF and DMSO. In order to obtain the target stoichiometry and keep the preparation easy and reproducible, we prepared the stock solution as described in the following method. First, we prepared stock solutions of PbI₂ and PbBr₂ with an actual molarity of 1.155 M (1 M = 1 mol/L) in a solvent mixture of DMF and DMSO (4:1, v:v, denoted by DMF/DMSO). The solutions were heated at 180 °C for 3 min, resulting a fully dissolved clear solution. The CsI stock solution with a molarity of 1.155 M was prepared by adding CsI powder in DMSO and heating at 150 °C for 3 min. The amount of solvent added to obtain the target molarity is calculated by the following steps.

For convenience, we prepared the CsI solution with 2 mol CsI per L of DMSO, PbI₂ and PbBr₂ solutions with 2.5 mol per L in the DMF/DMSO mixture. The density of the stock solution was calculated from the weight of $100 \,\mu$ l of solution at around 25 °C. (Table 3.1) The molarity of the precursor solution can then be calculated from the densities of the solutions and the solvents by:

$$M_{solution} = \frac{mol_{powder}}{V_{solution}} = \frac{mol_{powder} \cdot d_{solution}}{Mw_{powder} \cdot mol_{powder} + V_{solvent} \cdot d_{solvent}}$$
(3.1)

where $M_{solution}$ is the molarity of the precursor solution; mol_{powder} is the number of moles of the precursor; $V_{solution}$ is the volume of the final solution; Mw_{powder} is the

molecular weight of the precursor; $d_{solution}$ is the density of the precursor solution; and $V_{solution}$ is the volume of the added solvent.

After the addition of precursors, the volume of the stock solutions increased compared to the volume of added solvents and the extent of this expansion in volume is calculated by:

$$\gamma_{\rm V} = \frac{V_{solution}}{V_{solvent}} = \frac{a}{M_{eg}} \tag{3.2}$$

where $\gamma_{\rm V}$ is the volume factor listed in Table 3.1 and assumed to be constant in the following calculations; *a* is the number of moles of the precursor per L solvent, and in this case equals 2.5 for PbI₂ and PbBr₂ and 2.0 for CsI, respectively; M_{eg} is the molarity of the stock solution in Table 3.1.

The volume of solvent needed to be added to obtain the target molarity ($M_{solution}$, in this case 1.155 M) was calculated using the following formulae:

$$V_{solvent} = \frac{W_{powder}/Mw_{powder}}{M_{solution}} - V_{exp}$$
(3.3)

$$V_{exp} = \frac{W_{powder}/Mw_{powder}}{a} \cdot (\gamma_{\rm V} - 1) \tag{3.4}$$

where a equals 2.5 for PbI₂ and PbBr₂ and 2.0 for CsI, respectively; W_{powder} is the weight of the precursor solid and Mw_{powder} is the molecular weight of the precursor.

Solvent	-	$egin{array}{c} { m Density} \ { m (g/ml)} \end{array}$	-	-
DMSO	-	1.100 ± 0.001	-	-
DMF/DMSO (4:1, v:v)	-	0.975 ± 0.001	-	-
Solution	Moles per L	$egin{array}{c} { m Density} \ { m (g/ml)} \end{array}$	Molarity (mol/L)	Volume Factor (γ_V)
CsI (DMSO)	2	1.440 ± 0.001	1.778	1.124
$PbI_2 (DMF/DMSO)$	2.5	1.810 ± 0.001	2.127	1.176
PbBr ₂ (DMF/DMSO)	2.5	1.664 ± 0.001	2.198	1.137

Table 3.1: Densities, molarities and volume factors of the solvents and examples of stock solutions.

The final inorganic stock solution $(1.1 \text{ M } \text{Cs}_{0.05} \text{PbI}_{1.75} \text{Br}_{0.3})$ was obtained by mixing the 1.155 M PbI₂, PbBr₂ and CsI stock solutions with a volume ratio of 0.85: 0.15: 0.05.

Methylammonium iodine (MAI) and formamidinium iodine (FAI) powders were weighted out into two separate vials. To obtain the solutions with the stoichiometric formulae of $Cs_{0.05}MA_{0.95}PbI_{2.7}Br_{0.3}$ and $Cs_{0.05}FA_{0.95}PbI_{2.7}Br_{0.3}$, the volume of the $Cs_{0.5}PbI_{1.75}Br_{0.3}$ stock solution to add was calculated by this formula:

$$V_{stocktoadd} = \frac{W_{powder}/Mw_{powder}}{M_{organic/Pb} \cdot M_{Pb}}$$
(3.5)

where $M_{organic/Pb}$ is the molar ratio of FA or MA to Pb and equals 0.95; M_{Pb} is the molar concentration of Pb and equals 1.10. It is important to note that the volume factors of FAI and MAI dissolving in the inorganic stock solution calculated by Equation 3.2 are approximately the same (Table B1, Appendix).

Finally, the perovskite solution with the stoichiometric formula of $Cs_{0.05}(MA_{0.17} FA_{0.83})_{0.95}PbI_{2.7}Br_{0.3}$ was obtained by mixing the $Cs_{0.05}MA_{0.95}PbI_{2.7}Br_{0.3}$ and $Cs_{0.05}FA_{0.95}PbI_{2.7}Br_{0.3}$ solutions in a 1:5 volume ratio.

The perovskite solution was spin coated first at 1000 rpm for 12 s and then at 6000 rpm for 28 s. During the second step, $200 \,\mu$ l of antisolvent was dripped onto the rotating sample 5 s before the end of the spinning step. The as-spun films were annealed directly at 100 °C for 30 min. Solvents used as antisolvent are listed in Table 3.2.

3.2.4 Electron Transport Layers

 $\mathbf{PC_{61}BM}$, short for [6,6]-phenyl C61 butyric acid methyl ester: It was dissolved in anhydrous chlorobenzene with a concentration of 20 mg/ml and stirred at 70 °C for 2-3 hours until fully dissolved. After cooling down to room temperature, the $\mathbf{PC_{61}BM}$ solution was filtered through a 0.25 μ m polytetrafluoroethylene (PTFE) filter before use. 20 μ l was then dynamically spin coated on the perovskite films at 2000 rpm. Afterwards, the samples were annealed at 100 °C for 10 min in a N₂ filled glovebox.

BCP, short for bathocuproine: It was dissolved in anhydrous 2-propanol with a concentration of 0.5 mg/ml and stirred at 70 °C overnight. The solution was prepared and stored in the N₂ filled glovebox. BCP solution was dynamically spin coated at 4000 rpm after PC₆₁BM deposition.

3.2.5 Materials

We listed all materials and solvents used in this thesis in Table 3.2.

ber.		<u></u>				
Chemical	Producer	CAS	Product number			
Perovskite Precursors						
CH_3NH_3I (MAI)	Greatcell Solar	14965-49-2	MS101000			
$CH(NH_2)_2I$ (FAI)	Greatcell Solar	879643-71-7	MS150000			
CsI	abcr GmbH	7789-17-5	AB207757			
$Pb(OAc)_2 \cdot 3H_2O$	Sigma-Aldrich	6080-56-4	316512			
PbI_2	TCI Deutschland	10101 - 63 - 0	L0279			
$PbBr_2$	TCI Deutschland	10031-22-8	L0288			
Solvents for Perovskite Solution						
Dimethylformamide (DMF)	Sigma-Aldrich	68-12-2	227056			
Dimethyl sulfoxide (DMSO)	Sigma-Aldrich	67-68-5	276855			
γ -Butyrolactone (GBL)	Sigma-Aldrich	96-48-0	B103608			
Solvents used as Antisolvent						
Anisole	Sigma-Aldrich	100-66-3	296295			
1-Butanol	Sigma-Aldrich	71-36-3	281549			
Butyl acetate	Sigma-Aldrich	123-86-4	287725			
Chlorobenzene	Sigma-Aldrich	108-90-7	284513			
Chloroform	Sigma-Aldrich	67-66-3	288306			
Diethyl ether	Sigma-Aldrich	60-29-7	346136			
1,2-Dichlorobenzene	Sigma-Aldrich	95-50-1	240664			
Ethanol	Sigma-Aldrich	64-17-5	32205			
Ethyl acetate	Sigma-Aldrich	141-78-6	270989			
Methanol	Sigma-Aldrich	67 - 56 - 1	322415			
2-Propanol	Sigma-Aldrich	67-63-0	278475			
Toluene	Sigma-Aldrich	108-88-3	244511			
Trifluorotoluene	Sigma-Aldrich	98-08-8	547948			
p-Xylene	Sigma-Aldrich	106-42-3	296333			
Organic Polymers						
PEDOT:PSS	Heraeus	-	Clevios P VPAI 4083			
PTAA	Sigma-Aldrich	1333317-99-9	702471			
$PFN-P_2$	1-Material	889672-99-5	OS0995			
$PC_{61}BM$	Solenne BV	160848-22-6	-			
Bathocuproine	Sigma-Aldrich	4733-39-5	699152			
(BCP)	0					
Others						
Hypophosporous acid (HPA)	Sigma-Aldrich	6303-21-5	1046330500			

Table 3.2: List of all chemicals, together with the producer, CAS and product number.

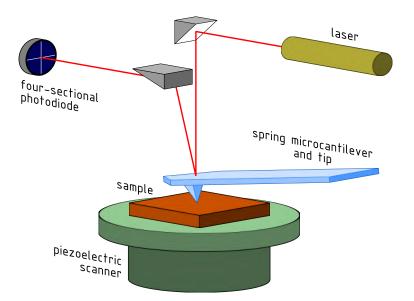


Figure 3.2: Schematic of an AFM. Reproduced with permission from [136]. Copyright GregorioW, licensed under Creative Commons Attribution-Share Alike 3.0 Unported license.

3.3 Film and Device Characterization

3.3.1 Profilometery

Layer thicknesses were measured with a Veeco Dektak 150 profilometer. A diamond stylus with a diameter of $12.5 \,\mu\text{m}$ comes into contact with a sample, and then scans the surface with specified contact force (1-15 mg) in a specified direction. The profilometer then records the vertical position of the stylus, which is then tracked and displayed in absolute values. In order to measure the thickness of a film on the substrate, the film was scratched without any damage to the substrate. Film thickness was determined by scanning the stylus across this scratch.

3.3.2 Atomic Force Microscopy (AFM)

AFM was performed with a Bruker MultiMode microscope in tapping mode in air with silicon tips (Bruker NTESPA) to study the surface morphology of the perovskite films. A cantilever with a very small radius tip is scanned over the sample surface within certain distance, where the cantilever is first attracted by Van-der-Waals forces and then repelled due to Pauli-principle. In the tapping mode, the cantilever is driven to oscillation by a piezo crystal close to its resonance frequency. When the tip rasters across the sample, affected by the attractive and repulsive forces, the resonance frequency of the cantilever is reduced. Consequently, the oscillation amplitude at the constant driving frequency is attenuated. The movement of the cantilever is monitored by the reflection of a laser off of the top of the cantilever by

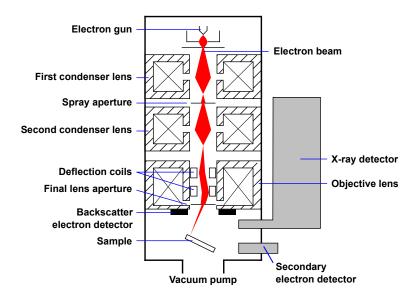


Figure 3.3: Schematic of an SEM. Reproduced with permission from [137]. Copyright Steff, ARTE and MarcoTolo, licensed under Creative Commons Attribution-Share Alike 1.0 Generic license.

a segmented photo diode, and fed back to the cantilever to maintain constant force (Figure 3.2). The average grain size of perovskite films in Chapter 5 was estimated using the software Gwyddion.

3.3.3 Scanning Electron Microscopy (SEM)

In typical scanning electron microscopy (SEM), a thermionically emitted electron beam generated from an electron gun is focused on the sample by one or two magnetic condenser lenses to a spot about 0.4 to 5 nm in diameter. The interaction of the electron beam with the sample results in reflected and backscattered highenergy electrons (BSE), emission of low-energy secondary electrons (SE), characteristic X-rays and light (cathodoluminescence), which can be detected by specialised detectors. The SEs originate from the inelastic scattering of beam electrons with the conduction, or valence, band of the atoms, and has lower energy than the BSE, which is a result of elastic collisions of electrons and specimen atoms. Due to its low energy, SEs emerge from within a few nanometers below the sample surface, and provide high-resolution topographic images of the sample.

The SEM images in this work is taken in the secondary electron imaging mode. The SEM measurements in Chapter 4 were performed by Dr. Paul Fassl, using an Ultra FE-SEM Gemini Ultra 55 (Zeiss) microscope with a working distance of 3 mm and an acceleration voltage of 1.6 kV. The SEM measurements in Chapter 5 were carried out by Dr. Paul Fassl, using an JSM-7610F FEG-SEM (Jeol) with a working distance of 2 mm and an acceleration voltage of 1.5 kV. The SEM measurements in Chapter 6 were performed by Dr. Fabian Paulus, using an JSM-7610F FEG-SEM

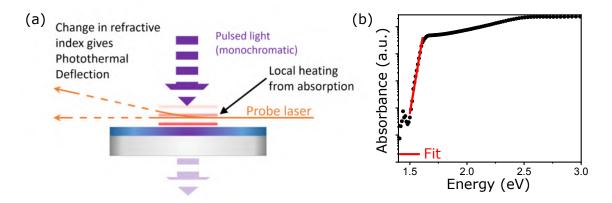


Figure 3.4: (a) Schematic of working principle of PDS. Courtesy of Dr. Paul Hopkinson. (b) Absorption spectra of a pristine CH₃NH₃PbI₃ perovskite sample. The absorption edge near the band gap is fitted linearly (red line) to estimate the Urbach energy.

(Jeol) with a working distance of 2 mm and an acceleration voltage of 1.5 kV.

3.3.4 X-Ray Diffraction (XRD)

X-ray diffraction (XRD) measurements were conducted by Bruker, Karlsruhe as part of demonstration measurements on a D8 ADVANCE DaVinci diffractometer equipped with a PILATUS3-100K detector. 2D-XRD measurements of thin perovskite films were collected in a stepwise $\theta/2\theta$ scan from 0-30°. After combining the individual frames to a full 2D diffraction map, a central slit profile was used to obtain the 1D diffraction profile, followed by background correction and stripping the K_{β} line using the Bruker DIFFRAC.EVA Software.

3.3.5 UV-vis Spectroscopy (UV-vis)

Optical absorption spectra were measured using a Jasco UV-660 spectrophotometer. The monochromated light passes through the sample, partly absorbed, and the transmitted light is then detected by a photodiode. A baseline correction was performed prior to measurements, where the absorption of substrate was subtracted.

3.3.6 Photothermal Deflection Spectroscopy (PDS)

Photothermal deflection spectroscopy (PDS) employs a mirage-like effect to very precisely measure absorption of thin films. The refractive index adjacent to the sample surface changes due to heating of the sample by absorption of light. Light radiated by a 150 W Xenon short-arc lamp (Ushio) is sent through a monochromator (Cornerstone 260 Oriel) to generate a monochromatic pump beam, which is absorbed by the sample and generates a pulsed heat wave. The heat changes the refractive index of an inert liquid (Fluorinert, FC-770 (IoLiTec)) surrounding the sample, and consequently causes a deflection of a probe He-Ne-laser (REO) close to the sample surface (Figure 3.4a). This deflection was monitored by a position-sensitive-detector (Thorlabs, PDP90A) and a lock-in amplifier (Amatec SR 7230).

Spectrosils from UQG Optics were used as substrates for all PDS measurements. In Chapter 4, PDS was employed to determine Urbach energies of the perovskite films. The Urbach energy E_U can be estimated by fitting the exponential Urbach tail in the range of the measured absorption edge near the band gap (Figure 3.4b).

$$A(E) = V \cdot exp(\frac{E}{E_U}) \tag{3.6}$$

where A is the absorption, V is a normalisation factor and the slope of the absorption edge in the semi-log is the inverse Urbach energy E_U .

The PDS measurements in Chapter 4 were performed by David Becker-Koch and Alexandra Bausch. The experimental setup was built by Stefan Lannig and David Becker-Koch [138, 139].

3.3.7 Photoemission Spectroscopy

Photoemission spectroscopy (PES) is based on the photoelectric effect, where electrons can be emitted when photons with the right amount of energy strike a material. Depending on the ionization energy, PES is divided into X-ray photoemission spectroscopy (XPS) and ultra-violet photoemission spectroscopy (UPS). Regardless of the incident photon beam, the PES is carried out in a ultra high-vacuum (10^{-10} mbar) system (Thermo Scientific ESCALAB 250Xi).

X-Ray Photoemission Spectroscopy (XPS)

XPS measurements were performed by irradiating the sample with an X-Ray beam generated by XR6 monochromated Al K α source ($h\nu = 1486.6 \text{ eV}$) with a pass energy of 20 eV. The kinetic energy of the emitted electrons from the core levels of specimen atoms is measured. Each element has a characteristic XPS spectrum, with peaks at certain binding energies corresponding to the configuration of electrons in the atoms, where the peak intensity (i.e., the number of detected photoelectrons at the specified peak) is related to amount of the element within the sampling region. To determine the chemical composition, the atomic percentage of each element is calculated by fitting the respective peak with a convolution of Gaussian and Lorentzian functions, and corrected by the relative sensitivity factor (RSF) (except for hydrogen, as it cannot be detected). The photoelectrons undergo inelastic collisions before escaping the sample and being detected, which exponentially attenuates the number of escaped photoelectrons with depth. Therefore, XPS is a surface-sensitive technique and only detects the top 10 nm of the sample.

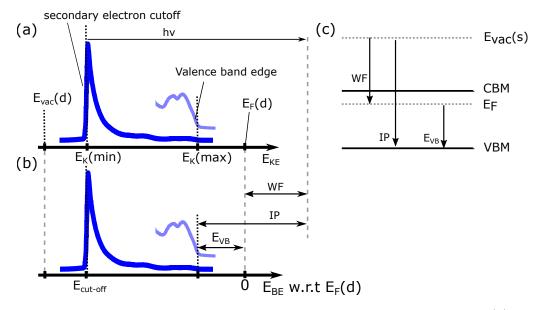


Figure 3.5: UPS spectrum of a pristine CH₃NH₃PbI₃ sample illustrated (a) in the scale of kinetic energy, (b) in the scale of binding energy with respect to the Fermi level. (c) Schematic of an energy band diagram illustrating the relevant energies.

Ultra-Violet Photoemission Spectroscopy (UPS)

UPS measurements were carried out using an He discharge lamp ($h\nu = 21.2 \text{ eV}$, He $I\alpha$) and a pass energy of 2 eV. The UPS spectrum shows the density of states, from which the work function (WF) and the ionization potential (IP) can be determined.

UPS measures the kinetic energy of the emitted electrons that are excited from the valence band. Electrons that have enough energy to escape the specimen, are emitted with a kinetic energy of vacuum level of the sample or above. Electrons that just escape the sample have the lowest kinetic energy ($E_K(min)$), which is the secondary electron cutoff (SECO) measured by the detector. Therefore, $E_K(min)$ is determined by the difference of the vacuum level of the sample $E_{vac}(s)$ and that of the detector $E_{vac}(d)$. In comparison to the primary electrons, secondary electrons undergo inelastic scattering before leaving the sample surface and produce a peak above the SECO in the emitted spectrum. The maximum kinetic energy $E_K(max)$ is the energy of electrons at the Fermi level for metals, or at the valence band maximum (VBM) for semiconductors.

In order to include this SECO in the spectrometer response, the sample is usually biased by applying a negative voltage (e.g. -10 V). This negative biasing ensures that the vacuum level of the sample lies above the vacuum level of the detector. Typically, the Fermi level of the detector $E_F(d)$ is obtained by measuring the Fermi edge of a clean metal sample (e.g. 31.2 eV for -10 V bias), and set to zero in the scale of binding energy. The conversion from the kinetic energy to binding energy is described as:

$$E_{BE} = h\nu - E_{KE} - E_F(d) \tag{3.7}$$

The WF is determined as the energy difference between the vacuum level and the Fermi level of the detector. The IP is obtained by the sum of the WF and the difference between Fermi level and the valence band edge E_{VB} , determined from the x-intercept of the linear fit of valence band onset (Figure 3.5b).

$$WF = h\nu - (E_F(d) - E_k(min))$$
(3.8)

$$IP = WF + E_{VB} \tag{3.9}$$

The UPS measurement collects the kinetic energy of the photoemitted electrons at some distance away from the sample surface. In order to avoid the dependence on the measurement system or the applied bias, the UPS spectra are normally plotted in the scale of binding energy with respect to (w.r.t) the Fermi level, as shown in Figure 3.5b and 3.6d.

3.3.8 XPS and UPS Mapping

In Chapter 5, XPS and UPS mapping were performed on $CH_3NH_3PbI_3$ films deposited on ITO/PEDOT:PSS substrates across the centre area of the samples. Total mapping area is $7.2 \text{ mm} \times 7.2 \text{ mm}$, and contains the active areas of the eight pixels of the final device (Figure A10b, Appendix). The measurements were performed in the dark and all samples had the same exposure time to vacuum and X-ray or UV light.

In order to generate an XPS map with surface compositional information, a serial acquisition of 324 XPS measurements is carried out, with an X-ray spot size of 400 μ m and a step size of 400 μ m. At each measurement point, Pb 4f and I 3d spectra are collected, and the I/Pb ratio is then calculated from the ratio of the atomic percentages of iodine and lead. An I/Pb map summarizes the distribution of the I/Pb ratio across the sampled area as a two-dimensional colour contour plot, with x- and y- axes representing the probed position and the colour difference representing the changes in surface composition.

UPS mapping experiments were performed in a similar fashion to the XPS mapping. UPS spectra are collected across the sample with a measurement spot size of around 400 μ m. At each probed spot, spectra containing information about the secondary electron cutoff and the valence band edge are collected, and the work function and IP are calculated in the way explained in Section 3.3.7. An IP map (two-dimensional colour contour plot with colour changes representing the IP variations) depicts electronic structure distribution across the sampled area.

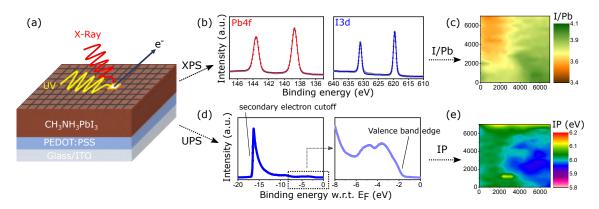


Figure 3.6: (a) Illustration of the experimental approach of XPS and UPS mapping.
(b) Pb 4f and I 3d spectra collected at each point. (c) I/Pb map, a two-dimensional plot of iodine to lead (I/Pb) ratio calculated by dividing the atomic percentage of iodine by the atomic percentage of lead for all measured points over the entire sample. (d) UPS spectra collected at each point. (e) IP map, a two-dimensional plot of ionization potential (IP) for all measured points over the entire sample. Adapted with permission from [140]. Copyright 2018, American Chemical Society.

3.3.9 Photoluminescence Quantum Efficiency (PLQE)

Photoluminescence quantum efficiency (PLQE) was measured using an integrating sphere to collect all emitted light. An integrating sphere (also known as an Ulbricht sphere) is a hollow sphere with its interior covered with a white diffuse coating. Light rays incident on the inner surface are scattered multiple times and redistributed isotropically, regardless of the original directions. Hence, light scattered from different locations on the inside of the sphere contributes equally to the measured spectrum. The experimental set up introduced by de Mello et.al.[141], is shown in Figure 3.7. A laser at a wavelength of 532 nm and a spot size of 2 mm is directed into the sphere, and all light is collected by an optical fiber that connects to a spectrometer. Directly in front of the fibre is a baffle coated with the same diffuse material, to prevent direct illumination of incident light.

Three measurements are made using the sphere. The first measurement (Figure 3.7a) measures the laser intensity in the empty sphere. For the second measurement ('off-axis'), the sample placed inside the sphere is out of the laser incoming path and only absorbs the scattered light (Figure 3.7b). For the third measurement ('on-axis'), the laser beam is now directed on to the sample (Figure 3.7c). There are two types of light striking the sample: direct and scattered light from the inner sphere which are reflected or transmitted by the sample. The PLQE can then be calculated with the following formula [141]:

$$\eta = \frac{P_{on} - (1 - A) \cdot P_{off}}{L_{empty} \cdot A}$$
(3.10)

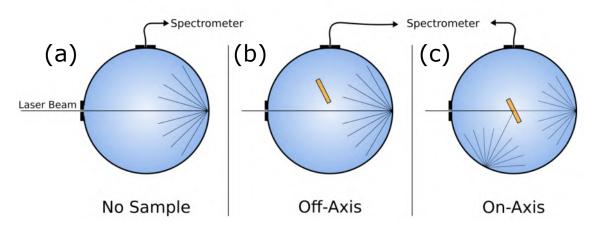


Figure 3.7: Schematic illustration of three configurations of the sphere for the PLQE measurement introduced by [141]. Reproduced with permission from [142]

with

$$A = 1 - \frac{L_{on}}{L_{off}} \tag{3.11}$$

where A is the fraction of the incident light absorbed by the sample in the on-axis measurement; L_{empty} , L_{off} and L_{on} are the areas under the laser peaks for empty, off-axis and on-axis measurements, respectively; P_{off} and P_{on} are the areas under the PL emission peaks for off-axis and on-axis measurements.

The atmosphere in the sphere can be changed by continuously flushing, for example in our setup with nitrogen or dry air.

The PL measurements in Chapter 4 were performed by Alexandra Bausch.

3.3.10 Solar Simulator and J-V measurement Setup

The current density-voltage (J-V) measurements were performed under the simulated sunlight with the Abet Class AAA 3000 sun simulator, which provide the light output matching the ASTM G-173-03 Standard Testing Conditions, 100 mW/cm^2 irradiance and AM 1.5G spectrum. The light intensity was corrected by a spectral mismatch factor, which was obtained by measuring the spectral response of the perovskite solar cell and a Si reference cell (NIST traceable, VLSI), and the spectral irradiance of the solar simulator and the AM 1.5G solar spectrum. The photovoltaic devices were measured with a Keithley 2400 Source Measure Unit and scanned from 1.2 V to 0 V and back, with a step size of 0.025 V and an equilibrium time of 0.1 s after holding under illumination at 1.2 V for 2 s. The active pixel size is 4.5 mm².

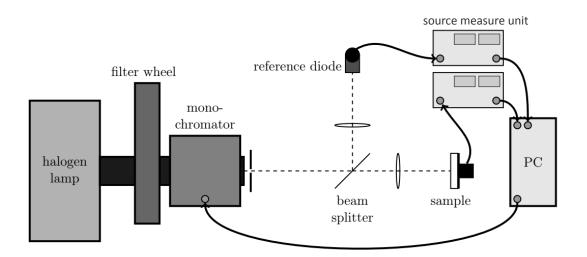


Figure 3.8: Schematic illustration of the EQE setup. Adapt with permission from [143]

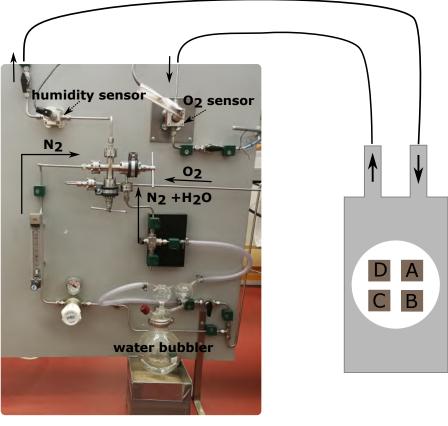
3.3.11 External Quantum Efficiency (EQE)

Light irradiated from a halogen lamp was monochromated in a range from 375 to 820 nm and split into two parts by a beam splitter. One part was focused on one pixel of a photovoltaic device and the other on a reference diode which was calibrated with a NIST-traceable Si diode (Thorlabs) prior to measurements (Figure 3.8). The signals from the illuminated pixel and the reference diode were measured with two Source Measure Units, respectively. A Labview program was then used to evaluate these signals in order to obtain the EQE spectrum.

3.3.12 Environmental Rig

The environmental rig and the sample holder was designed and built by A. Weu and Dr. Paul Hopkinson [144]. The environmental rig is flushed with nitrogen (Purity: 99.999%) as the carrier gas, and is mixed with oxygen (Purity: 99.99%) to obtain the desired O₂ percentage by adjusting the relative flow rate of O₂ to N₂. In order to control the level of humidity, the incoming N₂ gas travels through a water bubble where the humidity is introduced and can be adjusted by the ratio of humid and dry N₂. The oxygen and humidity levels are monitored by different sensors (oxygen sensor: Cambridge Sensotet, Rapidox 2100; humidity sensor: Michell instruments, Easy Dew) continuously before or after being connected to the sample holder, which can hold four samples at the same time.

For all the oxygen degradation studies, the perovskite films were stored in a nitrogen filled glovebox after fabrication, and transferred to a sealed sample holder without exposure to moisture or oxygen. All the samples were degraded at various oxygen levels under simulated AM 1.5G sunlight at $100 \,\mathrm{mW/cm^2}$ irradiance (Abet



Environmental rig

Sample holder

Figure 3.9: Image of the self-built environmental rig that flushes the sample holder with controlled atmosphere.

Sun 3000 Class AAA solar simulator) for 10 h.

4 Oxygen and Light Induced Photodegradation

4.1 Overview

This chapter is based on our recent publication in the journal Advanced Energy Materials with the title 'Role of Microstructure in Oxygen Induced Photodegradation of Methylammonium Lead Perovskite Films' [58]. Most graphs are reproduced or adapted with permission from Wiley VCH. This project was performed in collaboration with Dr. Paul Fassl, who fabricated the samples utilizing $Pb(OAc)_2$ recipe and performed SEM measurements. The PDS measurements were performed by David Becker-Koch and Alexandra Bausch. The PL measurements were performed by Alexandra Bausch.

4.2 Introduction

When we started this study at the beginning of 2016, although the perovskite community had already reported on device stability issues caused by various reasons, such as oxygen, humidity, UV light, heat and electrical stress [47, 51, 52, 145], the main cause for the perovskite instability was thought to be the degradation of CH₃NH₃PbI₃ (MAPbI₃) into PbI₂ under humid atmosphere based on numerous studies [50, 130, 131, 135, 146–151]. At that time, some pioneering papers [53–57] discussed oxygen-induced degradation under light exposure and revealed the degradation mechanism to be deprotonation of methylammonium cations by superoxide (O_2^-) , formed via the transfer of photogenerated electrons from perovskite to oxygen molecules. However, their study were carried out on complete solar cell devices, where limited information on the degradation of the actual perovskite layer was reported.

Therefore, we carried out the degradation experiments under precisely controlled oxygen levels (0-20 %) under 1 sun illumination. In order to understand the progression of oxygen induced degradation of bare perovskite films, we studied their optical, compositional and structural characteristics by UV-vis, PDS, XPS and SEM measurements. By completing the degraded films into solar cells, we correlated the degradation of active perovskite to the decay in photovoltaic performance. We utilised three deposition methods for fabrication of $CH_3NH_3PbI_3$ films, which allowed us to include the effects of perovskite microstructure on the progression of oxygen induced degradation.

4.3 Degraded Films in Oxygen under Illumination

The ETL (in most cases $PC_{61}BM$) in the inverted devices has been reported to serve as a blocking layer preventing oxygen diffusion [55, 152], whereas the metallic cathode could complicate the degradation by iodide diffusion [153, 154]. In another project, we studied the degradation of complete devices under light in various atmospheres and observed that the degradation rate in 20 % O₂ is comparable to in N₂, due to the protection of $PC_{61}BM$ from oxygen induced degradation [125]. Therefore, in order to study the progression of oxygen induced degradation on actual perovskite, we degraded $CH_3NH_3PbI_3$ films coated on ITO/PEDOT:PSS substrates for XPS, SEM and PV studies and on spectrosils for UV-vis and PDS measurements.

The CH₃NH₃PbI₃ films were fabricated by the Pb(OAc)₂, sol-eng and 1-1-1 recipes in a drybox (glovebox filled with dry air, RH < 1.0%) and directly transferred to a N₂-filled glovebox. Samples were then loaded into a sealed environmental box and exposed to various levels of O₂ under 1 sun illumination for 10 hours. The environmental box is flushed with N₂ and O₂, where oxygen levels were controlled by adjusting the relative flow rate of O₂ to N₂ and monitored throughout the experiment (Figure A1 in the Appendix shows an example of a degradation in 10% O₂). Details about film preparation are described in Section 3.2.3.

4.3.1 Optical Properties

To probe the effect of degradation on the bulk properties of degraded films, UVvis and PDS measurements were carried out. Figure 4.1a-c present the absorption of pristine and degraded perovskite films deposited on spectrosils using the Pb(OAc)₂, sol-eng and 1-1-1 recipes. Pristine films prepared by all three methods exhibit a clear onset at 780 nm from the perovskite film [4]. After degradation, the absorption of all films decreases with increasing oxygen concentration. In order to quantify the reduction in absorption, the evolution of absorbance at 700 nm is plotted in Figure 4.1d, with the 0 in the x-axis representing the pristine samples. Absorption of Pb(OAc)₂ films drops only slightly at low oxygen levels. When O₂ is above 12%, the absorption decreases dramatically, together with the appearance of an onset at 520 nm which corresponds to the formation of PbI₂ [53, 55]. For sol-eng films, the absorption drops drastically even for low levels of oxygen and the PbI₂ onset is already present at 5% O₂. The 1-1-1 recipe is the most stable, where the absorption of degraded films remains quite similar to the pristine and only starts to decrease slightly at 20% O₂.

We performed PDS measurements to investigate the change in optical absorption edge and sub-band gap upon degradation [82, 84, 155–158]. Urbach energy (E_u), calculated by fitting the exponential increase in absorption at the band edge (fitting range: 1.50-1.58 eV), represents the degree of energetic disorder in the bulk films [96, 156, 159]. Figure 4.2 presents the Urbach energies of degraded perovskite films after exposure to light and oxygen for 10 h. The Urbach energy for degraded Pb(OAc)₂ samples is unchanged up to 10% at around 18.5 meV, then rises to above 22.5 meV

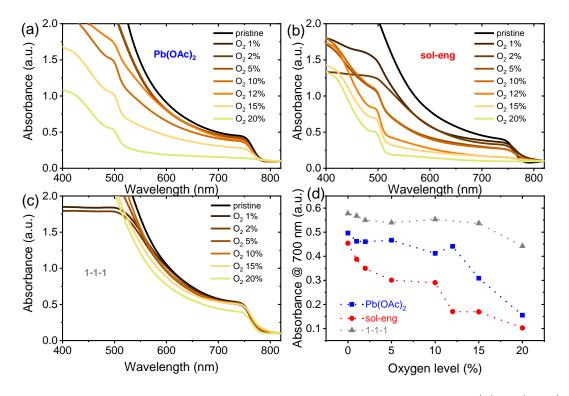


Figure 4.1: UV-vis measurements of degraded CH₃NH₃PbI₃ films for (a) Pb(OAc)₂,
(b) sol-eng and (c) 1-1-1 on spectrosils. (d) Absorbance at 700 nm determined by (a-c) for degraded perovskite films. Adapted with permission from [58]. Copyright 2017, Wiley-VCH.

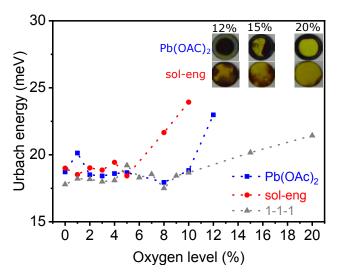


Figure 4.2: Urbach energies of degraded CH₃NH₃PbI₃ films for Pb(OAc)₂, sol-eng and 1-1-1 on spectrosils determined by PDS measurements. The measurements were performed by David Becker-Koch and Alexandra Bausch. Adapted with permission from [58]. Copyright 2017, Wiley-VCH.

at 12 % O_2 and cannot be determined from PDS spectra as films for levels above 12 % are dominated by PbI₂ (see pictures in inset of Figure 4.2 and the onset at 520 nm in Figure 4.1a). For the sol-eng degraded samples, E_u starts to increase at 8 % reaching around 22 meV, and continues to rise until 24 meV for 10 % O_2 . Degraded samples convert to PbI₂ above 10 % O_2 and therefore no value could be determined, while for the 1-1-1 recipe, E_u remains around 18 meV up to 10 % O_2 , increases slightly to 20 meV and 21 meV for 15 % and 20 %, respectively.

We noticed the discrepancy between the unchanged E_u and the fast reduction in the absorption of degraded sol-eng films at low oxygen levels. One possible explanation is that the Urbach energy draws information only about the sub-band gap absorption from perovskite, not the degradation by-products (e.g. PbI₂). Although the absorption intensity drops at low oxygen levels and possible early degradation by-products are formed, degraded films still exhibit high-quality perovskite domains suggested by the steep absorption onset, and therefore the charge carrier generation and separation remain quite efficient. At 8% for sol-eng and 12% for Pb(OAc)₂, Urbach energy rises as a result of the increased energetic disorder in the degraded films, which hampers the electronic properties of the final devices. For even higher oxygen levels, no onset could be determined for sol-eng and Pb(OAc)₂, as the films degraded to PbI₂ indicated by the dominating PbI₂ feature in absorption spectra.

4.3.2 Surface Compositional Properties

In order to investigate the degradation mechanism of perovskite, XPS measurements were performed. The chemical composition evolution is revealed by core level spectra of the main elements (I, Pb, N, C and O) for each degraded sample. As XPS is a surface sensitive technique that only probes the top 5-10 nm of the sample, all the following discussion is therefore focused on the surface properties of perovskite films.

Figure 4.3 presents the I3d and Pb4f XPS spectra collected from CH₃NH₃PbI₃ films fabricated by the $Pb(OAc)_2$, sol-eng and 1-1-1 recipes on ITO/PEDOT:PSS substrates, which had been degraded under constant simulated AM 1.5G solar illumination for 10 h at various oxygen levels. The I3d spectrum consists of $3d_{5/2}$ and $3d_{3/2}$ peaks, with a spin-orbit splitting of $11.5 \,\mathrm{eV}$. The binding energies of $I3d_{5/2}$ peaks for all the pristine films are 619.6 eV, consistent with the reported values from other groups |96|. The $I3d_{5/2}$ peak undergoes a slight shift towards higher binding energies upon degradation, which is monitored in all three recipes, suggesting a similar degradation process regardless of fabrication methods. The Pb4f spectrum consists of $4f_{7/2}$ and $4f_{5/2}$ peaks with a spin-orbit splitting of $4.9\,\mathrm{eV}$. The binding energies of $Pb4f_{7/2}$ peaks for pristine and degraded perovskite films prepared by all three recipes are 138.6 eV, in good agreement with the values reported in literature [96]. Although some groups measured the metallic lead (Pb⁰) peak at a lower binding energy $(136.9 \,\mathrm{eV})$ on freshly prepared samples [48, 160], we observed no Pb^0 for all the pristine films. Small peaks of Pb^0 start to appear on the $Pb(OAc)_2$ perovskite films at an oxygen level of 5% and on the sol-eng films at just 2%. Samples made

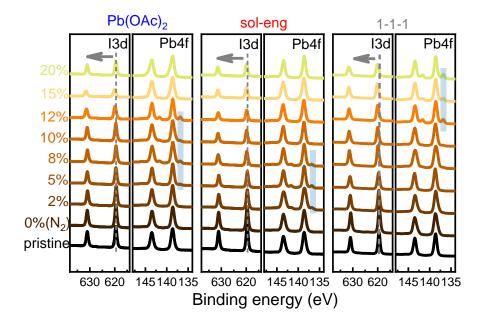


Figure 4.3: I3d and Pb4f spectra collected from X-ray photoemission spectroscopy measurements on pristine and degraded (10 h under various oxygen levels at 1 sun illumination) perovskite films fabricated using the Pb(OAc)₂, sol-eng and 1-1-1 recipes on PEDOT:PSS coated ITO substrates. Adapted with permission from [58]. Copyright 2017, Wiley-VCH.

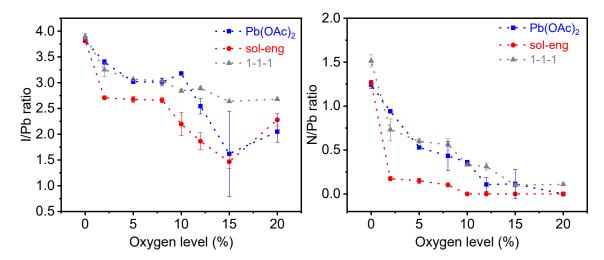


Figure 4.4: I to Pb (left) and N to Pb (right) ratio evolution calculated by X-ray photoemission spectroscopy measurements for degraded films fabricated using the Pb(OAc)₂, sol-eng and 1-1-1 recipes on PEDOT:PSS coated ITO. Adapted with permission from [58]. Copyright 2017, Wiley-VCH.

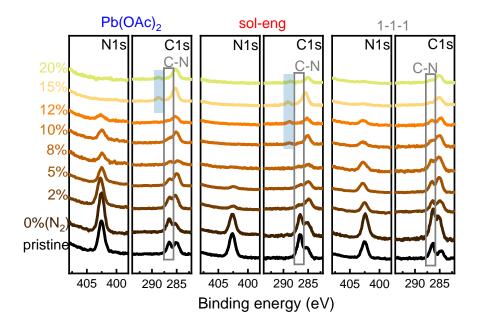


Figure 4.5: N1s and C1s spectra collected from X-ray photoemission spectroscopy measurements on pristine and degraded (10 h under various oxygen levels at 1 sun illumination) perovskite layers fabricated using the Pb(OAc)₂, sol-eng and 1-1-1 recipes on PEDOT:PSS coated ITO substrates. Adapted with permission from [58]. Copyright 2017, Wiley-VCH.

by the 1-1-1 recipe only start to show a small trace of Pb^0 at 12%. The Pb^0 found here is possibly a degradation product in the early phase, as for higher oxygen levels, the previously observed metallic Pb is no longer present for $Pb(OAc)_2$ and sol-eng recipes.

By fitting the Pb4f_{7/2} and I3d_{5/2} peaks we can quantify the atomic percentages of Pb and I in the perovskite and their changes upon degradation. We calculate the I to Pb (I/Pb) ratio by dividing the atomic percentage of I over Pb. The I/Pb ratio for pristine samples is above 3, higher than the expected stoichiometry, due to residual methylammonium iodide on the surface. In the left panel of Figure 4.4, we monitor I/Pb ratio evolution upon degradation and observe a gradual decrease to below 2, suggesting another possible by-product (PbO_x) besides PbI₂. Films prepared by the 1-1-1 recipe show a less steep decline compared to samples prepared by the other two recipes. The I/Pb ratio for the Pb(OAc)₂ films decreases at a comparable speed as the 1-1-1 below 10% O₂ and drops remarkably quickly at higher oxygen levels. The sol-eng recipe gives the least stable films in terms of I/Pb ratio, as it quickly drops to around 2.5 for 2%-5% O₂.

Figure 4.5 presents the N1s and C1s spectra collected from pristine and degraded $CH_3NH_3PbI_3$ films fabricated by the $Pb(OAc)_2$, sol-eng and 1-1-1 recipes. The C1s core level of the pristine perovskite films consists of two types C-bonds. One peak at

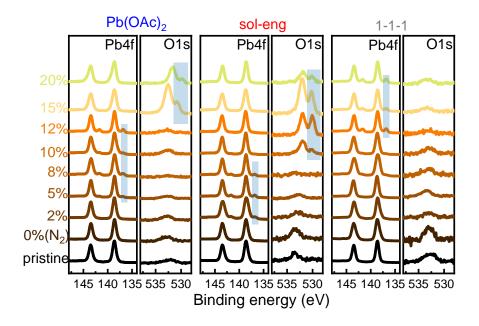


Figure 4.6: Pb4f and O1s spectra collected from X-ray photoemission spectroscopy measurements on pristine and degraded (10 h under various oxygen levels at 1 sun illumination) perovskite layers fabricated using the Pb(OAc)₂, sol-eng and 1-1-1 recipes on PEDOT:PSS coated ITO substrates. Adapted with permission from [58]. Copyright 2017, Wiley-VCH.

a lower binding energy (285.2 eV) represents the C-C and C-H bonds from surface contamination. The other peak, at a higher binding energy of 286.7 eV, is attributed to the C-N bond in the $CH_3NH_3^-$ cation. The corresponding peak in the N1s spectra is observed at 402.6 eV. During the degradation process the N1s core level intensity gradually decreases, coinciding with the disappearance of C-N bond, suggesting the loss of CH_3NH_2 . We calculate the N to Pb (N/Pb) ratio and monitor its evolution under various oxygen levels, as presented in the right panel of Figure 4.4. The N/Pb ratio is above 1 for all the pristine films, due to the excess of methylammonium iodide on the surface mentioned previously. It consistently decreases until both N1s and C-N peaks disappear completely, for oxygen levels above 15% in degraded Pb(OAc)₂ and 1-1-1 films. A sharp drop in the N/Pb ratio for the films prepared by the soleng recipe occurs at significantly lower oxygen levels—as low as 2%. The drastic difference in the compositional evolution suggests a quite different rate of surface degradation among three methods.

Now we turn our attention to the O1s spectra (Figure 4.6). The O1s spectrum of pristine samples only contains one small peak at around 532 eV as a result of fabrication in the dry air environment. A second O1s peak at a lower binding energy (530.2 eV) only appears at high oxygen levels for films prepared by the Pb(OAc)₂ and sol-eng recipes. It is observed together with the increase in the intensity of

the O1s peak around 532 eV. At the same time, N1s and C-N peaks (Figure 4.5) vanish and a new C1s peak at a higher binding energy (288.8 eV) that represents the oxidized carbon species (C=O, O-C=O) is detected. This second O1s peak can be attributed to the appearance of PbO_x during degradation process [161, 162], proposed by the ab initio studies by Zhang and Sit [163]. This result is consistent with the disappearance of Pb⁰ for the Pb(OAc)₂ and sol-eng recipes at higher oxygen levels (Figure 4.3). There is no trace of a second O1s peak for the degraded 1-1-1 films, possibly suggesting a slower break down in the perovskite structure.

Haque and coworkers [53] report that superoxide generated in a light and oxygen environment causes the degradation of perovskite. Upon illumination, electron transfers from photoexcited perovskite to atmospheric oxygen and superoxide (O_2^-) is formed. O_2^- attacks the CH₃NH₃PbI₃, leading to the formation of methylamine (CH₃NH₂), PbI₂, I₂ and H₂O (Reaction 2.26). Zhang and Sit [163] show, based on their ab initio study, that peroxide (OOH⁻) forms through the deprotonation of CH₃NH₃⁻ cations by superoxide, destroying the Pb-I octahedral structure. At the same time Pb-O bonds are formed, together with the possible formation of I⁺, O²⁻, OH⁻ and local PbO structure. Ouyang et al. [164] further propose that the degradation products from the rapid surface oxidation, PbO and Pb(OH)₂, serve as a protective layer to prevent further oxidation of inner perovskite. The other oxidation product, H₂O, causes hydration-based degradation of the inner perovskite at a slower pace.

It is important to note that, even at low oxygen levels (2%-5%), although the surface of perovskite films is degraded as suggested by the loss of N and I (the drop in I/Pb and N/Pb in Figure 4.4), the perovskite under the surface is still intact and partly consists of high-quality perovskite structures, as indicated by the similar Urbach energies to the pristine (Figure 4.2).

We emphasize that the perovskite films degraded in the dark at 20% oxygen for 24 hours (Figure A2 in the Appendix), or aged in N₂ under illumination for 10 h, do not show noticeable differences to the pristine, suggesting that perovskite is stable under these conditions. This observation demonstrates that oxygen alone does not cause degradation on this timescale, and therefore, the observed degradation is a consequence of both illumination and oxygen together, in good agreement with precious reports [55, 57, 165].

So far, all the discussions are focused on the perovskite sample coated on PE-DOT:PSS. We also performed the same series of degradation and XPS studies for the $CH_3NH_3PbI_3$ films on bare ITO substrates. Figure A3, A4 and A5 in the Appendix present the XPS spectra collected from pristine and degraded $CH_3NH_3PbI_3$ films fabricated by the $Pb(OAc)_2$, sol-eng and 1-1-1 recipes on ITO substrates. In comparison to the films on PEDOT:PSS for the same recipe, the decrease in N1s and C-N peaks are much more dramatic, and the appearance of PbO_x starts at lower oxygen levels, indicating that rapid degradation is present without a PEDOT:PSS layer.

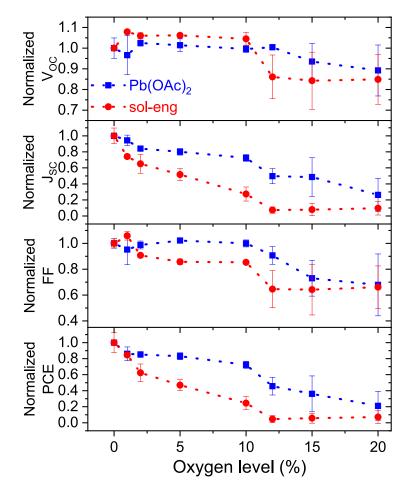


Figure 4.7: Solar cell characteristics of ITO/PEDOT:PSS/CH₃NH₃PbI₃/PC₆₁BM/ BCP/Ag solar cells, with the bare perovskite films degraded for 10 h under various oxygen levels at 1 sun illumination, normalized to the values of the device degraded in N₂ atmosphere. Adapted with permission from [58]. Copyright 2017, Wiley-VCH.

4.4 Photovoltaic Performance of Degraded films

In this section, we focus only on the $Pb(OAc)_2$ and sol-eng recipes. The reason to leave out the 1-1-1 recipe is that it yields very thick perovskite films (approximately 650 nm), nearly three times thicker than the other two (around 275 nm). Since the oxygen-induced degradation starts at the surface, it takes a much longer time for thicker films to undergo the same oxidation process. This explains the stable optical properties, and the slowest degradation rate of the 1-1-1 recipe, compared to other methods. To avoid the possible influence of perovskite film thinness on the rate of degradation, we choose not to further discuss the 1-1-1 in the following sections.

In order to investigate the oxygen induced degradation on the photovoltaic performance, degraded films were transferred to a N₂-filled glovebox immediately following degradation, then completed with $PC_{61}BM/BCP/Ag$ layers into fully functional so-

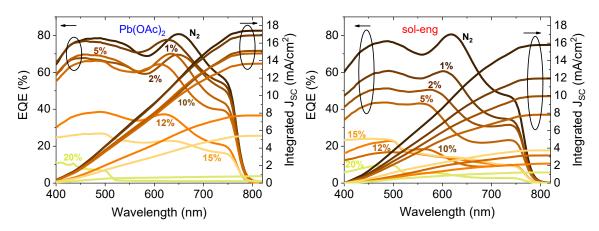


Figure 4.8: EQE spectra of the degraded Pb(OAc)₂ (left) and sol-eng (right) devices, respectively. Adapted with permission from [58]. Copyright 2017, Wiley-VCH.

lar cells, and tested under a solar simulator. The J-V characteristics and EQE spectra of the champion devices are presented in Figure A6 in the Appendix, with the statistics of PV parameters derived from more than 100 devices shown in Table B2. Both recipes result in hysteresis-free solar cells with stable power output (Figure A7 in the Appendix), thanks to the application of $PC_{61}BM$ as ETL [118, 123]. The highest PCEs for solar cells from the Pb(OAc)₂ and sol-eng recipes are 15.5% and 10.5%, respectively. Similar to the previous XPS results (Figure A2 in the Appendix), there is no obvious change in their PV performance compared to the pristine, when the perovskite films were kept in N₂ atmosphere under light for 10 h or stored in 20% O₂ in the dark for 72 h (Figure A8 in the Appendix).

Figure 4.7 shows the evolution of the PV characteristics of the degraded devices, normalised to the performance of a device whose active layer was degraded for 10 h in nitrogen only atmosphere. Below 10 % O₂, V_{OC} and FF remain almost unchanged; the drop in PCE is mostly due to the reduction in J_{SC}. While the Pb(OAc)₂ devices exhibit relatively stable J_{SC} , the sol-eng devices start to show reduced J_{SC} at even 1% O₂. When the oxygen level is above 10%, V_{OC} and FF also start to drop. Soleng solar cells degrade much more significantly than Pb(OAc)₂, with a reduction in all parameters. The PCE of the sol-eng devices reaches nearly zero at 12%, while Pb(OAc)₂ devices at 20% O₂ still exhibit 25% of their initial value. We observe a similar trend in EQE spectra (Figure 4.8) as in the J_{SC}. The drop in EQE of the sol-eng devices is more pronounced than in Pb(OAc)₂, and starts already at very low oxygen levels. The decrease of EQE across the whole visible range suggests the loss in absorption of the active perovskite layers, which is confirmed by the UV-vis results discussed previously.

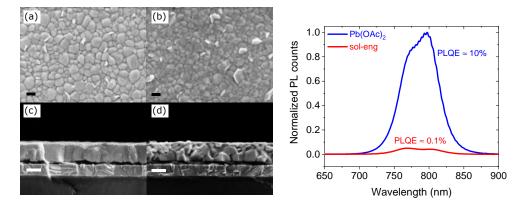


Figure 4.9: Left: Scanning electron microscopy top view and cross-section images of pristine perovskite films prepared by (a,c) Pb(OAc)₂ and (b,d) sol-eng recipes. Scale bars represent 200 nm. The measurements were carried out by Dr. Paul Fassl. Right: Photoluminescence spectra of Pb(OAc)₂ and sol-eng films prepared on spectrosils. The measurements were carried out by Alexandra Bausch. Adapted with permission from [58]. Copyright 2017, Wiley-VCH.

4.5 Discussion—Role of Microstructure

So far, we observe drastic differences in optical properties, surface composition and PV performance of degraded $CH_3NH_3PbI_3$ films prepared by the $Pb(OAc)_2$ and sol-eng methods, which yield perovskite films with highly different microstructures. AFM images (Figure A9 Appendix) of pristine $Pb(OAc)_2$ and sol-eng films coated on ITO/PEDOT:PSS substrates show that both recipes yield pinhole-free and compact films with low root-mean-squared surface roughness $(9.36 \text{ nm for Pb}(OAc)_2)$ and 11.97 nm for sol-eng). The top view SEM images in Figure 4.9 reveal large and regular-shaped grains for $Pb(OAc)_2$, and small and irregular gains for sol-eng samples. The cross-section images show that grains in $Pb(OAc)_2$ film protrude throughout the entire film, which allows the generated charges in perovskite to be transported within single grains. The sol-eng film consists of smaller grains that do not extend vertically through the perovskite film. They are relatively larger at PEDOT:PSS/perovskite interface but smaller at the surface, as a result of the antisolvent treatment [74, 166, 167]. The dramatic difference in PL intensity and PLQE value (Figure 4.9) between two methods also confirms great difference in their microstructure. Sol-eng films show a much reduced PL peak and only 0.1%PLQE, while Pb(OAc)₂ films yields 10 % PLQE. Recent publications [168–173] suggest that the defects and traps at the surface and grain boundaries of perovskite crystals serve as recombination centres that cause the low PL intensity. As a result of smaller grains and a higher defect density [100, 174, 175], sol-eng devices exhibit a lower J_{SC} and FF in comparison to $Pb(OAc)_2$.

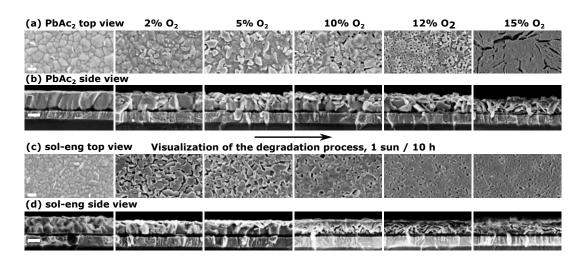


Figure 4.10: Scanning electron microscopy top view and cross-section images of perovskite films prepared by (a,b) Pb(OAc)₂ and (c,d) sol-eng recipes showing changes in morphology for different degradation states from left, degraded in N₂, to right, degraded in 15% O₂, all for 10h at 1 sun illumination. Scale bars represent 200 nm. The measurements were carried out by Dr. Paul Fassl. Adapted with permission from [58]. Copyright 2017, Wiley-VCH.

4.5.1 Evolution of Microstructure in Degradation

The degradation process was monitored by recording SEM images of the surface and cross-section of degraded ITO/PEDOT:PSS/CH₃NH₃PbI₃ films at various oxygen levels. These experiments were performed by Dr. Paul Fassl. Figure 4.10 presents the different evolutions in the morphology of degraded perovskite films prepared by the $Pb(OAc)_2$ and sol-eng recipes. It must be noted that this figure only gives a rough visualization of the degradation process. Since the size of the shown SEM images is $1.8 \,\mu\text{m} \times 0.9 \,\mu\text{m}$, the small sections here are only representative areas for each sample, not the whole film. Films degraded in N_2 do not show any sign of degradation. For $Pb(OAc)_2$ films up to $10\% O_2$, the grain structure is still identifiable, except for the appearance of smaller structures at the surface and the formation of gaps in the cross-section images. For 10% and 12% O₂, small holes appear in the top view images, and they penetrate into and damage the film under the surface, as shown in the cross-section images. For 15% O₂, the morphology alters completely and the layer becomes thinner, suggesting complete conversion to PbI_2 . These results are in good agreement with the PV studies, where degraded $Pb(OAc)_2$ devices still possess 80 % of the initial performance up to 10 % O₂, after which complete deterioration.

For sol-eng films, at only 2% and 5% O₂, the grain structure at the surface is already different. As revealed by the cross-section images, the film is about 40-50 nm thinner, but the grains underneath remain pristine. For 10% and higher oxygen levels, the gaps between grain structures appear, and the entire film is degraded and

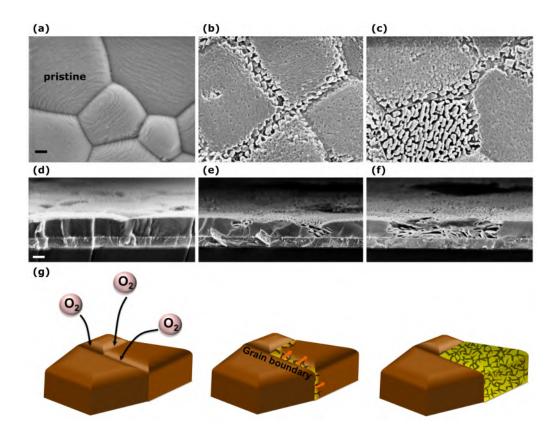


Figure 4.11: Scanning electron microscopy top view and cross-section images of Pb(OAc)₂ on ITO: (a,d) pristine layers, (b,e) initial stages of degradation, and (c,f) later stages of degradation. The scale bars represent 200 nm. (g) Illustration of the progression of oxygen induced degradation from the grain boundaries. The measurements were carried out by Dr. Paul Fassl. The cartoon in (g) was provided through the courtesy of Prof. Yana Vaynzof. Reproduced with permission from [58]. Copyright 2017, Wiley-VCH.

converted to PbI₂. These results are consistent with the previous observation of a severely degraded surface at low oxygen levels (by XPS) but still intact grains in the bulk by UV-vis and PDS measurements. The rapid degradation at the surface of sol-eng samples explains the drop in J_{SC} at very low oxygen levels, while reasonable Voc and FF are maintained by the high-quality perovskite grains in the bulk. The reduction in J_{SC} is related to the loss in absorption and increasing recombination pathways due to the degradation by-products [29, 176, 177].

4.5.2 Role of Grain Boundaries

The observation that smaller grains degrade faster indicates that the oxygen induced degradation is related to the grain size, and possibly the grain boundaries. To investigate this, we monitored the degradation progress of larger perovskite grains. By depositing $Pb(OAc)_2$ films directly on ITO substrates, we increased the grain size to several micrometers, thanks to the non-wettability of the precursor solution on ITO [114]. Figure 4.11 presents top view and cross-section images of pristine and degraded films at 5% oxygen. At the early stages of degradation, the surface of the grains become rougher, and small holes and gaps start to appear at grain boundaries, suggesting that the degradation is initialized at the surface and grain boundaries. Later, the degradation extends to entire grains. However, some grains are still intact, while others are strongly altered. The fact that degraded films still possess some mostly intact grains explains the decent photovoltaic performance in the degraded samples. The whole progress of degradation is illustrated in Figure 4.11g. Previous studies have shown that density of defects is considerable larger at the surface and grain boundaries than in the interior regions of the grains [169, 171, 172, 178–180. Oxygen molecules are able to infiltrate more easily in these areas of higher defect density, generating superoxide that leads to further degradation, which is in agreement with our observation of the fast-degrading sol-eng films, which exhibit small grains, and therefore, more grain boundaries and a high defect density, confirmed by the low PL and PLQE.

4.6 Conclusions

In this chapter, we systematically degraded CH₃NH₃PbI₃ films prepared by various methods, with different microstructures, in controlled oxygen atmospheres under 1 sun illumination. Using UV-vis, PDS, XPS, PV and SEM measurements, we demonstrate that the microstructure of perovskite films has a tremendous effect on the oxygen and light induced degradation. Our study shows that perovskite films with small, irregular grains and high defect density degrade much faster and more severely than films with large uniform grains and better electronic properties. Therefore, the stability of perovskite films, or devices, can be improved in oxygen and light by producing high-quality perovskite films with large uniform grains and reduced defect density. We also reveal that the light and oxygen induced degradation occurs even at low oxygen levels, which emphasizes the importance of avoiding unnecessary light exposure during device fabrication in non-inert atmospheres, in order to reduce the degradation occurring during the production phase.

5 Effects of Large-Scale Compositional and Electronic Inhomogeneities on Device Performance

5.1 Overview

This chapter is based on our recent publication in the journal of ACS Applied Energy Materials with the title 'Large-Scale Compositional and Electronic Inhomogeneities in $CH_3NH_3PbI_3$ Perovskite and Their Effect on Device Performance' [140]. Most graphs are reproduced or adapted with permission from the American Chemistry and Materials Society. The SEM measurements in this chapter were performed by Dr. Paul Fassl.

5.2 Introduction

Since the first application of hybrid organic-inorganic halide perovskite solar cells in 2009 [3], significant work has been devoted to enhance the PCE of perovskite solar cells via optimization of the perovskite composition [6, 8, 65, 155, 181–183] and film morphology [25, 29], and also by improving the adjacent transport layers [184, 185]. Early publications often only reported the maximum PCE of the champion solar cells [3, 5, 186–189]. Recently, researchers have started to pay more attention to the statistics of photovoltaic device performance parameters [29, 74– 76, 190]. A wide distribution in PCE, up to 6-7%, was reported in most studies from 2013 to 2017 (left panel of Figure 5.1). In order to monitor our statistics of device performance, we fabricated and measured a set of 8 samples with 8 pixels on each (in total of 64 solar cells) under identical conditions, with the structure of ITO/PEDOT:PSS/CH₃NH₃PbI₃/PC₆₁BM/BCP/Ag. As presented in the right panel of Figure 5.1, we observed large variations in solar cell performance in the entire batch (black line) of all 64 devices, comparable with previous studies from different groups. We also noticed variation in the pixels from individual devices (8) coloured lines), where both the average and standard deviation of PCE differ from sample to sample.

Although we can attribute the differences in the reported photovoltaic performance distribution from different groups to different fabrication methods and conditions [25, 128, 191–193], the origin of the wide distribution within single studies remains unclear and merely mentioned. In addition to the wide distribution in the performance of the perovskite devices, different IP values (from 5.1 to 6.6 eV)

Chapter 5: Effects of Large-Scale Compositional and Electronic Inhomogeneities on Device Performance

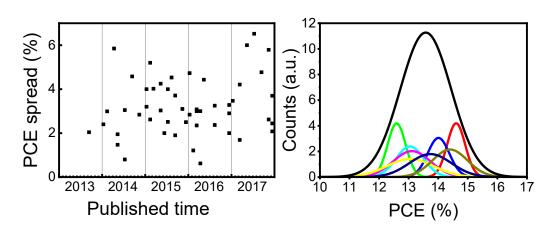


Figure 5.1: Left: PCE spread (i.e., $PCE_{max} - PCE_{min}$) reported in 50 representative studies (for a detailed list of references, see Table B3 in the Appendix) from 2013 to 2017. Right: Variation in PCE of eight identical devices $(8 \times 8 = 64 \text{ pixels})$ from one batch (black line), as well as distributions among eight pixels from each device (coloured lines). Lines displayed are the Gaussian fit to each distribution. The actual histograms of all eight devices are shown in Figure A10a. Adapted with permission from [140]. Copyright 2018, American Chemical Society.

are reported in the literature for the same perovskite $CH_3NH_3PbI_3$ [92–97]. The work from Olthof's group examined the origin of the various reported IP values and consequently the device performances [194]. They performed UPS and XPS measurements on $CH_3NH_3PbI_3$ films fabricated by deliberately varying the precursor composition ratio and with different deposition methods. UPS studies show that IP is significantly altered as a result of the variation in surface composition (i.e., N/Pb ratio) observed by XPS studies. The photovoltaic performance is enhanced when an optimized energy level alignment between the perovskite film and adjacent transport layers is obtained by varying the IP of perovskite, due to a higher accessible opencircuit voltage (V_{oc}) and more efficient charge transport [94, 195–198]. However, single XPS and UPS measurements are limited by the acquisition point size. As a result, IP values of certain spots on the film cannot represent the electronic structure of the whole sample. Therefore, a single-point photoemission spectroscopy experiment is insufficient to provide the full information about the surface compositional and electronic properties of the whole film.

Therefore, we performed XPS and UPS mapping experiments on perovskite films across a $7.2 \text{ mm} \times 7.2 \text{ mm}$ area in the centre of a $12 \text{ mm} \times 12 \text{ mm}$ sample, which covers the active areas of the eight pixels of the final device (see illustration in Figure A10b). Details about the mapping experiments are explained in the section 3.3.8. We observed large surface compositional and electronic variations between identically fabricated samples, and furthermore, a divergent extent of inhomogeneity across single samples. By employing three fabrication recipes for perovskite deposition, we show that these variations are not limited to certain fabrication methods.

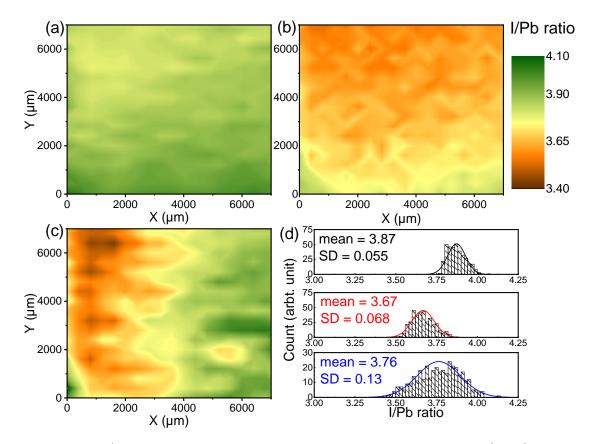


Figure 5.2: I/Pb maps of perovskite samples produced by the Pb(OAc)₂ recipe, where (a) is a uniform film, (b) is quite uniform but with a different average I/Pb ratio, and (c) is a film with a very high inhomogeneity. (d) Histograms of the I/Pb ratio distribution of these three samples (black, sample in panel (a); red, sample in panel (b); blue, sample in panel (c)). Adapted with permission from [140]. Copyright 2018, American Chemical Society.

XPS and UPS mapping experiments performed on the same samples reveal a correlation between compositional and electronic inhomogeneities. Lastly, we completed the films into solar cells and were able to correlate the distribution of the photovoltaic performance to the degree of electronic inhomogeneity.

5.3 XPS and UPS Mapping

5.3.1 I/Pb Maps

To probe the surface compositional properties of the entire sample, we performed two-dimensional XPS mapping experiments across the measured $CH_3NH_3PbI_3$ films, where a serial acquisition of 324 XPS measurements are carried out, with an X-ray

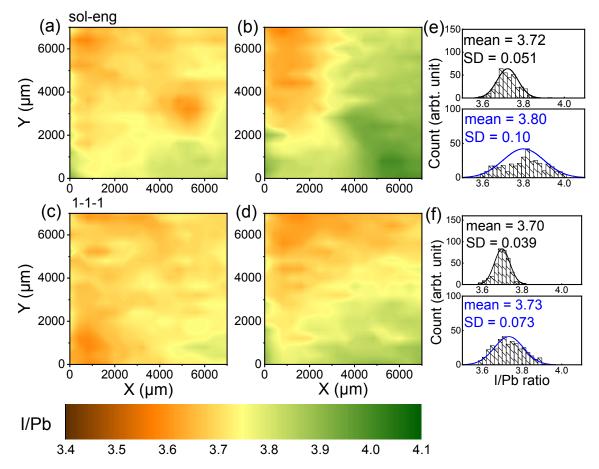


Figure 5.3: I/Pb maps of perovskite films produced by different recipes: (a) uniform and (b) inhomogeneous films produced by the sol-eng recipe; (c) uniform and (d) inhomogeneous films produced by the 1-1-1 recipe. Corresponding histograms of I/Pb ratio distribution of (e) sol-eng and (f) 1-1-1 samples. Adapted with permission from [140]. Copyright 2018, American Chemical Society.

spot size of $400 \,\mu\text{m}$ and step size of $400 \,\mu\text{m}$. At each measurement point, Pb 4f and I 3d spectra are collected, and the I/Pb ratio is then calculated from the ratio of the atomic percentages of iodine and lead. An I/Pb map summarizes the distribution of the I/Pb ratio across the sampled area as a two-dimensional colour contour plot, with x- and y- axes representing the probed position, and the colour difference representing the changes in surface composition.

Figure 5.2 presents I/Pb maps of three $CH_3NH_3PbI_3$ films on ITO/PEDOT:PSS substrates prepared by the $Pb(OAc)_2$ recipe under identical conditions. The first two samples (Figure 5.2a,b) are relatively uniform, but have different average values of I/Pb ratio. In contrast, the third film (Figure 5.2c) exhibits large compositional inhomogeneities across the surface. Histograms of the I/Pb ratio distribution (Figure 5.2d) illustrate the clear differences amongst them. In order to show that the observed compositional inhomogeneities are not limited to certain fabrication methods, I/Pb maps of samples prepared by the sol-eng and 1-1-1 recipes are presented in Figure 5.3. Similar to the Pb(OAc)₂ films, uniform (Figure 5.3a,c) or inhomogeneous samples (Figure 5.3b,d) can be produced by the sol-eng and 1-1-1 recipes under identical conditions as well.

5.3.2 IP maps

In order to monitor the electronic properties of the entire sample, UPS mapping experiments were performed in a similar fashion to the XPS mapping. UPS spectra were collected across the sample with a measurement spot size of around $400 \,\mu$ m. At each probed spot, the work function was acquired by measuring the difference between the cut-off binding energy obtained from the secondary photoemission onset, and the Fermi Level of the spectrometer determined by a standard gold sample (Equation 3.8). The IP is, then obtained by the sum of the work function and the difference between Fermi level and the valence band edge, as determined from the x-intercept of the linear fit of valence band onset (Equation 3.9). An IP map (twodimensional colour contour plot with colour changes representing the IP variations) depicts the distribution of electronic structure across the sampled area.

Figure 5.4 presents IP maps of two representative samples prepared by both the $Pb(OAc)_2$ and sol-eng recipes. Similar to the XPS mapping results, despite the identical fabrication recipe and process, some films are relatively uniform (Figure 5.4a,c), while others exhibit large-scale inhomogeneities in electronic properties (Figure 5.4b,d). It is important to note that the perovskite films prepared by the sol-eng (Figure 5.4c,d) and 1-1-1 recipe (Figure 5.5) tend to be more inhomogeneous than $Pb(OAc)_2$ recipe. A droplet-like feature is observed in the inhomogeneous IP maps of sol-eng (Figure 5.4d) and 1-1-1 films (Figure 5.5b), which can attributed to the effect of antisolvent dropping on the wet film during the fabrication process, suggesting clearly different electronic properties at this dripping spot compared to the rest of the film. Therefore, recipes employing antisolvent quenching methods are more likely to form energetically inhomogeneous samples.

We note a correlation between the surface chemical compositional inhomogeneities

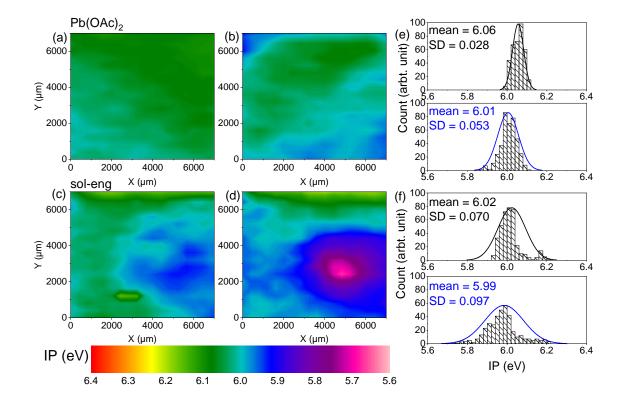


Figure 5.4: IP maps of perovskite films produced by different recipes: (a) uniform and (b) inhomogeneous films produced by the Pb(OAc)₂ recipe; (c) uniform and (d) inhomogeneous films produced by the sol-eng recipe. Corresponding histograms of IP distribution of (e) Pb(OAc)₂ samples and (f) sol-eng samples. Adapted with permission from [140]. Copyright 2018, American Chemical Society.

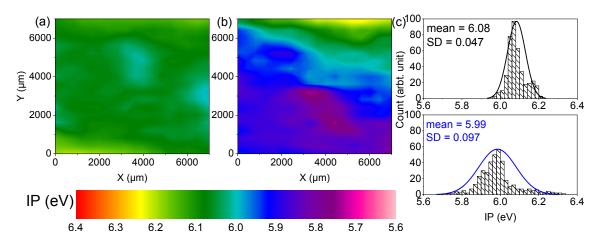


Figure 5.5: IP maps of (a) uniform and (b) inhomogeneous perovskite films prepared by the 1-1-1 recipe. (c) Histograms of the IP distribution of sample (a) and (b). Note the spot in panel (b) resulting from the deposition of the antisolvent during film fabrication. Adapted with permission from [140]. Copyright 2018, American Chemical Society.

and the variation in the electronic properties. Figure A19 in the Appendix shows a linear-like correlation between the standard deviations of IP and I/Pb, where the XPS and UPS mapping were carried out on the same region of the same sample, indicating that samples with a large compositional variation tend to be more inhomogeneous in electronic properties.

5.3.3 Influence of Storage Conditions

To study the effect of storage conditions after fabrication on the surface composition, we stored compositionally inhomogeneous perovskite films in three different conditions (N₂-filled glovebox, dry air filled glovebox and vacuum (10E-10 mbar, same as the photoemission spectroscopy experiments)) for 50 hours, and carried out XPS mapping experiments afterwards. Figure A16 in the Appendix shows that the I/Pb maps of the Pb(OAc)₂ films stored in N₂ and dry air remain unchanged in both I/Pb ratio and its distribution, while the I/Pb value is dramatically reduced over the entire sample for the films kept in vacuum. A similar reduction in I/Pb across the entire sample is observed for the sol-eng and 1-1-1 perovskite films stored in vacuum as well (Figure A17, Appendix). This decrease in I/Pb can be attributed to the loss of methylammonium iodine from perovskite surface as a result of long-time vacuum exposure [199]. In order to minimize the effect of exposure to vacuum on the mapping results, we kept the XPS and UPS mapping measurement time of a single sample to below 1.5 h.

Chapter 5: Effects of Large-Scale Compositional and Electronic Inhomogeneities on Device Performance

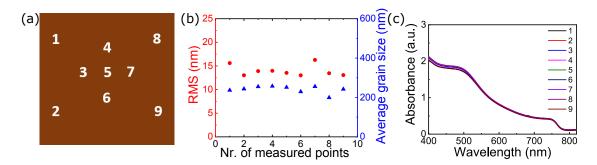


Figure 5.6: (a) Illustration of nine different spots on the sample where AFM, SEM and UV-vis measurements are carried out. (b) Roughness and average grain size of an inhomogeneous perovskite film produced by the Pb(OAc)₂ recipe measured at various spots, calculated from AFM measurements. (c) Absorbance spectra of an inhomogeneous perovskite film produced by the Pb(OAc)₂ recipe measured at various spots. Adapted with permission from [140]. Copyright 2018, American Chemical Society.

5.3.4 Possible Causes for the Surface Inhomogeneities

Microstructure, Absorption or Phase Impurities?

In order to investigate the microstructural variation across the perovskite samples, we carried out AFM and SEM measurements on nine different spots across the perovskite film (Figure 5.6a). As presented in Figure 5.6b, for perovskite films prepared by the $Pb(OAc)_2$ recipe on ITO/PEDOT:PSS substrates, the root-meansquare roughness and average grain size (obtained from AFM measurements) show no obvious difference at different spots across the film, in good agreement with the unchanging grain structures in SEM images (Figure A11, Appendix). Figures A13 and A14 in the Appendix further demonstrate that the microstructure of perovskite films prepared by the sol-eng and 1-1-1 recipes also remain unchanged across the sample. Therefore, the compositional inhomogeneities at the film surface are not related to variations in the microstructure of perovskite samples. We also measured the UV-vis absorption of the perovskite at different points. As presented in Figure 5.6c and A12, the absorption remains unaltered across the film for all three recipes, suggesting that the observed inhomogeneities are unrelated to the changes in absorption. It is important to note that the surface compositional inhomogeneities are not the result of phase impurities, indicated by the X-ray diffraction spectrum of $Pb(OAc)_2$ film (Figure A15, Appendix), as no residual PbI_2 or lead acetate is detected.

Fabrication Conditions

The impact of solvent atmosphere during perovskite fabrication on the crystal formation and film morphology has already been discussed in literature [186, 191, 200], which possibly affects the chemical composition at the surface of perovskite films. The perovskite layers were solution-deposited in a spin-coater with a closed lid located in a dry air filled glovebox. It is likely that the surface compositional

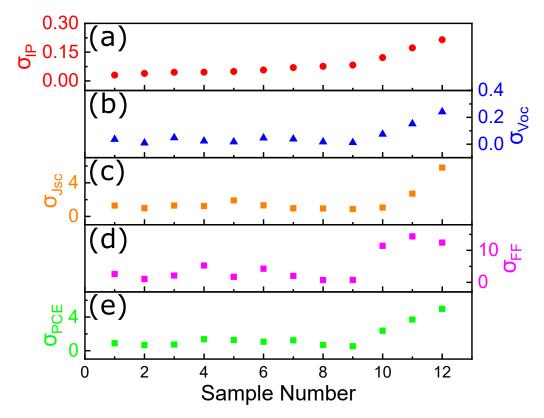


Figure 5.7: (a) Standard deviation of IP ($\sigma_{\rm IP}$) derived from UPS mapping experiments of 12 perovskite films, which afterwards were completed into solar cells and measured under AM1.5 G simulated sunlight (100 mW/cm²). Standard deviation of V_{OC} (b), J_{SC} (c), FF (d), and PCE (e) extracted from J-V measurements of the corresponding solar cell devices. Reproduced with permission from [140]. Copyright 2018, American Chemical Society.

inhomogeneities are related to the fact that films cannot be prepared in identical environmental conditions. For example, the first sample is prepared in the atmosphere with the lowest solvent vapor concentration, and samples coated later face a higher solvent vapor concentration atmosphere than the previous ones, since residual solvent from the precursor solutions cannot easily escape the spin-coater and continuously form a solvent bath inside. The variations observed across the sample could also be attributed to uneven annealing of the perovskite films, possibly induced by temperature gradient across the heating plates, or by the fact that some region of the sample touches the heating plate earlier due to manual handling errors.

5.4 Effect on Device Performance

To investigate the influence of electronic inhomogeneities on the photovoltaic performance, perovskite films prepared by the $Pb(OAc)_2$ recipe were completed with $PC_{61}BM/BCP/Ag$ into fully functional solar cells following the UPS mapping experiments and tested under a solar simulator. Since the observed inhomogeneities are not limited to certain fabrication methods, this study was carried out using only the Pb(OAc)₂ recipe. Average device parameters and standard deviations of all twelve devices are listed in Table B4 in the Appendix, with the J-V curves of their champion pixels shown in Figure A18 in the Appendix. We calculated standard deviations of V_{OC}, J_{SC}, FF and PCE for each sample to quantify their degree of inhomogeneity in photovoltaic performance. As presented in Figure 5.7, there is a clear correlation between the standard deviations of IP and device performance parameters. Samples with a narrow distribution in ionization potential (Sample nr. 1-9) exhibit a small performance variation, while the electronically inhomogeneous perovskite films (sample nr. 10-12) result in devices with widespread PV parameters.

It has been reported that the IP of the perovskite film influences the built-in potential in the final photovoltaic device and consequently the accessible V_{OC} [194]. As a result of the unchanged absorption across the perovskite film, the band gap remains the same; however, the observed variation in the IP suggests corresponding changes in the electron affinity (i.e., the conduction band position of perovskite) across the film. This in turn affects the energetic alignment at the perovskite/PC₆₁BM interface, and consequently the electron extraction efficiency and V_{OC} [94, 195–198]. Some publications have shown that the IP of perovskite layers also affects the hole blocking and charge carrier recombination efficiency at the perovskite/PC₆₁BM interface [201, 202]. The differences in the charge carrier extraction and blocking efficiency explain the observed changes in the J_{SC} and FF, despite the uniform microstructure and unchanged absorption across the sample. The wide spread in PCE observed in our study and by other groups (Figure 5.1 left panel) is hence the combined result of these factors.

5.5 Conclusions

In this chapter, large-scale chemical compositional and electronic inhomogeneities of $CH_3NH_3PbI_3$ films are observed with XPS and UPS mapping experiments. By analysing perovskite films prepared by different fabrication methods, we demonstrate that these inhomogeneities are not limited to certain recipes, and are therefore commonly observed by other research groups as well. Our study shows that the extent of the variation in PV parameters across the entire sample are related to the extent of the observed inhomogeneities, which explains the wide spread of device performance reported in the literature. Therefore, in order to obtain devices with reliable and reproducible performance, more attention should be devoted to improving the perovskite fabrication process, in order to obtain an active layer with a uniform surface composition and electronic properties. Additionally, our results highlight the importance of studying the surface properties of perovskite layers, instead of simply focusing on bulk properties such as microstructure and absorption.

6 The Relationship between Antisolvent Type and Dripping Speed, Determining the Device Performance

6.1 Overview

This project was performed in collaboration with Dr. Alex Taylor, who inspired the project idea and contributed to the samples fabrication and device characterisation (J-V measurements). The SEM and optical microscope measurements in this chapter were carried out by Dr. Fabian Paulus and Maximilian Litterst, respectively.

6.2 Introduction

Solution-processed perovskite has drawn great attention in the photovoltaic community, as it ensures highly efficient solar cell at a low manufacturing cost. One of the most employed fabrication methods uses an antisolvent dripping technique during spin coating the perovskite precursor solution on the substrate. It results in high-quality perovskite films that yield reproducible solar cells with efficiencies over 22% [8, 65]. Since the introduction of this technique in 2014 [74, 85], the choice of antisolvents in early times were always among toluene [74], diethyl ether [75] and chlorobenzene [8, 87]. Interestingly, these three solvents produce high-quality perovskite films with high photovoltaic efficiencies, despite their different physiochemical properties (such as boiling point and polarity (Table B5 in the Appendix)). Recently, three promising new candidates, namely ethyl acetate [88], anisole [90] and trifluorotoluene [89], were reported to yield champion efficiencies of 19.4\%, 19.8% and 20.3\%, respectively.

Although diethyl ether has been reported to result in a rapid crystallization of high-quality pinhole-free $CH_3NH_3PbI_3$ perovskite film [75, 203], it does not seem to work well for the $(FAPbI_3)_{0.85}(MAPbBr_3)_{0.15}$ perovskite. Bu et al. reported a poorly functioning device using diethyl ether, due to its low boiling point [88]. Paek et al. also showed that pinholes are present on the surface of films treated with diethyl ether [89]. They proposed that only the DMF/DMSO miscible solvents with a high boiling point (such as trifluorotoluene, toluene and chlorobenzene) assure a smooth and complete perovskite film, which yields much enhanced device performance in comparison to the immiscible solvents (e.g. xylene and diethyl ether).

It is interesting to notice that there are no clear/standard requirements for a 'good' antisolvent, and a 'poor' antisolvent for some compositions might work great on other

occasions. Therefore, we elucidate this issue with a systematic study comparing a large number of solvents in the following sections. In addition, parameters of the antisolvent application such as the dripping time [204] and the antisolvent volume [90] have been well studied, whereas the influence of the antisolvent dripping speed on the perovskite formation and its device performance has been merely discussed. Hence, we also compare two different dripping speeds (i.e., fast and slow) of a series of solvents and reveal the relationship between the antisolvent and its dripping speed, and their effects on the device performance.

6.3 The Relationship between Antisolvent type and Dripping Speed

We employed thirteen different solvents: ethanol (EtOH), 2-propanol (IPA), 1butanol (BuOH), ethyl acetate (EA), chloroform (CF), chlorobenzene (CB), butyl acetate (BA), 1,2-dichlorobenzene (DCB), anisole (Ani), trifluorotoluene (TFT), diethyl ether (DE), m-xylene (XYL) and toluene (Tolu). Their chemical structures are presented in Figure A20 with the physiochemical properties listed in Table B5 in the Appendix. These 13 solvents not only contain the widely used popular antisolvents (e.g. CB and Tolu), but also include some traditionally considered 'poor' solvents (e.g. EtOH, IPA and DE), in addition to some solvents which were never reported as the antisolvent (e.g. CF).

6.3.1 Device Performance

In order to investigate the relationship between the antisolvent and its dripping speed, we fabricated $Cs_{0.05}(MA_{0.17}FA_{0.83})_{0.95}Pb(I_{0.9}Br_{0.1})_3$ (CsMAFA) perovskite solar cells with the device architecture of ITO/PTAA/PFN-P₂/CsMAFA/PC₆₁BM/BCP/Ag and compared the resulting PV performance of two dripping speeds (fast and slow) using 13 different antisolvents. Their J-V characteristics and statistics are presented in Figure 6.1. The other parameters, such as dripping time and volume of the applied antisolvent were kept the same. (Fabrication details are described in Chapter 3.2.3.)

It is clear that the antisolvent dripping speed influences the perovskite device performance to different extents, depending on which antisolvent was employed during the fabrication of perovskite film. Therefore, we categorise these 13 antisolvents into three types based on the different device performances between fast and slow antisolvent application.

Methanol (MeOH) (type 0) does not produce a photoactive black perovskite film due to its high solubility of organic precursors [85]. The first type (type I) contains three antisolvents: EtOH, IPA and BuOH. They belong to the alcohol group, which has a hydroxyl functional group (-OH) bound to a carbon atom. Devices whose perovskite films were prepared with a fast dripping of the type I antisolvents obtain a higher V_{OC} , J_{SC} , FF and PCE with narrower distributions, in comparison to the

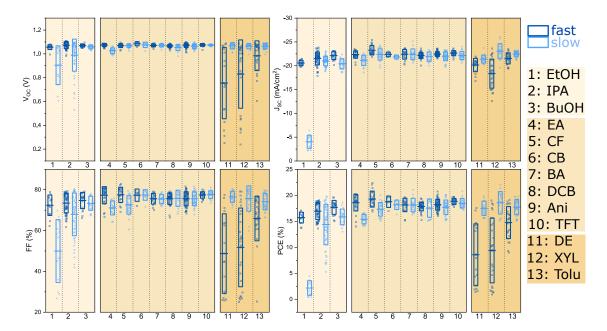


Figure 6.1: Solar cell characteristics of ITO/PTAA/PFN-P₂/CsMAFA/PC₆₁BM/BCP/Ag solar cells, with the perovskite films fabricated using 13 different antisolvents with fast (dark blue) and slow (light blue) dripping speeds. The used antisolvents are EtOH, IPA, BuOH, EA, CF, CB, BA, DCB, Ani, TFT, DE, XYL and Tolu, from 1 to 13.

devices fabricated by a slow antisolvent dripping. The difference in the PV performance between the fast and slow dripping speeds varies from extremely different to relatively comparable for EtOH, IPA and BuOH, respectively. This can be attributed to the continuous decrease in the solubility of organic precursors (i.e., MAI and FAI), which is further elaborated in Section 6.4. The solar cell fabricated by a slow EtOH dripping has a PCE of less than 5% with a low J_{SC} below 7 mA/cm², suggesting the formation of a poor-quality perovskite absorbing layer. This is confirmed by the non-stoichiometric chemical composition by its XPS measurements (Section 6.3.2) and also by its SEM images (Section 6.3.3).

For the type II antisolvents, namely EA, CF, CB, BA, DCB, Ani and TFT, the PV performance is not affected by the antisolvent application speed with a PCE around 20%. Therefore, the type II solvents are promising candidates, as it does not require a specific dripping speed to ensure the formation of high-quality perovskite films and well-functioning devices. However, there is a slight variation among device performances using different antisolvents. The use of EA and CF yields devices with slightly different PV parameters between fast and slow dripping speeds, while devices employing TFT as the antisolvent obtain comparable and reproducible PV performance, regardless of dripping speed. Further discussion is elaborated in Section 6.4.

The type III category consists of DE, XYL and Tolu. The devices using this type antisolvents behave the opposite to the type I: perovskite solar cells fabricated by a slow antisolvent dripping are more efficient with improved reproducibility, in comparison to these using a fast application. Possible explanations are discussed in Section 6.4.

6.3.2 Surface Chemical Composition

To make sure that the perovskite layer contains the correct stoichiometry, we carried out XPS measurements on CsMAFA films fabricated using 13 antisolvents with fast and slow dripping speeds coated on ITO/PTAA/PFN-P₂ substrates. Five acquisitions were taken at five different spots of each sample. As XPS is a surface sensitive technique that only probes the top 5-10 nm of the sample, all the following discussion is therefore focused on the surface properties of perovskite films.

By fitting the Pb4f_{7/2} and I3d_{5/2} peaks from their XPS spectra we can quantify the atomic percentages of Pb and I in the perovskite. We calculate the I to Pb (I/Pb) ratio by dividing the atomic percentage of I over Pb. The Br to Pb (Br/Pb), Cs to Pb (Cs/Pb), FA to Pb (FA/Pb) and MA to Pb (MA/Pb) ratios are obtained in a similar way. The N1s spectrum of the CsMAFA perovskite contains two peaks at around 400.6 eV and 402.9 eV representing the FA and MA, respectively (Figure A21a in the Appendix). The peak positions are consistent with the reported values from other groups [205]. The atomic percentages of FA and MA are quantified by fitting the corresponding N1s peaks. The theoretical values of the I/Pb, Br/Pb, Cs/Pb, FA/Pb and MA/Pb ratios based on the precursor concentration are listed in the Table 6.1. Note that FA contains two nitrogen atoms; therefore, the theoretical

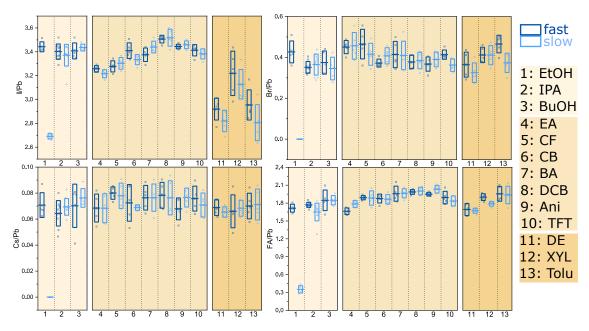


Figure 6.2: The I/Pb, Br/Pb, Cs/Pb and FA/Pb ratios of CsMAFA perovskite films fabricated using 13 antisolvents with fast (dark blue) and slow (light blue) dripping speeds on ITO/PTAA/PFN-P₂ substrates. The used antisolvents are EtOH, IPA, BuOH, EA, CF, CB, BA, DCB, Ani, TFT, DE, XYL and Tolu, from 1 to 13.

ratio of FA/Pb is 1.58, not 0.79. As shown in Figure A21a (Appendix), the N1s peak assigned to MA is low and noisy, which makes it difficult to perform a neat fitting with small errors. As a result, there is a large variation in the MA/Pb ratio among different acquisition spots on every single sample (Figure A21b, Appendix).

Figure 6.2 presents the I/Pb, Br/Pb, Cs/Pb and FA/Pb ratios of perovskite layers fabricated using 13 different antisolvents with fast and slow dripping speeds.

Table 6.1: Theoretical values of the I/Pb, Br/Pb, Cs/Pb, FA/Pb and MA/Pb ratios.

ratio(theo)	I/Pb	$\mathrm{Br/Pb}$	Cs/Pb	FA/Pb	MA/Pb
	2.7	0.3	0.05	1.58	0.16

For the perovskite samples prepared by these antisolvents, the calculated I/Pb and FA/Pb ratios are approximately 30 % higher than the theoretical values, except for the film fabricated by a EtOH slow dripping. These high ratios indicate the presence of residual FAI on the surface, proposed by Philippe et al. [205]. In addition, there is a slight enrichment in the Br/Pb and Cs/Pb ratios for these samples, which is also the consequence of the excess FAI on the sample surface [205].

The perovskite film fabricated using a slow EtOH dripping contains a much reduced I content, together with no traces of Br nor Cs. Moreover, it exhibits nearly zero percentage of MA and a much lower FA amount. These findings indicate the

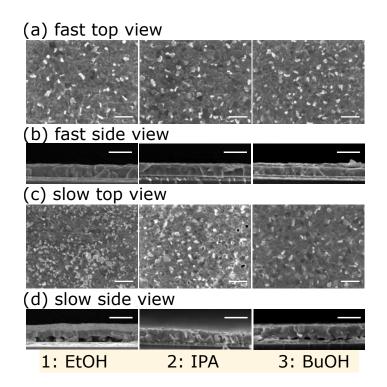


Figure 6.3: Scanning electron microscopy top view and cross-section images of the CsMAFA perovskite films fabricated by dripping the antisolvent fast (a,b) and slow (c,d). The antisolvents used are EtOH, IPA and BuOH, from left to right. Scale bars represent $1 \,\mu$ m. The measurements were carried out by Dr. Fabian Paulus.

presence of PbI₂, rather than the CsMAFA perovskite. Consequently, the devices prepared using a slow EtOH dripping exhibit a PCE below 5% with much reduced J_{SC} and FF.

6.3.3 Microstructure

In order to monitor the influence of antisolvent type and dripping speed on the perovskite microstructure, we recorded SEM top view and cross-section images of CsMAFA perovskite films fabricated using 13 antisolvents with fast and slow dripping speeds on ITO/PTAA/PFN-P₂ substrates. These experiments were performed by Dr. Fabian Paulus and are presented in Figure 6.3, 6.4 and 6.5.

Figure 6.3 demonstrates the clear difference in the microstructure of the perovskite films on the top surface and also at the PTAA/PFN-P₂/perovskite interface prepared using the type I antisolvents with different dripping speeds. The white content in the top view images are PbI₂ phases [206]. It is clear that the perovskite film prepared by a EtOH slow dripping exhibits many more PbI₂ phases than the one produced with a fast dripping. Moreover, the EtOH slow sample exhibits a different grain structure on the surface and in the bulk. As revealed by its cross-section

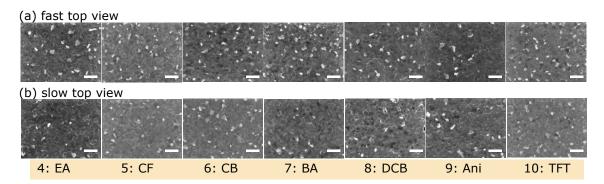


Figure 6.4: Scanning electron microscopy top view images of the CsMAFA perovskite films fabricated by dripping the antisolvent fast (a) and slow (b). The antisolvents used are EA, CF, CB, BA, DCB, Ani and TFT, from left to right. Scale bars represent $1\,\mu$ m. The measurements were carried out by Dr. Fabian Paulus.

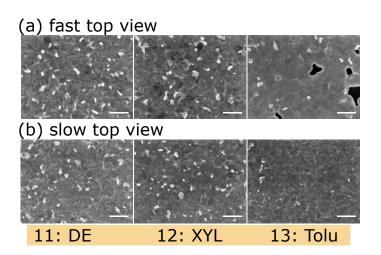


Figure 6.5: Scanning electron microscopy top view images of the CsMAFA perovskite films fabricated by dripping the antisolvent fast (a) and slow (b). The antisolvents used are DE, XYL and Tolu, from left to right. Scale bars represent $1 \,\mu$ m. The measurements were carried out by Dr. Fabian Paulus.

image, 'cave'-like structures are present at the PTAA/PFN-P₂/perovskite interface, suggesting an incomplete film coverage at this interface. We observe a similar trend when using IPA and BuOH as the antisolvent. The perovskite films prepared by fast dripping are dense and complete, while the slow application results in a slight increase of PbI₂ phases and incomplete film coverage with either holes penetrating through the whole layer (IPA) or caves present at the PTAA/PFN-P₂/perovskite interface (BuOH).

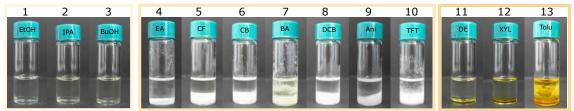
The top view and cross-section images of the perovskite films fabricated using type II antisolvents with different dripping speeds are presented in Figure 6.4 and A22 (Appendix). All the perovskite films prepared using type II antisolvents are pinhole-free and compact, with comparable grain sizes and less PbI_2 phases. The dripping speed of type II antisolvents does not affect the perovskite microstructure, nor the amount of the PbI_2 phases.

Figure 6.5 and A23 (Appendix) show the top view and cross-section images of the perovskite films fabricated using type III antisolvents with different dripping speeds. Many fewer PbI₂ phases are present in comparison to the films fabricated employing type I antisolvents. Note that it is difficult to form a complete perovskite film using a fast antisolvent dripping, as hinted at by the SEM image of the Tolu fast sample in Figure 6.5 and the optical microscope images in Figure A24a in the Appendix. When the type III antisolvent is applied fast, the photoactive black perovskite forms only in the middle, close to where the antisolvent droplet hits the sample, while regions closer to the edges exhibit a completely different microstructure. Perovskite films prepared by a slow dripping of the type III antisolvents are pinhole-free and compact.

6.4 Discussion

So far, we have reported that there are three types of antisolvents, each resulting in different device performance with different dripping speeds. For the first type, only fast application of the antisolvent ensures a pinhole-free and compact perovskite film. As a result, the devices prepared by fast dripping exhibit improved PV performance with narrower distributions than those with a slow dripping. For the type II antisolvents, the dripping speed is not as crucial as for the first type, as PV performance is not significantly altered when the antisolvents are applied fast or slow. The type III antisolvents work exactly opposite to the type I. Perovskite devices prepared using a fast dripping perform poorly with a large PV performance variation, whereas the solar cells prepared by a slow dripping exhibit comparable PV parameters to the type II antisolvents. This is likely the reason why some of the type III antisolvents were regarded as 'poor' choices for the fabrication of perovskite films in the literature (e.g. DE and XYL) [88, 89], because the researchers might have applied the antisolvent too rapidly.

In order to investigate the reasons why the dripping speed has quite different effects on device performance for these three types of antisolvents, we carried out a solubility test of MAI in all antisolvents (Figure 6.6). The MAI solubility test



Antisolvents added to MAI (2 M in DMF/DMSO) solution at a volume ratio of 6:1

Figure 6.6: Pictures of different solution mixtures, where 13 antisolvents were added to the MAI solution (2M in DMF/DMSO) at a volume ratio of 6:1. The used antisolvents are EtOH, IPA, BuOH, EA, CF, CB, BA, DCB, Ani, TFT, DE, XYL and Tolu, from 1 to 13. 1-3: Clear and transparent solution. 4-10: White precipitants at the bottom of the vials. 11-13: Yellow solution with two different phases.

is performed by adding different antisolvents to a 2 M MAI solution (dissolved in DMF/DMSO (v:v, 4:1)) at a volume ratio of 6:1. This ratio is very close to the volume ratio of the antisolvent to the perovskite precursor solution applied on each sample during spin coating process.

Furthermore, using the SEM top view images we determined the percentage of PbI₂ phases to the whole sampled area, and compared the amount of PbI₂ content in samples prepared using all 13 antisolvents with different dripping speeds (Figure 6.7). Details and the SEM images used for this calculation are further elaborated in the Appendix C. It is important to point out that the SEM images employed for calculating the PbI₂ amount are only representative areas for each sample, not the whole film, since the size of the shown SEM images in Figure C1, C2 and C3 is $12 \,\mu\text{m} \times 9.6 \,\mu\text{m}$.

Type I

MAI dissolves easily in the type I antisolvents. The solubility in EtOH, IPA and BuOH decreases exponentially from EtOH to BuOH, as presented in Figure A25 in the Appendix. The lowest volumes of solution required to fully dissolve 100 mg MAI powder in methanol (MeOH), EtOH, IPA and BuOH are listed in Table B6 (Appendix). FAI is even more soluble in the type I antisolvents than MAI. The solubility test of FAI is shown in Figure A26 and Table B7 in the Appendix. These results allow us to understand the performance trend for the type I antisolvents.

The organic precursors (i.e., MAI and FAI) are extremely soluble in MeOH. When it is used as the antisolvent (regardless of dripping speed), the MAI and FAI are washed away and the stoichiometry of the perovskite is significantly altered, and consequently a yellow (PbI₂) film is formed (Figure A27, Appendix). This is consistent with previous reports, as Xiao et al. also observed a yellow film using MeOH as the antisolvent [85].

When EtOH, the second most soluble solution for MAI and FAI, is applied slowly to the wet film in the spin coater, it stays in contact longer with the wet film. As a

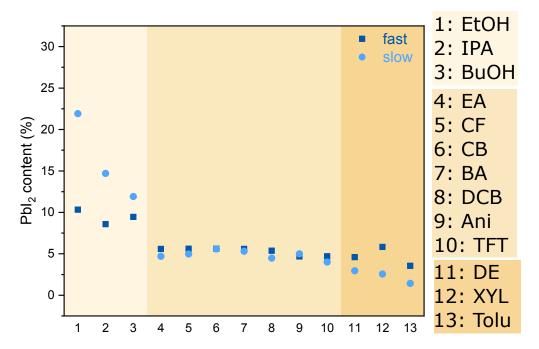


Figure 6.7: The PbI₂ content of the CsMAFA perovskite films fabricated using 13 antisolvents with fast (dark blue) and slow (light blue) dripping speeds on ITO/PTAA/PFN-P₂ substrates. The used antisolvents are EtOH, IPA, BuOH, EA, CF, CB, BA, DCB, Ani, TFT, DE, XYL and Tolu, from 1 to 13. The PbI₂ content is obtained by calculating the ratio of the white PbI₂ phases to the sampled area from the corresponding SEM images. Details are explained in the Appendix C.

consequence, it washes away more organic precursors, similar to the situation with MeOH. This explains the large PbI₂ content in the SEM images, the extremely low I/Pb ratio in the XPS results, and the poor PV performance (PCE below 5%). Fast dripping of EtOH alleviates this problem by reducing the time in contact with the wet film, as indicated by the reduction of the PbI₂ amount from 22% to 10%, but a best PCE above 18% is still never achieved.

IPA and BuOH exhibit a relatively lower solubility of MAI and FAI. Therefore, the differences in the PbI_2 content and device performance are still present between the fast and slow dripping speeds, but are much reduced in comparison to the EtOH case.

Perovskite films fabricated using type I antisolvents exhibit the largest PbI_2 content on the surface in comparison to the other two types, regardless of dripping speed, as a result of the soluble nature of MAI and FAI in the type I antisolvents. It is also important to note that when the type I antisolvents are dripped fast, the samples contain less PbI_2 content and have much improved device performance with narrower distributions.

Type II

The organic precursors are moderately soluble in the type II antisolvents when compared to type I. Unfortunately, the test with FAI with the same molar concentration and antisolvent volume as the MAI test is inconclusive, as shown in Figure A28 in the Appendix. For the MAI test, the white precipitants at the bottom of vials in Figure 6.6 suggest that the type II antisolvents extract the DMF/DMSO from the precursor solution, rather than dissolving the organic precursors, which helps to maintain the correct stoichiometry and facilitate the rapid formation of a high-quality perovskite film. As a result, the period that the antisolvent is in contact with the wet film on the spin coater is not critical. This explains the excellent device performances, regardless of dripping speed. It is also consistent with the much reduced PbI₂ content in comparison to the type I.

However, there is a slight variation among device performances using different antisolvents. For example, the devices using EA as the antisolvent obtain noticeably different PV parameters between fast and slow dripping speeds. This can be attributed to the fact that EA has the highest solubility for the organic precursors among all the type II antisolvents, as indicated by the smallest amount of the white precipitants in the EA mixture in Figure 6.6. In order to further investigate this issue, we fabricated perovskite devices using EA at six gradually changing dripping speeds. As presented in Figure A29 in the Appendix, the evolution in the V_{OC} , J_{SC} , FF and PCE from fast to slow is quite abrupt, rather than gradual. It seems that there is a critical point in the dripping speed; dripping slower than this will result in devices with reduced PV performances. Therefore, the difference in the device performance between the fast and slow dripping speeds for EA is probably due to the fact that our 'slow' speed is below this critical value. In contrast, the addition of TFT to the MAI solution leads to the formation of the largest amount of the white precipitants, which is consistent with the narrowest distribution in the device performance, regardless of dripping speed. Therefore, using an antisolvent that possesses a low solubility of MAI (such as TFT) will further expand the processing window of the dripping speed and increase the device reproducibility.

Type III

For the type III antisolvents, we observe a colour change and phase separation as shown in Figure 6.6 and A28 (Appendix). To further explore, we performed UV-vis measurements on these two phases (Figure A30 in the Appendix). An absorption onset in the bottom phase starting at around 500 nm is observed for all three solvents. This onset resembles the iodine (I₂) absorption spectrum [207], which suggests that a considerable amount of iodine is extracted by the antisolvent. The top phase of the Tolu mixture shows a similar onset as in the iodine, while the top phases of DE and XYL mixtures exhibit the two peaks, which are observed in the I₃⁻ absorption spectrum [207] as well, with I₃⁻ formed by combining iodine and iodide (I⁻).

Here we propose that the relatively low PbI_2 amount in the slow samples is the result of longer contact duration with the antisolvent during spin coating, as the antisolvent extracts a significant amount of iodine, resulting in perovskite films with less PbI_2 content. This is consistent with the slightly enrichment in the I/Pb ratio in Figure 6.2 for the samples prepared using a fast dripping.

The poor coverage of perovskite films using a fast dripping of type III antisolvents is probably due to the phase separation of the organic precursor solution and the antisolvent, as shown in Figure 6.6 and A28 (Appendix). This phase separation suggests that the type III antisolvents are poor at extracting the DMF/DMSO from the organic precursor solution, which further impedes the formation of a highquality perovskite film. When the antisolvent is applied fast to the spinning sample, it hits the centre area with a considerable amount of vertical force and is able to penetrate into the wet film and initiate crystallization of the perovskite. However, it then leaves the sample rapidly, without enough contact time with the outside regions to effectively extract the DMF/DMSO, yielding a film with poor uniformity. Consequently, a photoactive black perovskite is only formed in the middle of the substrate. In contrast to the type I antisolvents, where slow application prolongs contact and leads to the extraction of the organic compounds, for the type III antisolvents this leads to longer contact between the immiscible DMF:DMSO:perovskite and antisolvent phases, allowing for more complete extraction of the DMF/DMSO. This is supported by the formation of a uniform, compact perovskite film for the slow antisolvent application, and an inhomogeneous, pinhole-possessing film for fast application.

6.5 Conclusions

In this chapter, we systematically investigate the relationship between antisolvent type and dripping speed, and their effects on PV performance. Using PV, XPS, SEM and solubility measurements, we demonstrate that the organic precursors' solubility in the antisolvent has a tremendous effect on the perovskite formation, microstructure, and consequently, the final device performance. Our study shows that the antisolvents which exhibit a low solubility of organic precursors have an extended processing window, which is not limited to certain dripping speed. Furthermore, antisolvents which phase separate (type III) can lead to poor film formation with inappropriate dripping speed. Therefore, the solubility test serves as a good indicator of which application speed is required to obtain a good device for a new antisolvent. Although it is possible to obtain high-functioning devices with appropriate dripping speed for every antisolvent, the use of certain types of antisolvents eases the fabrication process and improves device reproducibility. We demonstrate that the devices fabricated using TFT as the antisolvent exhibit the best PV performance with the highest reproducibility.

7 Conclusions and Future Work

This thesis focuses on investigating two of the biggest challenges of perovskite solar cells: stability and reproducibility. Chapter 4 studies the role of microstructure on the oxygen and light induced degradation of perovskite films. Chapter 5 and 6 investigate the reproducibility issue from two different aspects: large-scale inhomogeneities of perovskite films, and how device performance is influenced by antisolvent type and dripping speed during film fabrication.

In Chapter 4, we investigate the degradation of $CH_3NH_3PbI_3$ films in controlled oxygen atmospheres under continuous illumination, using UV-vis, PDS, XPS, PV and SEM measurements. We reveal the dominant role of the microstructure on the oxygen and light induced degradation: perovskite films with small, irregular grains and high defect density degrade much faster and more severely than films with large uniform grains and better electronic properties. Therefore, by improving the perovskite film quality one could enhance the stability of perovskite films and devices under exposure to oxygen and light. In addition, we demonstrate that this degradation occurs even at low oxygen levels under illumination. This points out the importance of avoiding unnecessary light exposure during device fabrication in non-inert atmospheres, in order to reduce the degradation occurring during the production phase and further degradation caused by incorporated oxygen.

Chapter 5 presents the large-scale chemical compositional and electronic inhomogeneities of the $CH_3NH_3PbI_3$ perovskite films observed by XPS and UPS mapping experiments. We demonstrate that these inhomogeneities are not limited to certain fabrication methods, and are therefore likely present in films produced by other research groups. Furthermore, we find that the extent of the variation in photovoltaic performance across the entire sample is related to the extent of the observed inhomogeneities, which explains the wide spread of device performance reported in the literature. It suggests that a plausible way to produce devices with reliable and constant performance is to focus on improving the perovskite fabrication process in order to obtain an active layer with a uniform surface composition and electronic properties. Additionally, our results highlight the importance of studying the surface properties of perovskite layers, instead of simply focusing on bulk properties such as microstructure and absorption, which do not provide any information about surface inhomogeneities.

In Chapter 6, we turn our attention to improving device reproducibility by optimizing the antisolvent treatment procedure using PV, XPS and SEM measurements. We focus on the relationship between antisolvent type and dripping speed, and their effects on PV performance of triple cation perovskite $(Cs_{0.05}(MA_{0.17}FA_{0.83})_{0.95}Pb$ $(I_{0.9}Br_{0.1})_3)$ devices. We find that the solubility of organic precursors in the antisolvent has a tremendous influence on the perovskite formation, microstructure, and consequently, the final device performance. We demonstrate that using certain antisolvents (such as anisole or trifluorotoluene) which possess a low solubility of organic precursors improves device reproducibility by widening the dripping speed 'window', the dripping speeds at which a high-quality perovskite film is formed. In contrast, antisolvents which easily dissolve the organic precursors (such as alcohols) will extract these organics during film formation, significantly damaging the perovskite composition and microstructure. Lastly, we find that antisolvents (such as diethyl ether) which phase separate with the host solvents can lead to incomplete coverage. This suggests that the solubility test is an indicator of which application speed is required to obtain a good device for a new antisolvent. Moreover, our finding shows that these drawbacks can be mitigated by using the appropriate dripping speed, and that for every antisolvent it is possible to produce high-functioning devices by considering this factor.

With more attention attracted to the mixed cation perovskites, and initial performance approaching the records of conventional silicon photovoltaics, future studies should focus more on investigating the stability and reproducibility of these perovskites. For instance, the degradation processes of the triple or quadruple cation perovskite films and devices should be thoroughly investigated under various operating conditions. Furthermore, one can explore the role of different cations on the degradation process, since various cations such as Rb⁺, K⁺ and GA⁺ have been incorporated in the perovskite material system and have been shown to dramatically alter the device stability. Moreover, it would be interesting to study the chemical compositional and electronic inhomogeneities of various mixed cation perovskites, and investigate which perovskite material system yields the most homogenous film and device. Appendices

A Figures

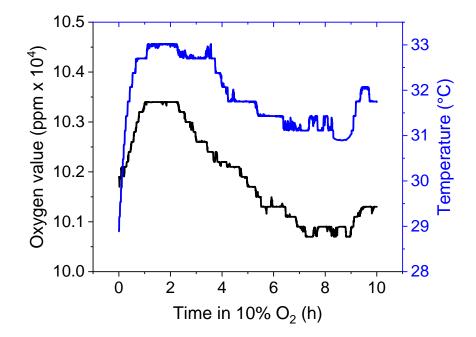


Figure A1: Oxygen level (left axis) and temperature (right axis) of the environmental box monitored throughout the 10% oxygen degradation experiment. Adapted with permission from [58]. Copyright 2017, Wiley-VCH.

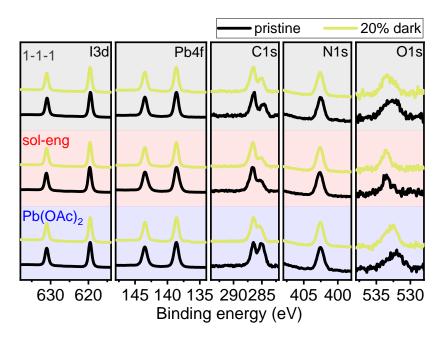


Figure A2: X-ray photoemission spectroscopy measurements on pristine perovskite layers (dark lines) and degraded films for 48 h under 20 % oxygen in dark. Perovskite films were fabricated using the Pb(OAc)₂, sol-eng and 1-1-1 recipes on PEDOT:PSS coated ITO substrates.

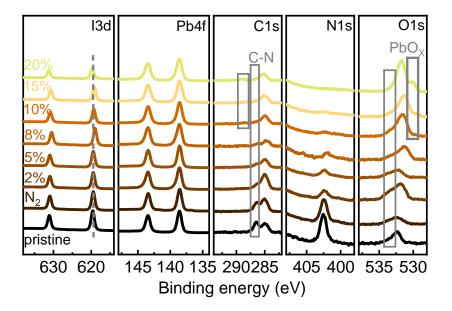


Figure A3: X-ray photoemission spectroscopy measurements on pristine and degraded (10 h under various oxygen levels at 1 sun illumination) perovskite layers fabricated using the Pb(OAc)₂ recipe on ITO substrates.

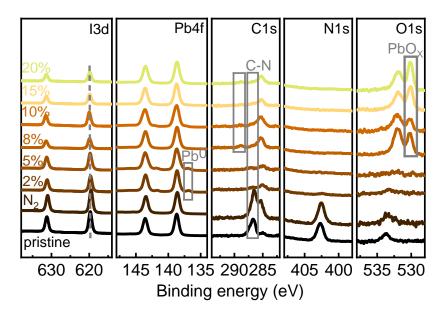


Figure A4: X-ray photoemission spectroscopy measurements on pristine and degraded (10 h under various oxygen levels at 1 sun illumination) perovskite layers fabricated using the sol-eng recipe on ITO substrates.

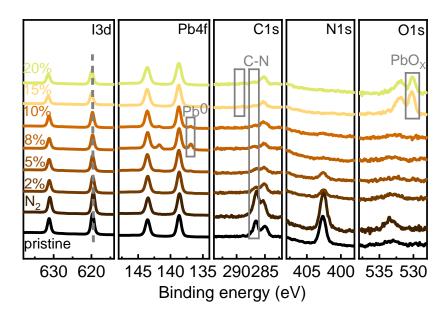


Figure A5: X-ray photoemission spectroscopy measurements on pristine and degraded (10 h under various oxygen levels at 1 sun illumination) perovskite layers fabricated using the 1-1-1 recipe on ITO substrates.

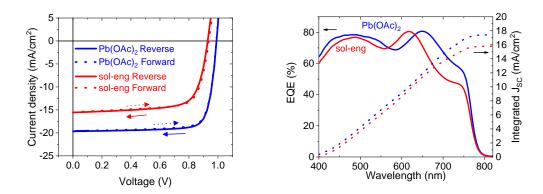


Figure A6: Current-voltage curves of champion devices with the structure ITO/PEDOT:PSS/CH₃NH₃PbI₃/PC₆₁BM/BCP/Ag with perovskite active layers prepared by the Pb(OAc)₂ and sol-eng recipes (left) as well as corresponding EQE spectra and integrated J_{SC} (right). Adapted with permission from [58]. Copyright 2017, Wiley-VCH.

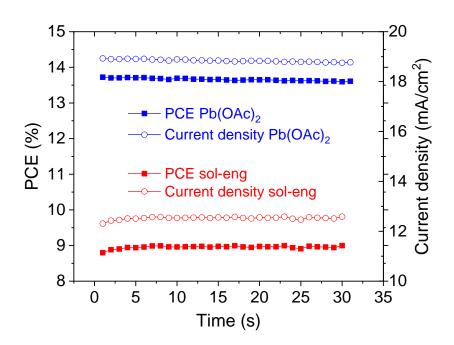


Figure A7: Power conversion efficiency and current density of solar cells fabricated by the Pb(OAc)₂ and sol-eng recipes over 30 s continuous illumination at maximum power point (0.78 V and 0.7 V for Pb(OAc)₂ and sol-eng, respectively.) Adapted with permission from [58]. Copyright 2017, Wiley-VCH.

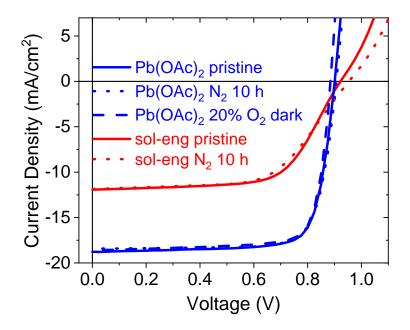


Figure A8: Current-voltage curves of ITO/PEDOT:PSS/CH₃NH₃PbI₃/PC₆₁BM/BCP/Ag solar cells, with perovskite active layers stored in glovebox (pristine), degraded in nitrogen for 10 h and in 20% oxygen for 72 h, for both the Pb(OAc)₂ and sol-eng recipes. Adapted with permission from [58]. Copyright 2017, Wiley-VCH.

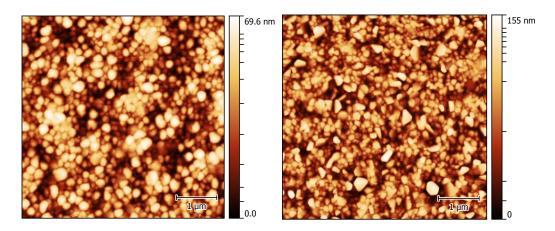


Figure A9: Atomic force microscopy picture of pristine perovskite films prepared by the Pb(OAc)₂ (left) and sol-eng recipes (right). Adapted with permission from [58]. Copyright 2017, Wiley-VCH.

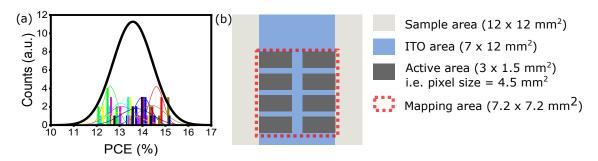


Figure A10: (a) The real histogram of Figure 5.1, which only shows the fitted Gaussian curves to each device (8 pixels, in total 8 devices, i.e., 64 pixels) histogram. (b) Illustration of the substrate area, pixel area and also the mapping area. Detailed numbers are given on the right. Adapted with permission from [140]. Copyright 2018, American Chemical Society.

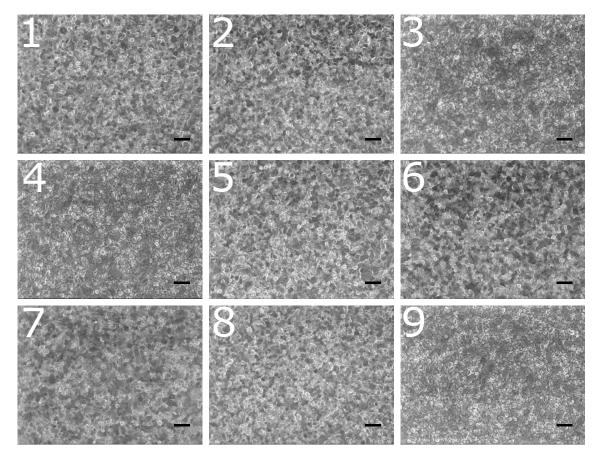


Figure A11: SEM images taken across different sample areas (1-9) on the film prepared by the Pb(OAc)₂ recipe. Scale bar is 1 µm. No strong variations in microstructure were observed. The measurements were carried out by Dr. Paul Fassl. Adapted with permission from [140]. Copyright 2018, American Chemical Society.

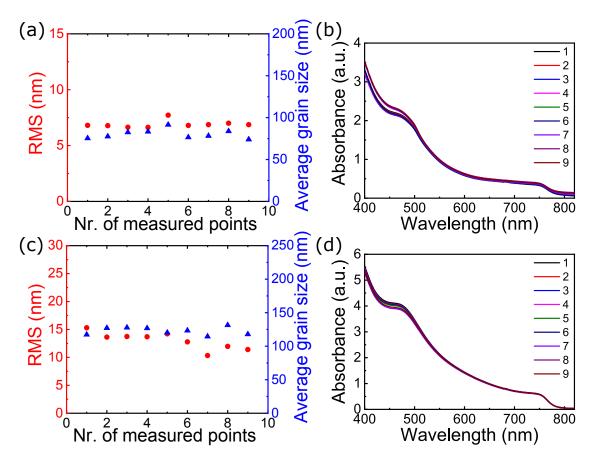


Figure A12: Roughness and average grain size of an inhomogeneous perovskite film produced by (a) sol-eng and (c) 1-1-1 recipes measured at various spots, calculated from AFM measurements. Absorbance spectra of an inhomogeneous perovskite film produced by (b) sol-eng and (d) 1-1-1 recipes measured at various spots. Adapted with permission from [140]. Copyright 2018, American Chemical Society.

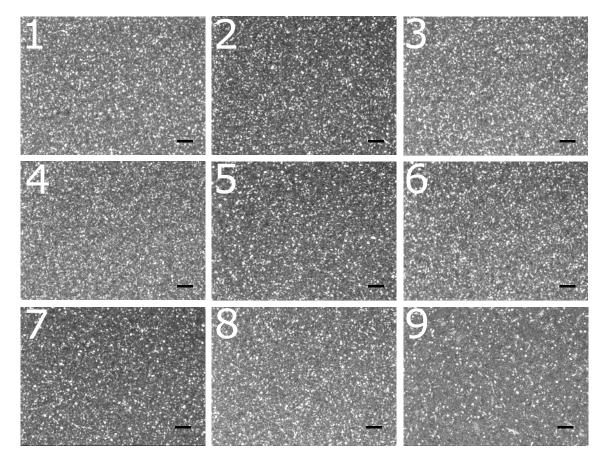


Figure A13: SEM images taken across different sample areas (1-9) on the film prepared by the sol-eng recipe. Scale bar is 1 µm. No strong variations in microstructure were observed. The measurements were carried out by Dr. Paul Fassl. Adapted with permission from [140]. Copyright 2018, American Chemical Society.

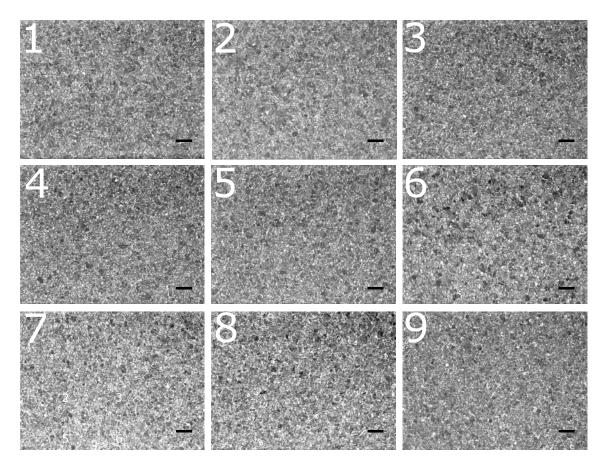


Figure A14: SEM images taken across different sample areas (1-9) on the film prepared by the 1-1-1 recipe. Scale bar is 1 µm. No strong variations in microstructure were observed. The measurements were carried out by Dr. Paul Fassl. Adapted with permission from [140]. Copyright 2018, American Chemical Society.

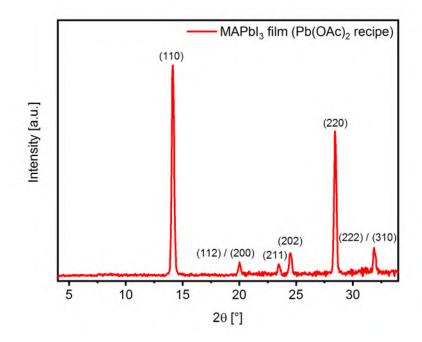


Figure A15: X-ray diffraction spectrum of the perovskite film prepared by the Pb(OAc)₂ recipe. The measurements were conducted by Bruker, Karlsruhe as part of demonstration measurements on a D8 ADVANCE DaVinci diffractometer equipped with a PILATUS3-100K detector. Reproduced with permission from [140]. Copyright 2018, American Chemical Society.

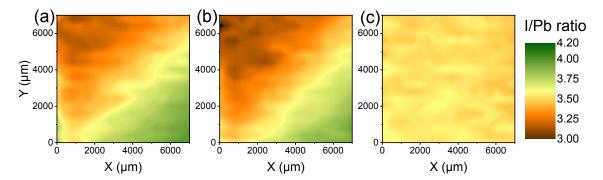


Figure A16: I/Pb maps of perovskite films produced by the Pb(OAc)₂ recipe after storing in (a) a N₂-glovebox, (b) a dry air filled glovebox and (c) vacuum for 50 hours. Reproduced with permission from [140]. Copyright 2018, American Chemical Society.

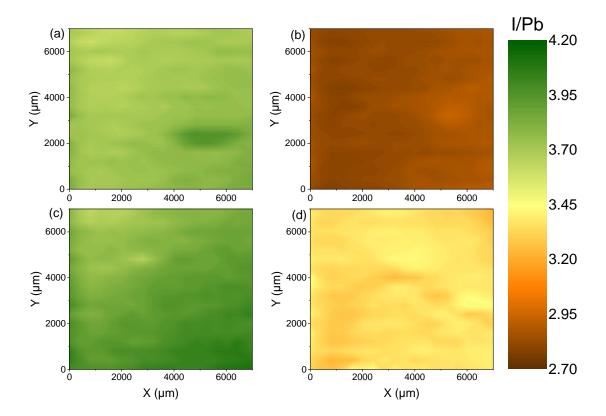


Figure A17: I/Pb maps of perovskite films produced by the sol-eng recipe after storing in (a) a dry air filled glovebox and (b) vacuum for 50 hours. I/Pb maps of perovskite films produced by the 1-1-1 recipe after storing in (a) a dry air filled glovebox and (b) vacuum for 50 hours.

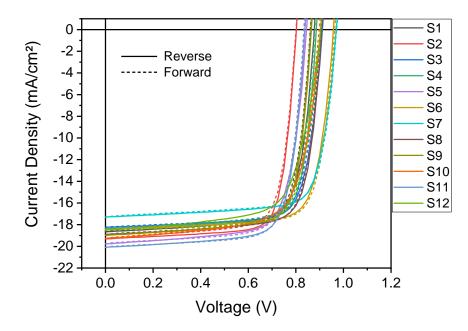


Figure A18: J-V curves of the champion pixels from the twelve solar cells in Figure 5.7. Solid line and dashed line indicates the reverse (open-circuit to short-circuit) and forward (short-circuit to open-circuit) scan directions, respectively. Reproduced with permission from [140]. Copyright 2018, American Chemical Society.

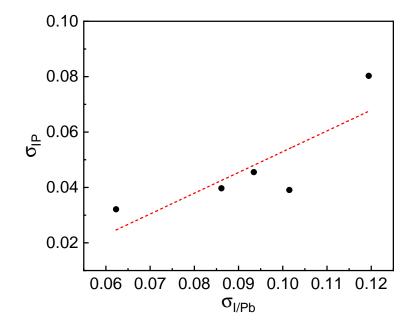


Figure A19: Standard deviation of IP ($\sigma_{\rm IP}$) compared to the standard deviation of I/Pb ($\sigma_{\rm I/Pb}$) of five perovskite films prepared by the Pb(OAc)₂ method, where the XPS and UPS mapping were performed to the same region of the same sample. Only five samples were measured due to the doubling of exposure time to vacuum which changes the surface properties of perovskite films. Reproduced with permission from [140]. Copyright 2018, American Chemical Society.

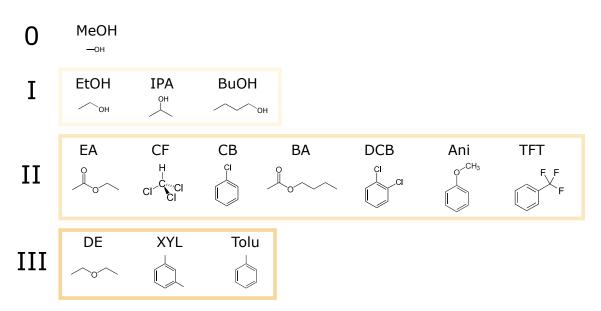


Figure A20: Chemical structures of all solvents used in Chapter 6 as the antisolvent. 0: methanol (MeOH). I: ethanol (EtOH), 2-propanol (IPA) and 1-butanol (BuOH). II: ethyl acetate (EA), chloroform (CF), chlorobenzene (CB), butyl acetate (BA), 1,2-dichlorobenzene (DCB), anisole (Ani) and trifluorotoluene (TFT). III: diethyl ether (DE), m-xylene (XYL) and toluene (Tolu).

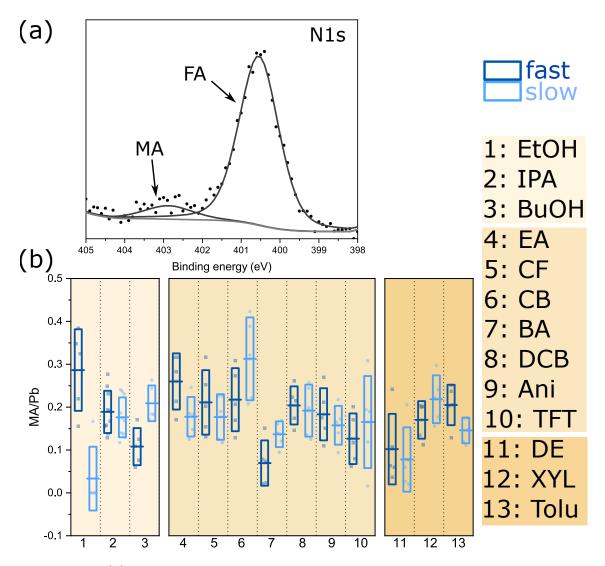


Figure A21: (a) N1s spectrum of the CsMAFA perovskite film coated on ITO/PTAA/PFN-P₂ substrate using TFT as the antisolvent. (b) The MA/Pb ratio of the CsMAFA perovskite films fabricated using 13 antisolvents with fast (dark blue) and slow (light blue) dripping speeds on ITO/PTAA/PFN-P₂ substrates. The used antisolvents are EtOH, IPA, BuOH, EA, CF, CB, BA, DCB, Ani, TFT, DE, XYL and Tolu, from 1 to 13.

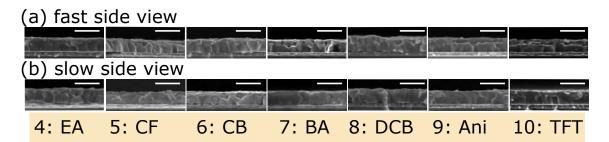


Figure A22: Scanning electron microscopy cross-section images of the CsMAFA perovskite films fabricated by dripping the antisolvent fast (a) and slow (b). The antisolvents used are EA, CF, CB, BA, DCB, Ani and TFT, from left to right. Scale bars represent $1 \,\mu$ m. The measurements were carried out by Dr. Fabian Paulus.

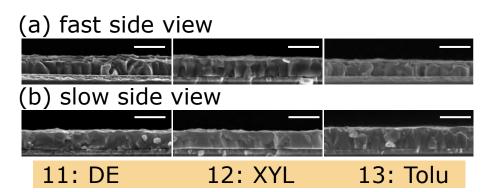


Figure A23: Scanning electron microscopy cross-section images of the CsMAFA perovskite films fabricated by dripping the antisolvent fast (a) and slow (b). The antisolvents used are DE, XYL and Tolu, from left to right. Scale bars represent $1 \,\mu$ m. The measurements were carried out by Dr. Fabian Paulus.

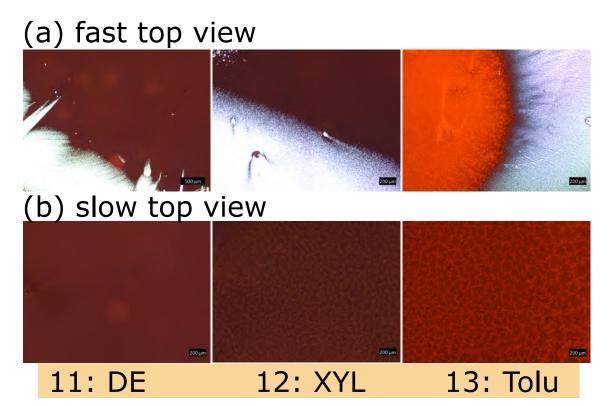


Figure A24: Optical microscopy images (transmission mode) of the CsMAFA perovskite films fabricated by dripping the antisolvent fast (a) and slow (b). The antisolvents used are DE, XYL and Tolu, from left to right. The measurements were carried out by Maximilian Litterst.

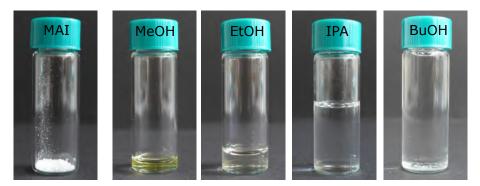


Figure A25: Pictures of 100 mg MAI powder, 100 mg MAI dissolved in 230 μ l MeOH, 100 mg MAI dissolved in 510 μ l EtOH, 100 mg MAI dissolved in 1540 μ l IPA and 100 mg MAI in 4240 μ l BuOH. Note that only the MAI in BuOH solution is not fully dissolved.



Figure A26: Pictures of 100 mg FAI powder, 100 mg FAI dissolved in 170 μ l MeOH, 100 mg FAI dissolved in 350 μ l EtOH, 100 mg FAI dissolved in 820 μ l IPA and 100 mg FAI in 1640 μ l BuOH.

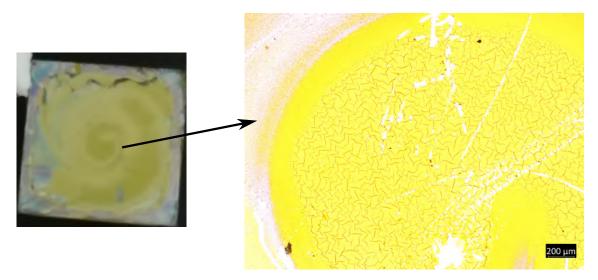
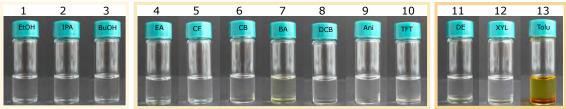


Figure A27: Picture (left) and optical microscopy image (transmission mode) (right) of the 'perovskite' fabricated using MeOH as the antisolvent. The measurements were carried out by Maximilian Litterst.



Antisolvents added to FAI (2 M in DMF/DMSO) solution at a volume ratio of 6:1

Figure A28: Pictures of different solution mixtures, where 13 antisolvents were added to the FAI solution (2M in DMF/DMSO) at a volume ratio of 6:1. The used antisolvents are EtOH, IPA, BuOH, EA, CF, CB, BA, DCB, Ani, TFT, DE, XYL and Tolu, from 1 to 13. 1-10: Clear and transparent solution. 11-13: Solution with two different phases.

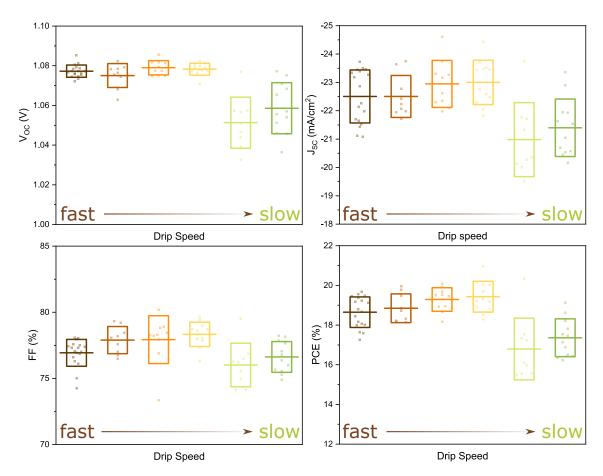


Figure A29: Solar cell characteristics of ITO/PTAA/PFN-P₂/CsMAFA/PC₆₁BM/BCP/Ag solar cells, with the perovskite films fabricated using EA with six different dripping speeds, which gradually change from left, fast, to right, slow. The measurements were carried out by Dr. Alex Taylor.

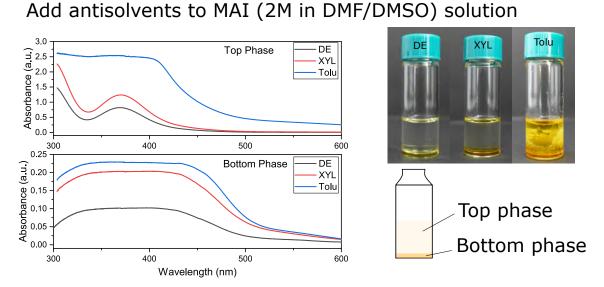


Figure A30: Absorbance spectra of the top and bottom phases of the solution mixtures of the antisolvents (i.e., DE, XYL and Tolu) and MAI solution (2M in DMF/DMSO) at a volume ratio of 6:1. The right panel shows the pictures of these solution mixtures and also an illustrate of the phase separation.

B Tables

Table B1: Densities, molarities and volume factors of the examples of stock solutions.

solution	moles per L	${ m density}\ ({ m g/mL})$	${ m molarity}\ { m (mol/L)}$	volume factor (γ_V)
FAI in DMF/DMSO	2.5	1.190 ± 0.001	2.117	1.181
MAI in DMF/DMSO	2.5	1.158 ± 0.001	2.109	1.185
ITO/PEDC permission	ng (130 dev)T:PSS/CH ₃ NH from [58]. Copy	vices) solar $_3 PbI_3/PC_{61}BN$ right 2017, Wi	cells with M/BCP/Ag. ley-VCH.	(158 devices) the structure Adapted with
Pb(O	$Ac)_2 Pb(0)_2$	$OAc)_2$ s	sol-eng	$\operatorname{sol-eng}$
(reve	rse) (for	ward) (r	everse)	(forward)
$V_{\rm OC}$ (V) 0.91 ±	: 0.04 0.90	± 0.04 0.	$.93 \pm 0.06$	0.92 ± 0.06
$ m J_{SC}~(mA/cm^2)$ 19.3 ±	= 0.9 19.3	± 0.9 14	4.0 ± 1.1	14.0 ± 1.1
FF (%) 78.1 ±	2.5 77.0	± 2.6 64	4.0 ± 10.9	63.3 ± 11.7
PCE (%) 13.68	+ 0.70 13.4	1 ± 0.77 8.	24 ± 0.87	8.05 ± 0.96

Table B3: List of references from 2013 to 2017 from which the values for the spread in PCE was extracted to obtain Figure 5.1 (left). Adapted with permission from [140]. Copyright 2018, American Chemical Society.

	from [140]. Copyright 2018, American Chemical Society.	
Year	Reference	PCE range
Ital	Itelefence	$(\max-\min)(\%)$
2013	W. A. Laban, L. Etgar, Energy Environ. Sci. 2013, 6, 3249.	6 - 8.04
2014	J. Cui, F. Meng, H. Zhang, K. Cao, H. Yuan, Y. Cheng, F. Huang,	7 - 9.84
	M. Wang, ACS Appl. Mater. Interfaces 2014, 6, 22862.	
2014	H. Zhou, Y. Shi, Q. Dong, H. Zhang, Y. Xing, K. Wang, Y. Du,	4.5 - 9.08
	T. Ma, J. Phys. Chem. Lett. 2014, 5, 3241.	
2014	W. Yan, Y. Li, W. Sun, H. Peng, S. Ye, Z. Liu, Z. Bian, C. Huang,	8.75 - 11.8
	RSC Adv. 2014, 4, 33039.	
2014	Y. S. Kwon, J. Lim, HJ. Yun, YH. Kim, T. Park, Energy Env-	6.8 - 9.2
	iron. Sci. 2014, 7, 1454.	
2014	W. H. Nguyen, C. D. Bailie, E. L. Unger, M. D. McGehee, J. Am.	3.9 - 4.7
	Chem. Soc. 2014, 136, 10996.	
2014	J. W. Lee, S. Park, M. J. Ko, H. J. Son, N. G. Park,	6.75 - 8.7
	ChemPhysChem 2014, 15, 2595.	
2014	S. Aharon, S. Gamliel, B. El Cohen, L. Etgar, Phys. Chem. Chem.	5 - 10.85
	Phys. 2014, 16, 10512.	
2014	J. Shi, J. Dong, S. Lv, Y. Xu, L. Zhu, J. Xiao, X. Xu, H. Wu, D.	7.5 - 10.49
	Li, Y. Luo, Q. Meng, Appl. Phys. Lett. 2014, 104, 63901.	
2014	J. Shi, Y. Luo, H. Wei, J. Luo, J. Dong, S. Lv, J. Xiao, Y. Xu, L.	9 - 10.47
	Zhu, X. Xu, H. Wu, D. Li, Q. Meng, ACS Appl. Mater. Interfaces	
	2014, 6, 9711.	
2015	G. Li, K. L. Ching, J. Y. L. Ho, M. Wong, H. S. Kwok, Adv.	11.2 -15.2
2015	Energy Mater. 2015, 5, 1401775.	
2015	K. Wang, C. Liu, P. Du, H. L. Zhang, X. Gong, Small 2015, 11,	6.25 - 11.45
2015	3369.	
2015	X. Bao, Q. Zhu, M. Qiu, A. Yang, Y. Wang, D. Zhu, J. Wang, R.	8.25 - 12.78
0015	Yang, J. Mater. Chem. A 2015, 3, 19294.	11.05 10.50
2015	J. Yin, J. Cao, X. He, S. Yuan, S. Sun, J. Li, N. Zheng, L. Lin, J.	11.25 - 13.76
0015	Mater. Chem. A 2015, 3, 16860.	
2015	C. Y. Chang, Y. C. Chang, W. K. Huang, K. T. Lee, A. C. Cho,	14.6 -16.5
0015	C. C. Hsu, Chem. Mater. 2015, 27, 7119.	0 10.00
2015	Y. C. Shih, L. Y. Wang, H. C. Hsieh, K. F. Lin, J. Mater. Chem.	8 - 12.02
0015	A 2015, 3, 9133.	0.0 10.2
2015	S. Li, K. Jiang, M. Su, X. Cui, J. Huang, Q. Zhang, X. Zhou, L. Vang, V. Sang, I. Mater. Cham. A 2015, 2, 0002	9.8 - 12.3
0015	Yang, Y. Song, J. Mater. Chem. A 2015, 3, 9092.	105 105
2015	D. Forgács, M. Sessolo, H. J. Bolink, J. Mater. Chem. A 2015, 3,	10.5 - 12.5
	14121.	

2015	H. Wei, J. Xiao, Y. Yang, S. Lv, J. Shi, X. Xu, J. Dong, Y. Luo,	10.5 - 13.53
	D. Li, Q. Meng, Carbon N. Y. 2015, 93, 861.	
2015	H. A. Abbas, R. Kottokkaran, B. Ganapathy, M. Samiee, L. Zhang,	10.5 - 13.7
	A. Kitahara, M. Noack, V. L. Dalal, APL Mater. 2015, 3, 016105.	
2015	Y. Li, Y. Zhao, Q. Chen, Y. Yang, Y. Liu, Z. Hong, Z. Liu, Y. T.	14.25 - 17.35
	Hsieh, L. Meng, Y. Li, Y. Yang, J. Am. Chem. Soc. 2015, 137,	
	15540.	
2015	S. Das, B. Yang, G. Gu, P. C. Joshi, I. N. Ivanov, C. M. Rouleau,	8.75 - 13
	T. Aytug, D. B. Geohegan, K. Xiao, ACS Photonics 2015, 2, 680.	
2015	L. Zhu, J. Xiao, J. Shi, J. Wang, S. Lv, Y. Xu, Y. Luo, Y. Xiao,	9.25 - 12.96
	S. Wang, Q. Meng, X. Li, D. Li, Nano Res. 2015, 8, 1116.	
2015	C. Bi, Q. Wang, Y. Shao, Y. Yuan, Z. Xiao, J. Huang, Nat. Com-	13.8 - 17.8
	mun. 2015, 6, 7747.	
2015	L. Zuo, Z. Gu, T. Ye, W. Fu, G. Wu, H. Li, H. Chen, J. Am.	13.05 - 15.67
	Chem. Soc. 2015, 137, 2674.	
2015	J. Zhang, F. Li, K. Yang, C. P. Veeramalai, T. Guo, Appl. Surf.	8 - 9.21
	Sci. 2016, 369, 308.	
2016	J. Huang, X. Yu, J. Xie, D. Xu, Z. Tang, C. Cui, D. Yang, ACS	11.3 - 14.55
	Appl. Mater. Interfaces 2016, 8, 21505.	
2016	A. Dubey, N. Kantack, N. Adhikari, K. M. Reza, S. Venkatesan,	12.4 - 16.83
	M. Kumar, D. Khatiwada, S. Darling, Q. Qiao, J. Mater. Chem.	
2016	A 2016, 4, 10231.	
2016	M. Qin, J. Ma, W. Ke, P. Qin, H. Lei, H. Tao, X. Zheng, L. Xiong,	11.75 - 14.83
	Q. Liu, Z. Chen, J. Lu, G. Yang, G. Fang, ACS Appl. Mater.	
2016	Interfaces 2016, 8, 8460.	15 4 16 09
2016	Y. Guo, W. Sato, K. Shoyama, E. Nakamura, J. Am. Chem. Soc. 2016, 138, 5410.	15.4 - 16.02
2016	T. Singh, T. Miyasaka, Chem. Commun. 2016, 52, 4784.	15.5 - 18.5
2010	C. Y. Chang, W. K. Huang, J. L. Wu, Y. C. Chang, K. T. Lee, C.	13.85 - 15.85
2010	T. Chen, Chem. Mater. 2016, 28, 242.	10.00 - 10.00
2016		15.3 - 18.3
2010	Liu, Y. Zhang, F. Liu, T. P. Russell, H. J. Snaith, R. Zhu, Q.	10.0 10.0
	Gong, Adv. Funct. Mater. 2016, 26, 3508.	
2016	T. P. Gujar, M. Thelakkat, Energy Technol. 2016, 4, 449.	12 - 14.84
2016	X. Zhao, H. Shen, Y. Zhang, X. Li, X. Zhao, M. Tai, J. Li, J. Li,	10.25 - 12.6
	X. Li, H. Lin, ACS Appl. Mater. Interfaces 2016, 8, 7826.	
2016	S. Chatterjee, A. J. Pal, J. Phys. Chem. C 2016, 120, 1428.	3.5 - 8.23
2016	D. Yang, X. Zhou, R. Yang, Z. Yang, W. Yu, X. Wang, C. Li, S.	17.25 - 19.62
	Liu, R. P. H. Chang, Energy Environ. Sci. 2016, 9, 3071.	
2016	H. Zhang, J. Cheng, F. Lin, H. He, J. Mao, K. S. Wong, A. KY.	11.25 - 14.53
	Jen, W. C. H. Choy, ACS Nano 2016, 10, 1503.	
2016	H. Kim, H. Ohkita, H. Benten, S. Ito, Adv. Mater. 2016, 28, 917.	16.5 - 19.4

2017	X. Ren, D. Yang, Z. Yang, J. Feng, X. Zhu, J. Niu, Y. Liu, W.	15.13 - 17.57
	Zhao, S. F. Liu, ACS Appl. Mater. Interfaces 2017, 9, 2421.	
2017	Q. An, P. Fassl, Y. J. Hofstetter, D. Becker-Koch, A. Bausch, P.	11.5 - 18.02
	E. Hopkinson, Y. Vaynzof, Nano Energy 2017, 39, 400.	
2017	J. Ye, X. Zhang, L. Zhu, H. Zheng, G. Liu, H. Wang, T. Hayat,	15 - 19.21
	X. Pan, S. Dai, Sustain. Energy Fuels 2017, 1, 907.	
2017	X. Zhang, Z. Li, Y. Ding, L. Hu, J. Ye, X. Pan, S. Dai, Sustain.	17.5 - 19.19
	Energy Fuels 2017, 1, 1056.	
2017	Y. Wang, S. Xu, J. Deng, L. Gao, R. Soc. Open Sci. 2017, 4,	10 - 15.79
	170980.	
2017	M. A. Mahmud, N. K. Elumalai, M. B. Upama, D. Wang, K. H.	4 - 8.77
	Chan, M. Wright, C. Xu, F. Haque, A. Uddin, Sol. Energy Mater.	
	Sol. Cells 2017, 159, 251.	
2017	S. You, S. Bi, J. Huang, Q. Jia, Y. Yuan, Y. Xia, Z. Xiao, Z. Sun,	16 - 19.7
	J. Liu, S. Sun, Z. Zhao, Chem A Eur. J. 2017, 23, 18140.	
2017	C. Liang, P. Li, Y. Zhang, H. Gu, Q. Cai, X. Liu, J. Wang, H.	13.5 - 16.12
	Wen, G. Shao, J. Power Sources 2017, 372, 235.	
2017	D. Liu, Y. Li, J. Yuan, Q. Hong, G. Shi, D. Yuan, J. Wei, C.	13.75 - 17.22
	Huang, J. Tang, M. Fung, J. Mater. Chem. A 2017, 5, 5701.	
2017	Q. Sun, P. Fassl, D. Becker-Koch, A. Bausch, B. Rivkin, S. Bai,	9.5 - 15.5
	P. E. Hopkinson, H. J. Snaith, Y. Vaynzof, Adv. Energy Mater.	
	2017, 7, 1700977.	
2017	S. Li, P. Zhang, Y. Wang, H. Sarvari, D. Liu, J. Wu, Y. Yang, Z.	14 - 16.08
	Wang, Z. D. Chen, Nano Res. 2017, 10, 1092.	

Table B4: Average device parameters and standard deviations of the twelve solar cells analysed in Figure 5.7. The active layers of these devices have been exposure to vacuum during the UPS mapping measurements (around 1.5 h) and therefore have slightly lower efficiency than the devices, which were fabricated without exposure to vacuum like in Figure 5.1 (right). Adapted with permission from [140]. Copyright 2018, American Chemical Society.

Sample number	V_{OC} (V)	$J_{SC}~(mA/cm^2)$	FF (%)	PCE (%)
1	0.91 ± 0.04	16.7 ± 1.3	73.8 ± 2.6	11.2 ± 0.9
2	0.79 ± 0.01	17.8 ± 1.0	77.3 ± 1.1	10.9 ± 0.7
3	0.91 ± 0.05	17.3 ± 1.3	75.8 ± 2.1	11.9 ± 0.7
4	0.86 ± 0.02	17.6 ± 1.2	75.3 ± 5.3	11.4 ± 1.4
5	0.88 ± 0.02	17.6 ± 1.9	71.7 ± 1.7	11.1 ± 1.3
6	0.92 ± 0.05	17.8 ± 1.3	73.5 ± 4.3	12.0 ± 1.1
7	0.92 ± 0.04	17.3 ± 1.0	73.8 ± 2.0	11.8 ± 1.2
8	0.91 ± 0.02	18.1 ± 0.9	76.0 ± 0.7	12.6 ± 0.7
9	0.86 ± 0.01	18.1 ± 0.9	77.7 ± 0.7	12.0 ± 0.5
10	0.85 ± 0.07	18.1 ± 1.0	59.9 ± 11.4	9.3 ± 2.4
11	0.80 ± 0.15	18.7 ± 2.7	63.6 ± 14.4	10.1 ± 3.8
12	0.78 ± 0.24	15.6 ± 5.8	60.6 ± 12.4	8.7 ± 5.0

Table B5: Physiochemical properties of all solvents used in Chapter 6 as the an-
tisolvent. 0-13: methanol (MeOH), ethanol (EtOH), 2-propanol (IPA),
1-butanol (BuOH), ethyl acetate (EA), chloroform (CF), chlorobenzene
(CB), butyl acetate (BA), 1,2-dichlorobenzene (DCB), anisole (Ani), tri-
fluorotoluene (TFT), diethyl ether (DE), m-xylene (XYL) and toluene
(Tolu). Values for density, boiling point and dipole moment are repro-
duced from PubChem and Wikipedia.

Nr.	Solvent	$\frac{1}{\text{Density (g/mL)}}$	Boiling point (°C)	Dipole moment
0	MeOH	0.79	65	1.69 D
1	EtOH	0.79	78	1.69 D
2	IPA	0.79	83	1.66 D
3	BuOH	0.81	118	1.66 D
4	$\mathbf{E}\mathbf{A}$	0.90	77	1.78 D
5	CF	1.49	61	1.15 D
6	CB	1.11	131	1.69 D
7	BA	0.88	126	1.87 D
8	DCB	1.30	180	$2.50 \ { m D}$
9	Ani	1.00	154	2.30 D
10	TFT	1.19	103	2.86 D
11	DE	0.71	35	1.15 D
12	XYL	0.86	139	0.33-0.37 D
13	Tolu	0.87	111	0.36 D

Table B6: The volume that requires to dissolve 100 mg of MAI.

MAI (mg)	MeOH (μl)	EtOH (μl)	IPA (μl)	BuOH (μl)
100	230	510	1540	${>}4240$

Table B7: The volume that requires to dissolve 100 mg of FAI.

					· · · · · · · · · · · · · · · · · · ·
FA	I (<i>mg</i>)	MeOH (μl)	EtOH (μl)	IPA (μl)	BuOH (μl)
100)	170	350	820	1640

C Calculate the Pbl₂ Amount from SEM Images

To calculate the PbI_2 content on the surface of the perovskite film, we choose the SEM top view images with a magnification of 10,000, as presented in the a and c panels in the Figure C1, C2 and C3. Then, the SEM images are converted to the binary images using the ImageJ software. The binary images of the corresponding SEM images are shown in the b and d panels in the Figure C1, C2 and C3, where the black pixels represent the PbI₂ phases in the SEM. The PbI₂ content is obtained by dividing the number of the black pixels to the total pixels. The SEM measurements were carried out by Dr. Fabian Paulus.

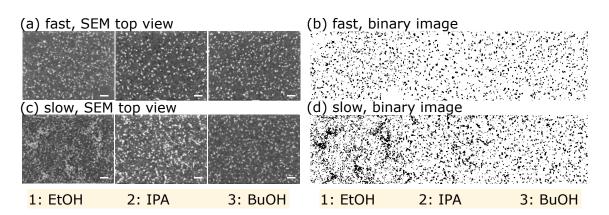


Figure C1: Scanning electron microscopy top view images of the CsMAFA perovskite films fabricated by dripping the antisolvent fast (a) and slow (c). The antisolvents used are EtOH, IPA and BuOH, from left to right. Scale bars represent $1 \,\mu$ m. (b) Binary images of the corresponding SEM images in (a). (d) Binary images of the corresponding SEM images in (c).

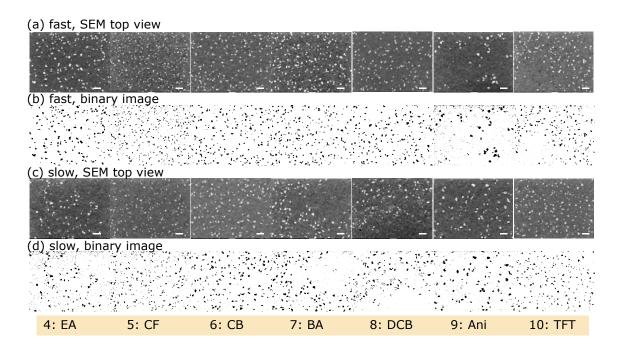


Figure C2: Scanning electron microscopy top view images of the CsMAFA perovskite films fabricated by dripping the antisolvent fast (a) and slow (c). The antisolvents used are EA, CF, CB, BA, DCB, Ani and TFT, from left to right. Scale bars represent $1 \,\mu$ m. (b) Binary images of the corresponding SEM images in (a). (d) Binary images of the corresponding SEM images in (c).

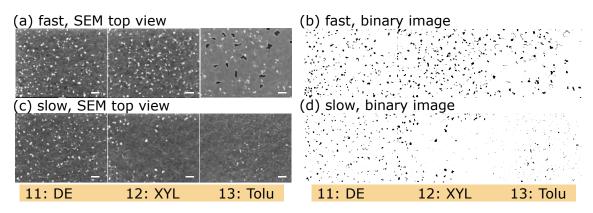


Figure C3: Scanning electron microscopy top view images of the CsMAFA perovskite films fabricated by dripping the antisolvent fast (a) and slow (c). The antisolvents used are DE, XYL and Tolu, from left to right. Scale bars represent $1 \,\mu$ m. (b) Binary images of the corresponding SEM images in (a). (d) Binary images of the corresponding SEM images in (c).

D Abbreviations

Ani	anisole
AFM	atomic force microscopy
BCP	bathocuproine
BA	butyl acetate
BuOH	1-butanol
CB	chlorobenzene
CBM	conduction band minimum
CF	chloroform
CsI	caesium iodide
CsMAFA	triple cation perovskite $Cs_{0.05}(MA_{0.17}FA_{0.83})_{0.95}Pb(I_{0.9}Br_{0.1})_3$ or
	$Cs_{0.05}(MA_{0.17}FA_{0.83})_{0.95}Pb(I_{0.83}Br_{0.17})_3$
DE	diethyl ether
DCB	1,2-dichlorobenzene
DMF	dimethyl formamide
DMSO	dimethyl sulfoxide
$\mathrm{DMF}/\mathrm{DMSO}$	a solution mixture of DMF and DMSO at a volume ratio of 4:1
EA	ethyl acetate
ETL	electron transport layer
EtOH	ethanol
FAI	formamidinium iodide, $CH(NH_2)_2I$
FF	fill factor
GBL	γ -butyrolactone
HBL	hole blocking layer
HPA	hypophosporous acid solution
HTL	hole transport layer
IPA	2-propanol (also called isopropanol)
ITO	indium tin oxide
IP	ionization potential
$ m J_{SC}$	short-circuit current density
Li-TFSI	lithium bis (trifluoromethyl sulfonyl) imide
MAI	methylammonium iodide, $\mathrm{CHNH}_3\mathrm{I}$
$MAPbI_3$	methylammonium lead triiodide, $ m CH_3NH_3PbI_3$
MeOH	$\mathrm{methanol}$
OPVs	organic photovoltaics
TFT	trifluorotoluene
Tolu	toluene
$Pb(OAc)_2 \cdot 3H_2O$	lead acetate trihydrate

PbI_2	lead iodide
$PbBr_2$	lead bromide
$PC_{61}BM$	[6,6]-phenyl C61 butyric acid methyl ester
PCE	power conversion efficiency
PDS	photothermal deflection spectroscopy
PEDOT:PSS	poly (3, 4-ethylenedioxythiopene)-poly (styrenesulfonate)
$PFN-P_2$	[(9,9-bis(3'-(N,N-dimethylamino)propyl)-2,7-fluorene)-alt-2,7-
	(9,9-dioctylfluorene)]
PLQE	photoluminescence Quantum Efficiency
PTAA	poly(triaryl amine), poly[bis(4-phenyl)(2,4,6-trimethyl-
	phenyl)amine]
PV	photovoltaic
RbCsMAFA	quadruple cation perovskite
SEM	scanning Electron Microscopy
Spiro-OMeTAD	2,20,7,70-Tetrakis[N,N-di(4-methoxyphenyl)amino]-9,90-
	spirobifluorene
VBM	valence band maximum
UV	ultra-violet
UV-vis	ultra-violet and visible spectroscopy
UPS	ultra-violet photoemission spectroscopy
V _{OC}	open-circuit voltage
WF	work function
w.r.t.	write with respect to
XPS	X-ray photoemission spectroscopy
XRD	X-ray diffraction
XYL	p-xylene

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List of my Publications

Main Publications:

[1] <u>Q. Sun</u>^{*}, P. Fassl^{*}, D. Becker-Koch, A. Bausch, B. Rivkin, S. Bai, P. E. Hopkinson, H. J. Snaith, Y. Vaynzof, Adv. Energy Mater. 2017, 7, 1700977. ^{*}These authors contributed equally.

[2] Q. Sun, P. Fassl, Y. Vaynzof, ACS Appl. Energy Mater. 2018, 1, 2410.

[3] <u>Q. Sun</u>^{*}, A. D. Taylor^{*}, F. Paulus, Y. Vaynzof. Under preparation. ^{*}These authors contributed equally.

Additional Publications:

[1] B. Rivkin, P. Fassl, <u>Q. Sun</u>, A. D. Taylor, Z. Chen, Y. Vaynzof, ACS Omega 2018, 3, 10042.

[2] Q. An, Q. Sun, A. Weu, D. Becker-Koch, F. Paulus, S. Arndt, F. Stuck, A. S. K. Hashmi, N. Tessler, Y. Vaynzof, Adv. Energy Mater. 2019, 1901257.

[3] Z. Li, A. Ottmann, Q. Sun, A. K. Kast, K. Wang, T. Zhang, H.-P. Meyer, C. Backes, C. Kübel, R. R. Schröder, J. Xiang, Y. Vaynzof, R. Klingeler, J. Mater. Chem. A 2019, 7, 7553.

[4] Z. Li, A. Ottmann, E. Thauer, C. Neef, H. Sai, Q. Sun, K. Cendrowski, H.-P. Meyer, Y. Vaynzof, E. Mijowska, J. Xiang, R. Klingeler, RSC Adv. 2016, 6, 76084.

[5] Z. Li, A. Ottmann, T. Zhang, Q. Sun, H.-P. Meyer, Y. Vaynzof, J. Xiang, R. Klingeler, J. Mater. Chem. A 2017, 5, 3987.

[6] H. Eggers, F. Schackmar, T. Abzieher, Q. Sun, U. Lemmer, Y. Vaynzof, B. S. Richards, G. Hernandez-Sosa, U. Paetzold, Adv. Energy Mater. Submitted.

[7] M. A. Siguan, D. Becker-Koch, A. D. Taylor, Q. Sun, V. Lami, P. G. Oppenheimer, F. Paulus, Y. Vaynzof, ACS Nano. Submitted.

[8] J. F. Butscher, Q. Sun, Y. Wu, F. Stuck, M. Hoffmann, A. Dreuw, A. S. K. Hashmi, N. Tessler, Y. Vaynzof, JPhys Materials. Under preparation.

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